

Resummation for Physics Beyond the Standard Model ¹

University of Oslo

Richard Ruiz

Institute for Particle Physics Phenomenology,
Durham University

24 January 2018

elusives
neutrinos, dark matter & dark energy physics



inVisiblesPlus

¹Based on lots of work: See slides for Refs. (*) = IPPP student.

What this talk is about

I would like to...

- motivate the existence of new physics from a neutrino perspective
- breakdown of fixed order (**FO**) perturbation theory in collider predictions
- factorization and exponentiation in mass gauge theories (QED/QCD)
- impact of QCD resummation on BSM searches at colliders

Central Idea: The property of **universality** of gauge interactions implies what works for SM largely works for BSM

- E.g., NLO+NNLL corrections for W_{SM} and a 4 TeV W_{KK} are identical
- Lots will be covered, so please ask questions!
- This is a BSM seminar, not a formal lecture.

Where we are today

The LHC is operating spectacularly! $\sim 60 \text{ fb}^{-1}$ at 13 TeV ($\sim 5x$ Tevatron)

- Higgs : Not just a hep-th problem but now also a hep-ex problem.
- ν masses , mass hierarchy, particle nature of dark matter, origin of EWSB, etc., require more data and thought

Where we are today

The LHC is operating spectacularly! $\sim 60 \text{ fb}^{-1}$ at 13 TeV ($\sim 5x$ Tevatron)

- Higgs : Not just a hep-th problem but now also a hep-ex problem.
- ν masses , mass hierarchy, particle nature of dark matter, origin of EWSB, etc., require more data and thought

After Run I and early Run II (Fall '17), data is clear:

| Interaction Strength \ Mass Scale | $\Lambda_{\text{BSM}} \lesssim \langle \Phi_{\text{EW}} \rangle$ | $\Lambda_{\text{BSM}} \gg \langle \Phi_{\text{EW}} \rangle$ |
|--|--|---|
| $g_{\text{BSM}} \gtrsim g_{\text{SM}}$ | × | Need more data! |
| $g_{\text{BSM}} \ll g_{\text{SM}}$ | Need more data! | Cannot probe :(|

Picture first suggested by LEP + Belle I + Tevatron is telling:

- No “low hanging fruit”
- hep-ph from 90s-00s designed for “day 1” discoveries, not for extreme regions of BSM parameter space (and hence collider phase space)

Issue: “Day 1” pheno = simple channels with moderately good signal/bkg, e.g., Drell-Yan process like $q\bar{q}' \rightarrow W_R \rightarrow Ne^\pm \rightarrow e^\pm e^\pm + q\bar{q}'$

In “exotic processes”, e.g., VBF and mono- X , contributions from phase space integration ($\int dk^2$) over add'l propagators ($1/k^2$) generates logs:

$$\sigma(pp \rightarrow Y) \sim \log \Lambda_{BSM} / \langle \Phi_{EW} \rangle \text{ and } \sim \log \Lambda_{BSM} / \Lambda_{QCD}$$

and can spoil the validity of **BSM predictions** and/or **collider signatures**.

Issue: “Day 1” pheno = simple channels with moderately good signal/bkg, e.g., Drell-Yan process like $q\bar{q}' \rightarrow W_R \rightarrow Ne^\pm \rightarrow e^\pm e^\pm + q\bar{q}'$

In “exotic processes”, e.g., VBF and mono- X , contributions from phase space integration ($\int dk^2$) over add'l propagators ($1/k^2$) generates logs:

$$\sigma(pp \rightarrow Y) \sim \log \Lambda_{BSM} / \langle \Phi_{EW} \rangle \text{ and } \sim \log \Lambda_{BSM} / \Lambda_{QCD}$$

and can spoil the validity of **BSM predictions** and/or **collider signatures**.

Solution: These are issues long-understood by the **pQCD** community: exploit **soft/collinear factorization**, **resummation**, and **IRC-safety**. From this perspective, BSM collider pheno looks **qualitatively different**.

Results focus on Seesaw partners (N, W_R) but are applicable/ necessary for other high-mass, colorless systems, e.g., $W^\pm h, \tilde{l}\tilde{\nu}_l$.

Issue: “Day 1” pheno = simple channels with moderately good signal/bkg, e.g., Drell-Yan process like $q\bar{q}' \rightarrow W_R \rightarrow Ne^{\pm} \rightarrow e^{\pm}e^{\pm} + q\bar{q}'$

In “exotic processes”, e.g., VBF and mono- X , contributions from phase space integration ($\int dk^2$) over add'l propagators ($1/k^2$) generates logs:

$$\sigma(pp \rightarrow Y) \sim \log \Lambda_{BSM} / \langle \Phi_{EW} \rangle \text{ and } \sim \log \Lambda_{BSM} / \Lambda_{QCD}$$

and can spoil the validity of **BSM predictions** and/or **collider signatures**.

Solution: These are issues long-understood by the **pQCD** community: exploit **soft/collinear factorization**, **resummation**, and **IRC-safety**. From this perspective, BSM collider pheno looks **qualitatively different**.

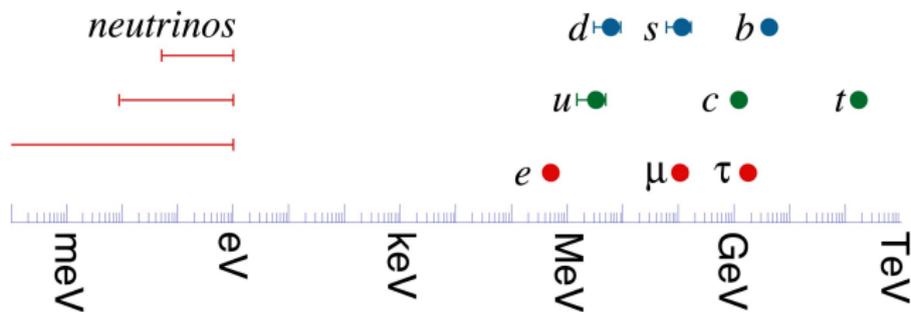
Results focus on Seesaw partners (N , W_R) but are applicable/ necessary for other high-mass, colorless systems, e.g., $W^{\pm}h$, $\tilde{l}\tilde{\nu}_l$.

Message: *QCD is a useful and powerful tool for BSM@Colliders*

Motivation for new physics

Our Motivation

The SM, via the Higgs Mechanism, explains *how* elementary fermions obtain mass, i.e., the $m_f = y_f \langle \Phi \rangle$, **not** the values of m_f .



Spanning many orders of magnitudes, the relationship of fermion masses is still a mystery. Two observations:

- 1 Neutrinos have mass (BSM physics and **2015** 🏆!)
- 2 Neutrinos have unusually small mass (**Seesaw Mechanism?**)

Evidence for New Physics from Neutrinos

To generate ν masses similar to other SM fermions, we need N_R

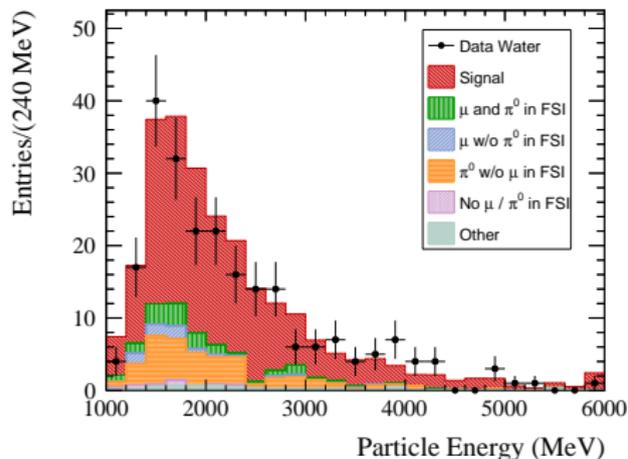
$$\mathcal{L}_{\nu \text{ Yuk.}} = -y_\nu (\overline{\nu}_L \quad \overline{\ell}_L) \begin{pmatrix} \langle \Phi \rangle + h \\ 0 \end{pmatrix} N_R + H.c. \implies -m_D \overline{\nu}_L N_R + H.c.$$

$m_D = y_\nu \langle \Phi \rangle$, and y_ν is the neutrino's Higgs Yukawa coupling.

Since N_R^i do not exist in the SM, massless neutrinos are predicted.

However, we have learned through neutrino oscillations that massless neutrinos is not an accurate description.

[T2K ν_e appearance, 1503.08815v3]



Neutrino masses is evidence of physics beyond the SM.

Collider Connection to Neutrino Mass Models

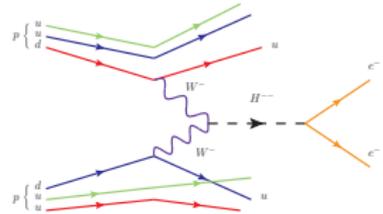
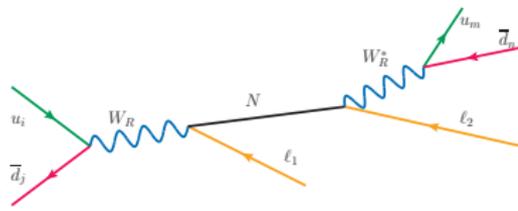
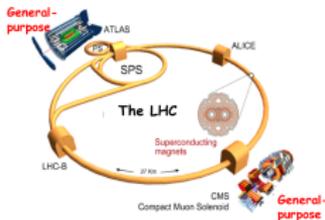
ν mass models predict new partners of all shapes, spins, and color, e.g.,

$$N \text{ (Type I), } T^{0,\pm} \text{ (Type III), } Z_{B-L}, H_R^{\pm,\pm\pm} \text{ (Type I+II)}$$

Through gauge couplings and mixing, production in $ee/ep/pp$ collisions²

$$\text{DY} : q\bar{q} \rightarrow \gamma^*/Z^* \rightarrow T^+T^- \quad \text{and} \quad q\bar{q}' \rightarrow W_R^\pm \rightarrow N\ell^\pm$$

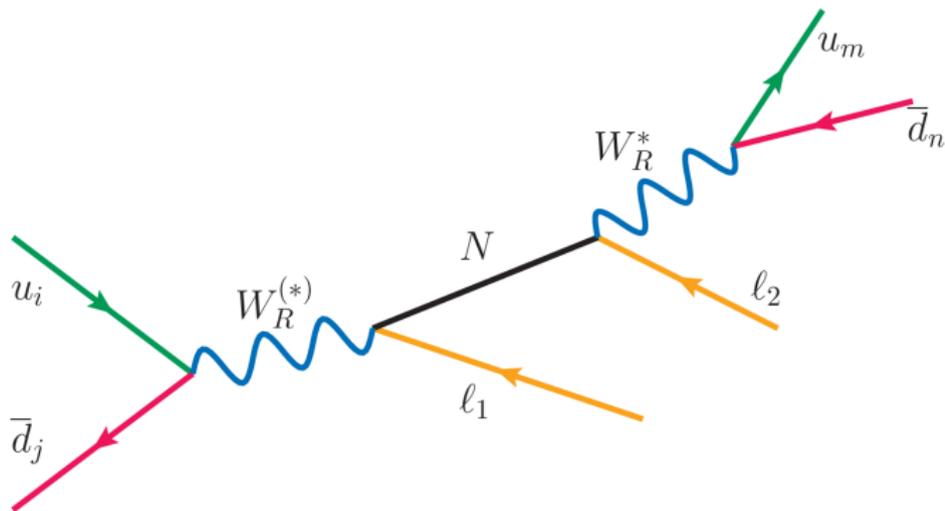
$$\text{VBF} : W^\pm W^\pm \rightarrow H^{\pm\pm} \quad \text{GF} : gg \rightarrow h^*/Z^* \rightarrow N\nu_\ell$$



Identification of Seesaw partners is then inferred by their decays to SM particles and the associated final-state kinematics

²Review on ν mass models at colliders, Y. Cai, T. Li, T. Han, and RR [1711.02180]

Benchmark Scenario: Left-Right Symmetry at Hadron Colliders



Left-Right Symmetric Models (**LRSM**) postulate that the SM's $V - A$ structure originates from the spontaneous breakdown of parity symmetry:

$$SU(3)_c \otimes SU(2)_L \otimes \underbrace{SU(2)_R \otimes U(1)_{B-L}}$$

After scalar Δ_R acquires a vev $v_R \gg v_{SM}$: $\hookrightarrow U(1)_Y$

Higgs field Φ then breaks down the EW group $SU(2)_L \otimes U(1)_Y \rightarrow U(1)_{EM}$

Left-Right Symmetric Models (**LRSM**) postulate that the SM's $V - A$ structure originates from the spontaneous breakdown of parity symmetry:

$$SU(3)_c \otimes SU(2)_L \otimes \underbrace{SU(2)_R \otimes U(1)_{B-L}}$$

After scalar Δ_R acquires a vev $v_R \gg v_{SM}$: $\hookrightarrow U(1)_Y$

Higgs field Φ then breaks down the EW group $SU(2)_L \otimes U(1)_Y \rightarrow U(1)_{EM}$

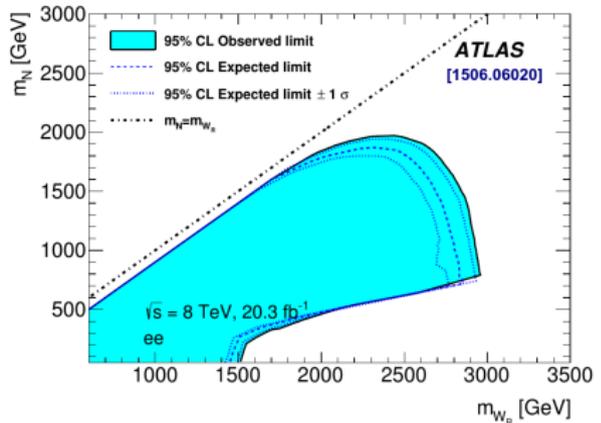
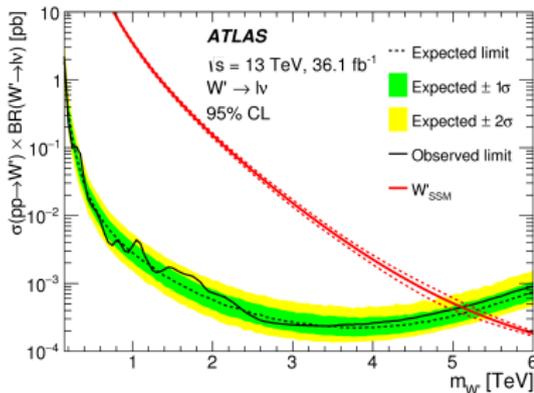
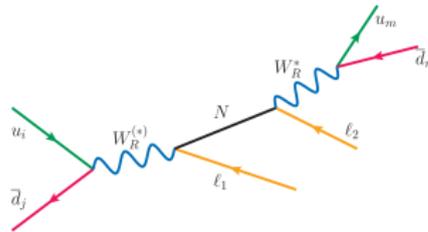
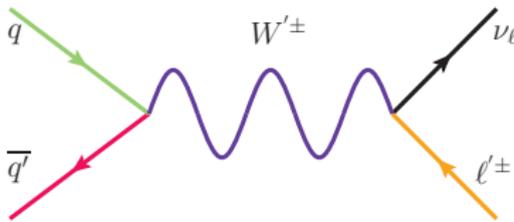
With N_R , all SM fermions can be grouped in $SU(2)_L$ and $SU(2)_R$ doublets. Dirac masses generated in (mostly) usual way with Φ , i.e., $\Delta\mathcal{L} \ni \bar{Q}_L \Phi Q_R$

Neutrinos obtain LH (RH) Majorana masses from triplet scalar Δ_L (Δ_R):

$$m_{\text{light}}^\nu = \underbrace{y_L \langle \Delta_L \rangle}_{\text{Type II}} - \underbrace{\left(y_D y_R^{-1} y_D^T \right) \langle \Phi \rangle^2 \langle \Delta_R \rangle^{-1}}_{\text{Type I a la Type II}} \sim \mathcal{O}(0) + \text{symm.-breaking}$$

Major pheno: heavy N , W'/Z' ($\approx W_R/Z_R$), and $H_i^{\pm\pm}$, H_j^\pm , H_k^0

LHC Tests of Left-Right Symmetry



Moriond: very light, long-lived N : $M_{W_R} \gtrsim 5 \text{ TeV}$ [1706.04786]

Question: Can 13 TeV LHC say anything about $M_{W_R} \gtrsim 5 \text{ TeV}$?

Threshold (or Soft Gluon) Resummation in pQCD³

³**Non-experts:** Roughly speaking, resummation is a procedure for collecting most (or next-to-most or next-to-next-...) divergent radiation terms at each order of perturbation theory to obtain a finite result. Useful since FO results breakdown near poles. ▶

Hadron colliders like the LHC are ultimately counting experiments

$$\underbrace{N_{W_R}}_{\text{No. of } W_R} = \underbrace{\mathcal{L}}_{\text{Luminosity [barns}^{-1}\text{]}} \times \underbrace{\sigma(pp \rightarrow W_R + \text{anything})}_{\text{Production cross section [barns]}}$$

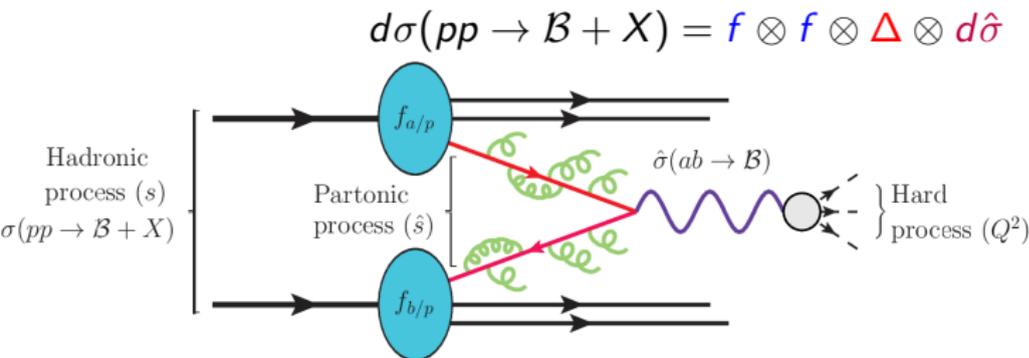
⁴Collins, Soper, Sterman ('85,'88,'89); Collins, Foundations of pQCD (2011)

Hadron colliders like the LHC are ultimately counting experiments

$$\underbrace{N_{W_R}}_{\text{No. of } W_R} = \underbrace{\mathcal{L}}_{\text{Luminosity [barns}^{-1}\text{]}} \times \underbrace{\sigma(pp \rightarrow W_R + \text{anything})}_{\text{Production cross section [barns]}}$$

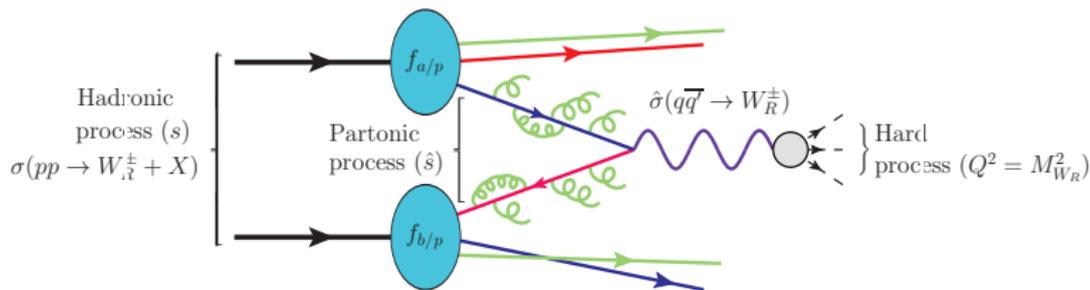
We usually employ the **Collinear Factorization Theorem**⁴
(master equation for colliders) to get hadronic scattering rate:

Hadron-level scattering probabilities are the product (convolution) of parton-dist. (PDFs), -emission (Sudakov), and -scattering probab. ($|\mathcal{M}|^2$)



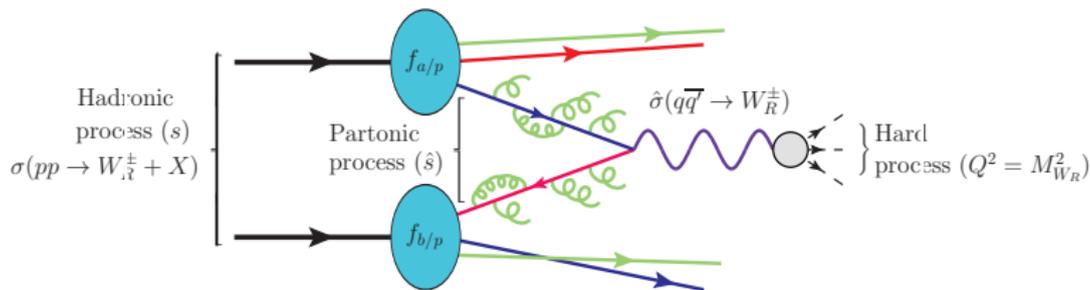
⁴Collins, Soper, Sterman ('85,'88,'89); Collins, Foundations of pQCD (2011)

W_R production is analogous to W_{SM} , except $M_{W_R} \gtrsim 3 - 5 \text{ TeV}$



⁵For total [Altarelli, et al ('79); Sullivan ('02)] and inclusive differential observables [Harris and Owens ('02); Sullivan ('02)], for **any** chiral structure and mass [RR ('15)]

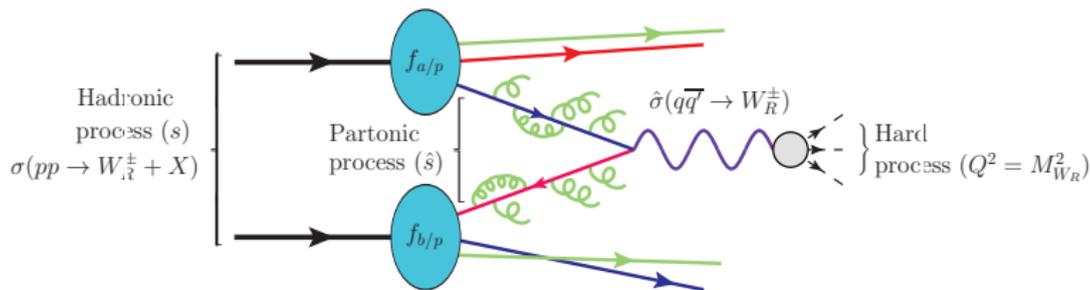
W_R production is analogous to W_{SM} , except $M_{W_R} \gtrsim 3 - 5 \text{ TeV}$



Away from phase space boundaries,

⁵For total [Altarelli, et al ('79); Sullivan ('02)] and inclusive differential observables [Harris and Owens ('02); Sullivan ('02)], for **any** chiral structure and mass [RR ('15)]

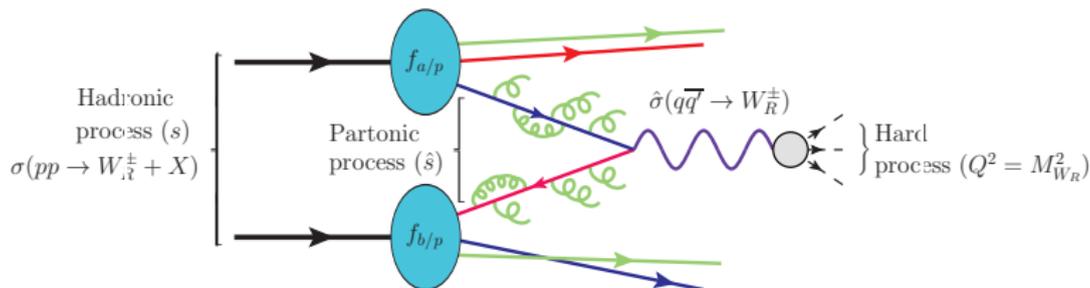
W_R production is analogous to W_{SM} , except $M_{W_R} \gtrsim 3 - 5 \text{ TeV}$



Away from phase space boundaries, QCD corrections for DY are⁵ 20-30%.

⁵For total [Altarelli, et al ('79); Sullivan ('02)] and inclusive differential observables [Harris and Owens ('02); Sullivan ('02)], for **any** chiral structure and mass [RR ('15)]

W_R production is analogous to W_{SM} , except $M_{W_R} \gtrsim 3 - 5 \text{ TeV}$

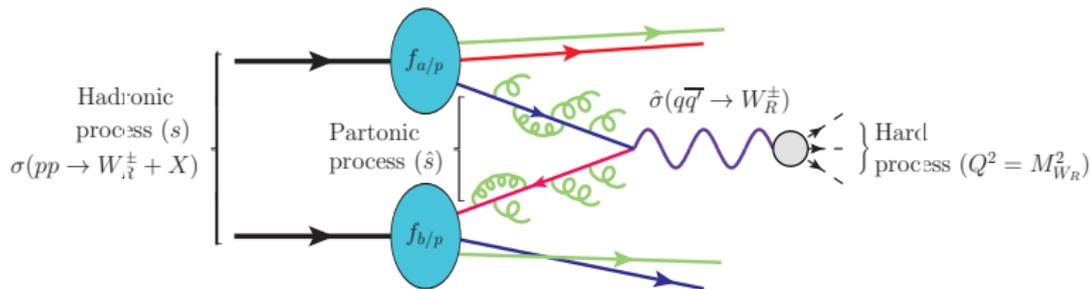


Away from phase space boundaries, QCD corrections for DY are⁵ 20-30%. However, **near** boundaries, e.g., near mass **threshold**, where $E_g \ll E_q$,

$$\begin{aligned} \sigma(q\bar{q}' \rightarrow W_R + g) &\sim \int d^{4-2\epsilon} PS_2 \sim \lambda^{\frac{1-2\epsilon}{2}} \left(1, \frac{Q^2=M_{W_R}^2}{\hat{s}}, \frac{k_g^2=0}{\hat{s}} \right) \\ &= \left(1 - \frac{M_{W_R}^2}{\hat{s}} \right)^{1-2\epsilon} \sim 2\epsilon \log \left(1 - \frac{M_{W_R}^2}{\hat{s}} \right) \end{aligned}$$

⁵For total [Altarelli, et al ('79); Sullivan ('02)] and inclusive differential observables [Harris and Owens ('02); Sullivan ('02)], for **any** chiral structure and mass [RR ('15)]

W_R production is analogous to W_{SM} , except $M_{W_R} \gtrsim 3 - 5 \text{ TeV}$



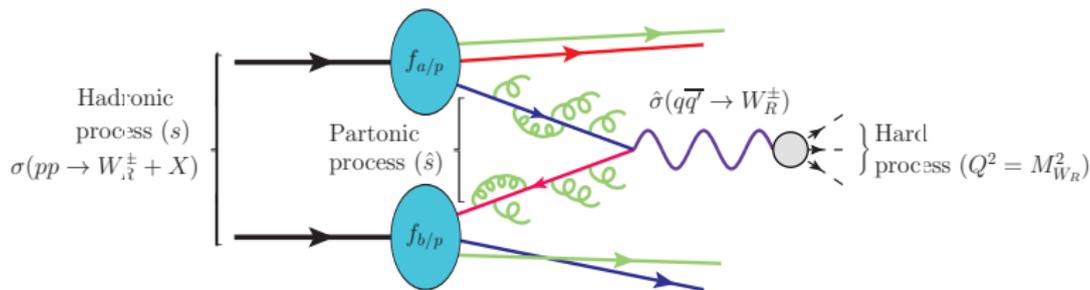
Away from phase space boundaries, QCD corrections for DY are⁵ 20-30%. However, **near** boundaries, e.g., near mass **threshold**, where $E_g \ll E_q$,

$$\begin{aligned} \sigma(q\bar{q}' \rightarrow W_R + g) &\sim \int d^{4-2\epsilon} PS_2 \sim \lambda^{\frac{1-2\epsilon}{2}} \left(1, \frac{Q^2=M_{W_R}^2}{\hat{s}}, \frac{k_g^2=0}{\hat{s}} \right) \\ &= \left(1 - \frac{M_{W_R}^2}{\hat{s}} \right)^{1-2\epsilon} \sim 2\epsilon \log \left(1 - \frac{M_{W_R}^2}{\hat{s}} \right) \end{aligned}$$

As $M_{W_R}^2 \rightarrow s$, logs $> 1/\alpha_s$ since $M_{W_R}^2 \rightarrow \hat{s} < s$ forces g to be soft.

⁵For total [Altarelli, et al ('79); Sullivan ('02)] and inclusive differential observables [Harris and Owens ('02); Sullivan ('02)], for **any** chiral structure and mass [RR ('15)]

W_R production is analogous to W_{SM} , except $M_{W_R} \gtrsim 3 - 5 \text{ TeV}$



Away from phase space boundaries, QCD corrections for DY are⁵ 20-30%. However, **near** boundaries, e.g., near mass **threshold**, where $E_g \ll E_q$,

$$\begin{aligned} \sigma(q\bar{q}' \rightarrow W_R + g) &\sim \int d^{4-2\epsilon} PS_2 \sim \lambda^{\frac{1-2\epsilon}{2}} \left(1, \frac{Q^2 = M_{W_R}^2}{\hat{s}}, \frac{k_g^2 = 0}{\hat{s}} \right) \\ &= \left(1 - \frac{M_{W_R}^2}{\hat{s}} \right)^{1-2\epsilon} \sim 2\epsilon \log \left(1 - \frac{M_{W_R}^2}{\hat{s}} \right) \end{aligned}$$

As $M_{W_R}^2 \rightarrow s$, logs $> 1/\alpha_s$ since $M_{W_R}^2 \rightarrow \hat{s} < s$ forces g to be soft.
 More logs $\implies \mathcal{O}(\alpha_s^{k+1}) > \mathcal{O}(\alpha_s^k) \implies$ expansion in α_s **not** justified.

⁵For total [Altarelli, et al ('79); Sullivan ('02)] and inclusive differential observables [Harris and Owens ('02); Sullivan ('02)], for **any** chiral structure and mass [RR ('15)]

In PT, one (Taylor) expands in powers of coupling constant:

$$\sigma = \sum_k \alpha_s^k \sigma^{(k)} = \sigma^{(0)} + \alpha_s \sigma^{(1)} + \alpha_s^2 \sigma^{(2)} + \mathcal{O}(\alpha_s^3)$$

Stop/truncate at finite/fixed order only if $\mathcal{O}(\alpha_s^{k+1}) < \mathcal{O}(\alpha_s^k)$.

⁶See, e.g., Appell, Sterman, Mackenzie ('88); Forte and Ridolffi ('03)

In PT, one (Taylor) expands in powers of coupling constant:

$$\sigma = \sum_k \alpha_s^k \sigma^{(k)} = \sigma^{(0)} + \alpha_s \sigma^{(1)} + \alpha_s^2 \sigma^{(2)} + \mathcal{O}(\alpha_s^3)$$

Stop/truncate at finite/fixed order only if $\mathcal{O}(\alpha_s^{k+1}) < \mathcal{O}(\alpha_s^k)$.

Soft radiation near **threshold** spoil perturbative convergence since

Higher order terms $>$ lower order terms.

⁶See, e.g., Appell, Sterman, Mackenzie ('88); Forte and Ridolffi ('03)

In PT, one (Taylor) expands in powers of coupling constant:

$$\sigma = \sum_k \alpha_s^k \sigma^{(k)} = \sigma^{(0)} + \alpha_s \sigma^{(1)} + \alpha_s^2 \sigma^{(2)} + \mathcal{O}(\alpha_s^3)$$

Stop/truncate at finite/fixed order only if $\mathcal{O}(\alpha_s^{k+1}) < \mathcal{O}(\alpha_s^k)$.

Soft radiation near **threshold** spoil perturbative convergence since

Higher order terms $>$ lower order terms.

All is not lost! Kinematics of soft, massless radiation are special:⁶
Soft rad. factorizes \rightarrow **all-orders summation** \rightarrow **exponentiation**
 \Rightarrow All-orders (re)summation of $\alpha_s \log(1 - M^2/\hat{s})$

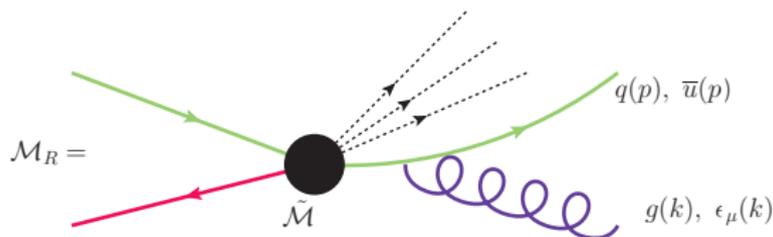
⁶See, e.g., Appell, Sterman, Mackenzie ('88); Forte and Ridolffi ('03) 

Factorization, Exponentiation, and Renormalization Group-Improved Summation⁷

⁷**Non-experts:** Roughly speaking, resummation is a procedure for collecting most (or next-to-most or next-to-next-...) divergent radiation terms at each order of perturbation theory to obtain a finite result. Useful since FO results breakdown near poles. ▶

Soft Factorization in Gauge Theories

Factorization in gauge theories is where a radiation amplitude \mathcal{M}_R in certain kinematic limits can be written as the no-radiation amplitude \mathcal{M}_B and a **universal**, i.e., process-independent, piece:

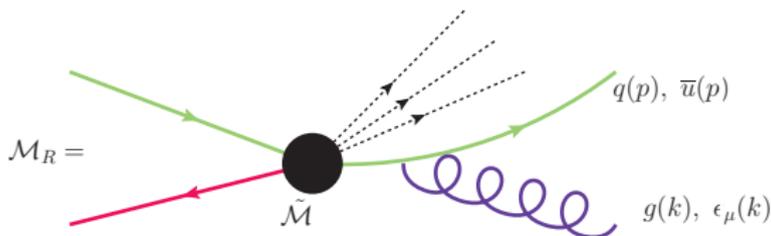


For radiation $q^*(p + k_g) \rightarrow q(p) + g(k_g)$, $E_g \ll E_q$, the amplitude is

$$\mathcal{M}_R \equiv \bar{u}(p)\epsilon_\mu^*(k)(ig_s T^A)\gamma^\mu \frac{\not{p} + \not{k}_g}{(p+k_g)^2} \cdot \tilde{\mathcal{M}} \approx (ig_s T^A)\bar{u}(p) \frac{\epsilon_\mu^* \gamma^\mu \not{p}}{(2p \cdot k_g)} \cdot \tilde{\mathcal{M}}$$

Soft Factorization in Gauge Theories

Factorization in gauge theories is where a radiation amplitude \mathcal{M}_R in certain kinematic limits can be written as the no-radiation amplitude \mathcal{M}_B and a **universal**, i.e., process-independent, piece:



For radiation $q^*(p + k_g) \rightarrow q(p) + g(k_g)$, $E_g \ll E_q$, the amplitude is

$$\mathcal{M}_R \equiv \bar{u}(p) \epsilon_\mu^*(k) (ig_s T^A) \gamma^\mu \frac{\not{p} + \not{k}_g}{(p+k_g)^2} \cdot \tilde{\mathcal{M}} \approx (ig_s T^A) \bar{u}(p) \frac{\epsilon_\mu^* \gamma^\mu \not{p}}{(2p \cdot k_g)} \cdot \tilde{\mathcal{M}}$$

Anti-commute and applying Dirac Eq. gives us

$$\underbrace{\mathcal{M}_R}_{\text{Soft rad. amp.}}|_{\text{Soft}} = (ig_s T^A) \bar{u}(p) \cdot \frac{(p^\mu \epsilon_\mu^*)}{(p \cdot k_g)} \cdot \tilde{\mathcal{M}} = \underbrace{(ig_s T^A) \frac{p^\mu \epsilon_\mu^*}{(p \cdot k_g)}}_{\text{Process independent}} \cdot \underbrace{\tilde{\mathcal{M}}}_{\text{Born amp.}}$$

Factorization of Virtual α_s Corrections to Currents

QCD corrections to colorless currents with massless quarks are special



At one-loop, corrections also **factorize(!)** for *generic* V-A structure:

$$\begin{aligned} \bar{v}(p_d)\gamma^\mu (g_L P_L + g_R P_R) u(p_u) &\rightarrow \bar{v}(p_d)\Gamma^\mu(p_u, p_d)u(p_u), \\ \bar{v}(p_d)\Gamma^\mu(p_u, p_d)u(p_u) &= \bar{v}(p_d)\gamma^\mu (g_L P_L + g_R P_R) u(p_u) \times \mathcal{F} \\ \mathcal{F} &\equiv \frac{\alpha_s(\mu_r^2)}{4\pi} C_F C_\varepsilon(\hat{s}) (-1)^\varepsilon \Gamma(1+\varepsilon) \Gamma(1-\varepsilon) \left(\frac{-2}{\varepsilon^2} - \frac{3}{\varepsilon} - 8 \right) \\ C_\varepsilon(\hat{s}) &= \left(\frac{4\pi\mu_r^2}{\hat{s}} \right)^\varepsilon \frac{\Gamma(1-\varepsilon)}{\Gamma(1-2\varepsilon)}, \quad C_F = 4/3. \end{aligned}$$

$$\sum |\mathcal{M}^{1-Loop}|^2 = \sum |\mathcal{M}^{Born}|^2 (1 + 2\Re[\mathcal{F}]) + \mathcal{O}(\alpha_s)$$

Hold for all phase space at 1-loop, and in soft/coll. limit beyond that.

Sketch of Factorization and Exponentiation

Is it possible to study soft/collinear radiation with perturbative QCD? **Yes!**
Combine our factorized results:

$$\mathcal{M}_{W_R+1 \text{ soft/collinear radiation}} = \left(\underbrace{\text{rad. pole} + \text{loop pole}}_{\text{universal factor}} \right) \times \mathcal{M}_{W_R}^{\text{FO}}$$

The squaring, averaging, and integrating over $(n+1)$ -body phase space

$$d\sigma_{W_R+1 \text{ soft/collinear radiation}} = \underbrace{\int dPS_1(\text{universal piece})}_{\text{finite, } \equiv \mathcal{S}} \Big|_{\text{soft/collinear}} \times \sigma_{W_R}^{\text{FO}}$$

Sketch of Factorization and Exponentiation

Is it possible to study soft/collinear radiation with perturbative QCD? **Yes!**
Combine our factorized results:

$$\mathcal{M}_{W_R+1 \text{ soft/collinear radiation}} = \underbrace{\left(\text{rad. pole} + \text{loop pole} \right)}_{\text{universal factor}} \times \mathcal{M}_{W_R}^{\text{FO}}$$

The squaring, averaging, and integrating over $(n+1)$ -body phase space

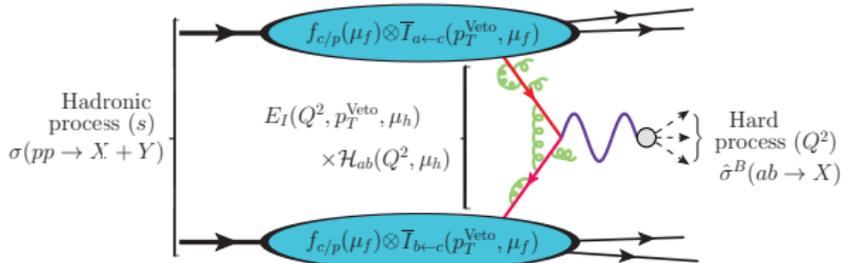
$$d\sigma_{W_R+1 \text{ soft/collinear radiation}} = \underbrace{\int dPS_1(\text{universal piece})}_{\text{finite, } \equiv \mathcal{S}} \Big|_{\text{soft/collinear}} \times \sigma_{W_R}^{\text{FO}}$$

Keeping track of symmetry factors lets us do this for k -emissions:

$$d\sigma_{W_R+k \text{ soft/collinear}} = \frac{1}{k!} [\mathcal{S}]^k \times \sigma_{W_R}^{\text{FO}}$$

Summing over **all** such emissions gives us a closed result:

$$d\sigma_{W_R+\text{any soft/collinear}} = \sum_k \frac{1}{k!} [\mathcal{S}]^k \times \sigma_{\text{DY}}^{\text{FO}} = \exp[\mathcal{S}] \times \sigma_{W_R}^{\text{FO}}$$



A different perspective⁸: In general, scattering rates have the form

$$\frac{d^3\sigma}{d\xi_1 d\xi_2 dz dPS} = \sum_{i,j=q,g,\dots} \underbrace{[f_i(\xi_1, \mu)f_j(\xi_2, \mu) + (1 \leftrightarrow 2)]}_{\text{parton flux, } \hat{s}=\xi_1\xi_2s} \times \underbrace{C(z)}_{\text{soft emissions, } C(z=Q^2/\hat{s})=\delta(1-z)+\mathcal{O}(\alpha_s(\mu))} \times \underbrace{d\hat{\sigma}(ij \rightarrow A)}_{\text{hard process, } Q^2}$$

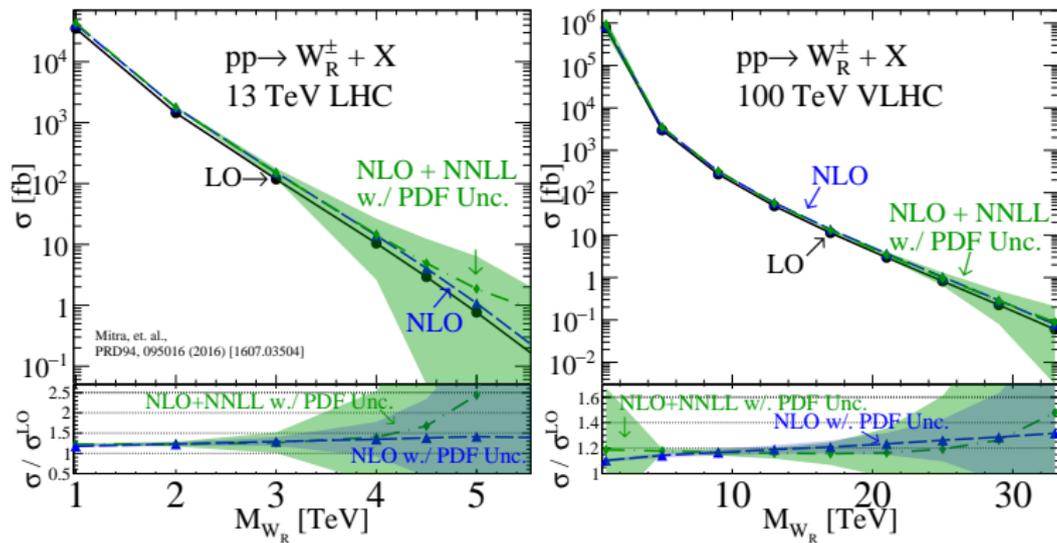
Multi-scale problem: \sqrt{s} , $\sqrt{\hat{s}}$, Q , m_A , but also μ (put in by hand).
Nature works independent of us:

$$\frac{d}{d \log \mu} d\sigma = 0 \implies \frac{d}{d \log \mu} C(z, \mu) = f(z, \mu) C(z, \mu)$$

$$\implies C(z, \mu) = \exp[S(\mu, \mu_0)] C(z, \mu_0) \quad \text{Each piece follows RG evolution}$$

⁸Contopanagos, Laenen, Sterman ('96); Becher, Neubert etc; Stewart, Tackmann etc

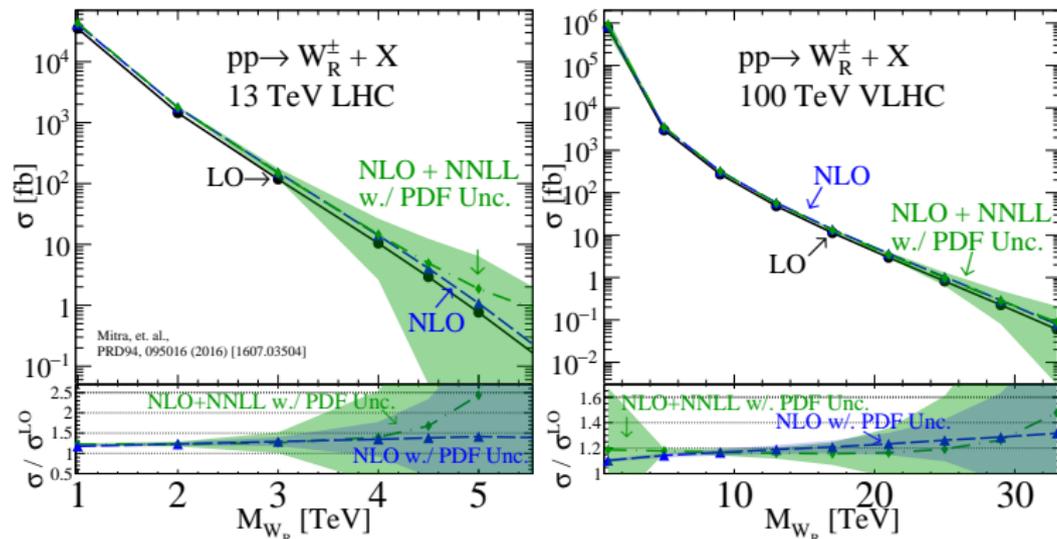
$pp \rightarrow W_R^\pm + X$ at NLO+NNLL(Thresh.)⁹



At 13 TeV, corrections to production rate $> +100\%$ for $M_{W_R} \gtrsim 4.5$ TeV

⁹Mitra, RR, Scott*, Spannowsky [1607.03504]

$pp \rightarrow W_R^\pm + X$ at NLO+NNLL(Thresh.)⁹

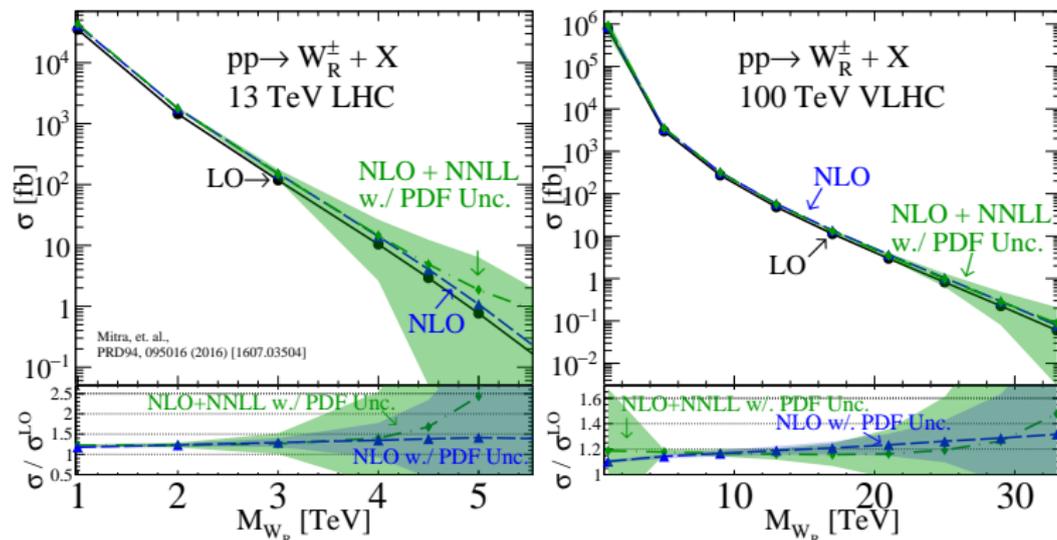


At 13 TeV, corrections to production rate $> +100\%$ for $M_{W_R} \gtrsim 4.5$ TeV

- $\sigma^{LO}(M_{W_R} = 5 \text{ TeV}) \sim 0.7 \text{ fb} \implies \sigma \times (1 \text{ ab}^{-1}) = 700 \text{ events}$

⁹Mitra, RR, Scott*, Spannowsky [1607.03504]

$pp \rightarrow W_R^\pm + X$ at NLO+NNLL(Thresh.)⁹

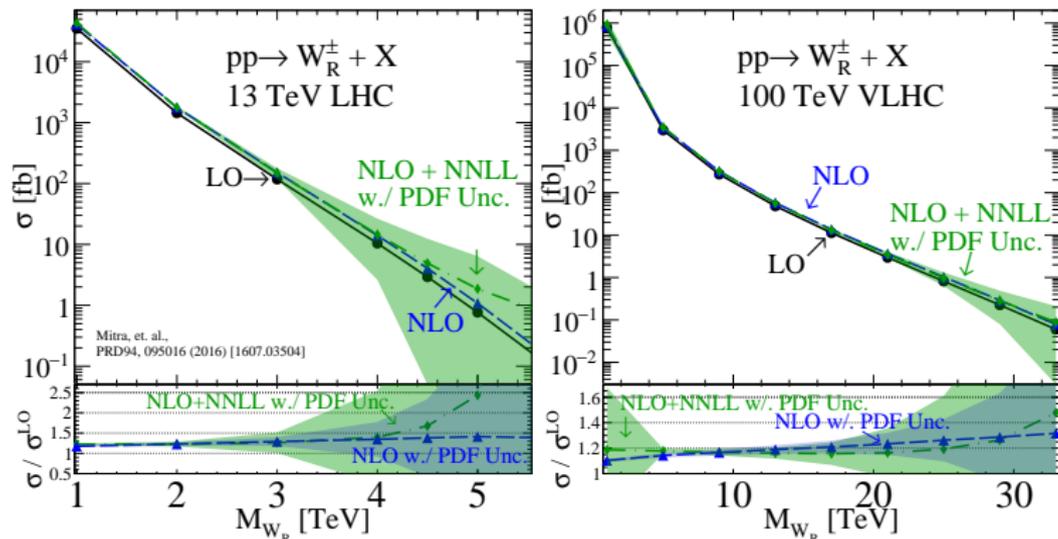


At 13 TeV, corrections to production rate $> +100\%$ for $M_{W_R} \gtrsim 4.5$ TeV

- $\sigma^{LO}(M_{W_R} = 5 \text{ TeV}) \sim 0.7 \text{ fb} \implies \sigma \times (1 \text{ ab}^{-1}) = 700 \text{ events}$
- $\sigma^{NLO+NNLL} \sim 1.7 \text{ fb} \implies \sigma \times (1 \text{ ab}^{-1}) = 1.7\text{k events}$

⁹Mitra, RR, Scott*, Spannowsky [1607.03504]

$pp \rightarrow W_R^\pm + X$ at NLO+NNLL(Thresh.)⁹



At 13 TeV, corrections to production rate $> +100\%$ for $M_{W_R} \gtrsim 4.5$ TeV

- $\sigma^{LO}(M_{W_R} = 5 \text{ TeV}) \sim 0.7 \text{ fb} \implies \sigma \times (1 \text{ ab}^{-1}) = 700 \text{ events}$
- $\sigma^{NLO+NNLL} \sim 1.7 \text{ fb} \implies \sigma \times (1 \text{ ab}^{-1}) = 1.7\text{k events}$

Assuming $\text{BR} \times \epsilon \times \mathcal{A} = 2\% \implies N \approx 34 \text{ events } (\sim 6\sigma \text{ vs } \sim 4\sigma)$

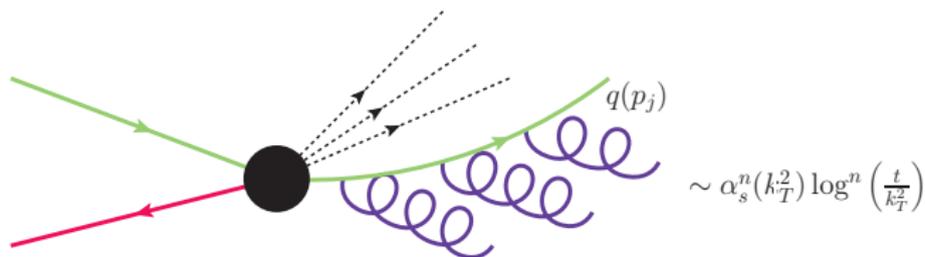
⁹Mitra, RR, Scott*, Spannowsky [1607.03504]

Parton Shower Resummation¹⁰

¹⁰**Non-experts:** Roughly speaking, resummation is a procedure for collecting most (or next-to-most or next-to-next-...) divergent radiation terms at each order of perturbation theory to obtain a finite result. Useful since FO results breakdown near poles. ▶

Parton Showers

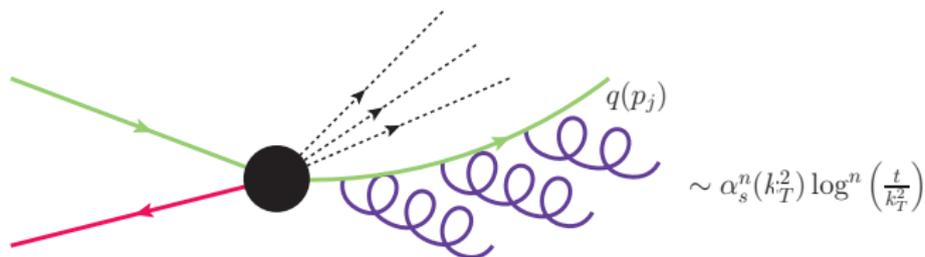
The idea of a parton shower is to capture the emission of soft/collinear (but mostly collinear) emissions in the initial and final state.



Such emissions are not naively $\mathcal{O}(\alpha_s)$ suppressed, i.e., *power suppressed*, since momentum scale of emission is small compared to outgoing particle

Parton Showers

The idea of a parton shower is to capture the emission of soft/collinear (but mostly collinear) emissions in the initial and final state.

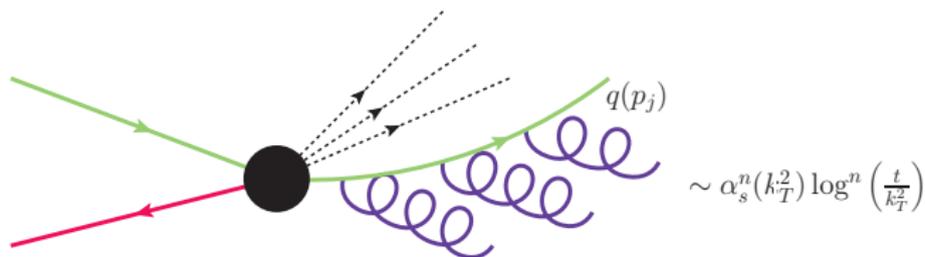


Such emissions are not naively $\mathcal{O}(\alpha_s)$ suppressed, i.e., *power suppressed*, since momentum scale of emission is small compared to outgoing particle

- Logs from additional propagators are large: $dp_T^2/p_T^2 \sim \log p_T^2 \sim 1/\alpha_s$

Parton Showers

The idea of a parton shower is to capture the emission of soft/collinear (but mostly collinear) emissions in the initial and final state.



Such emissions are not naively $\mathcal{O}(\alpha_s)$ suppressed, i.e., *power suppressed*, since momentum scale of emission is small compared to outgoing particle

- Logs from additional propagators are large: $dp_T^2/p_T^2 \sim \log p_T^2 \sim 1/\alpha_s$

Using collinear factorization and unitarity, we can build an evolution factor that accounts for all such radiations, up to leading logarithmic accuracy

For virtuality t of internal line, collinear factorization give us:

$$\sigma_{(n+1)} \sim \sigma_n \times \int dz \frac{dt}{t} \frac{\alpha_s C_i}{2\pi} P_{ji}(z), \quad z = E_g/E_{\text{parent}}$$

The differential splitting probability is then given as

$$d\mathcal{P}_{\text{Split}} \sim \frac{\sigma_{(n+1)}}{\sigma_n} = \frac{dt}{t} \int dz \frac{\alpha_s C_i}{2\pi} P_{ji}(z)$$

For virtuality t of internal line, collinear factorization give us:

$$\sigma_{(n+1)} \sim \sigma_n \times \int dz \frac{dt}{t} \frac{\alpha_s C_i}{2\pi} P_{ji}(z), \quad z = E_g/E_{\text{parent}}$$

The differential splitting probability is then given as

$$d\mathcal{P}_{\text{Split}} \sim \frac{\sigma_{(n+1)}}{\sigma_n} = \frac{dt}{t} \int dz \frac{\alpha_s C_i}{2\pi} P_{ji}(z)$$

By unitarity, the likelihood of a parton at t_0 **not** radiating down to t_1 is

$$\begin{aligned} \mathcal{P}_{\text{No Split}}(t_1, t_0) &= 1 - \mathcal{P}_{\text{Split}} = 1 - \int \frac{dt}{t} \int dz \frac{\alpha_s C_i}{2\pi} P_{ji}(z) \\ &\approx \exp \left[- \int_{t_0}^{t_1} \frac{dt}{t} \int dz \frac{\alpha_s C_i}{2\pi} P_{ji}(z) \right] \end{aligned}$$

For virtuality t of internal line, collinear factorization give us:

$$\sigma_{(n+1)} \sim \sigma_n \times \int dz \frac{dt}{t} \frac{\alpha_s C_i}{2\pi} P_{ji}(z), \quad z = E_g/E_{\text{parent}}$$

The differential splitting probability is then given as

$$d\mathcal{P}_{\text{Split}} \sim \frac{\sigma_{(n+1)}}{\sigma_n} = \frac{dt}{t} \int dz \frac{\alpha_s C_i}{2\pi} P_{ji}(z)$$

By unitarity, the likelihood of a parton at t_0 **not** radiating down to t_1 is

$$\begin{aligned} \mathcal{P}_{\text{No Split}}(t_1, t_0) &= 1 - \mathcal{P}_{\text{Split}} = 1 - \int \frac{dt}{t} \int dz \frac{\alpha_s C_i}{2\pi} P_{ji}(z) \\ &\approx \exp \left[- \int_{t_0}^{t_1} \frac{dt}{t} \int dz \frac{\alpha_s C_i}{2\pi} P_{ji}(z) \right] \end{aligned}$$

Note: As the leading splitting $\alpha \log(t_1/t_0)$ is exponentiated, i.e., sums to all orders in (couplings \times log), this **resummation** is *leading log* accurate

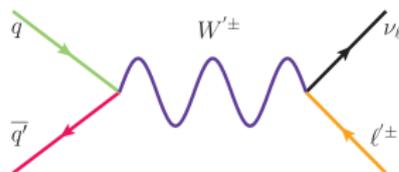
Pheno: NLO+PS/LL(q_T) = **lowest order** at which first QCD radiation is **qualitatively** correct / physically meaningful [CSS ('85)]

BSM @ NLO+PS/LL(q_T)

Major focus of MC community past decade was automation of NLO

LRSM Ex: $(m_N/M_{W_R}) \sim (y_N^\Delta/g_R) \ll 1$

- N is light and "long"-lived
- $pp \rightarrow W_R \rightarrow Ne$ looks like SSM W'



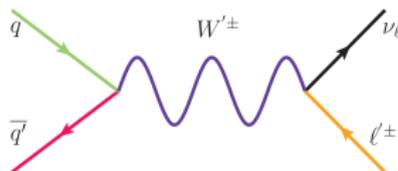
BSM @ NLO+PS/LL(q_T)

Major focus of MC community past decade was automation of NLO

LRSM Ex: $(m_N/M_{W_R}) \sim (y_N^\Delta/g_R) \ll 1$

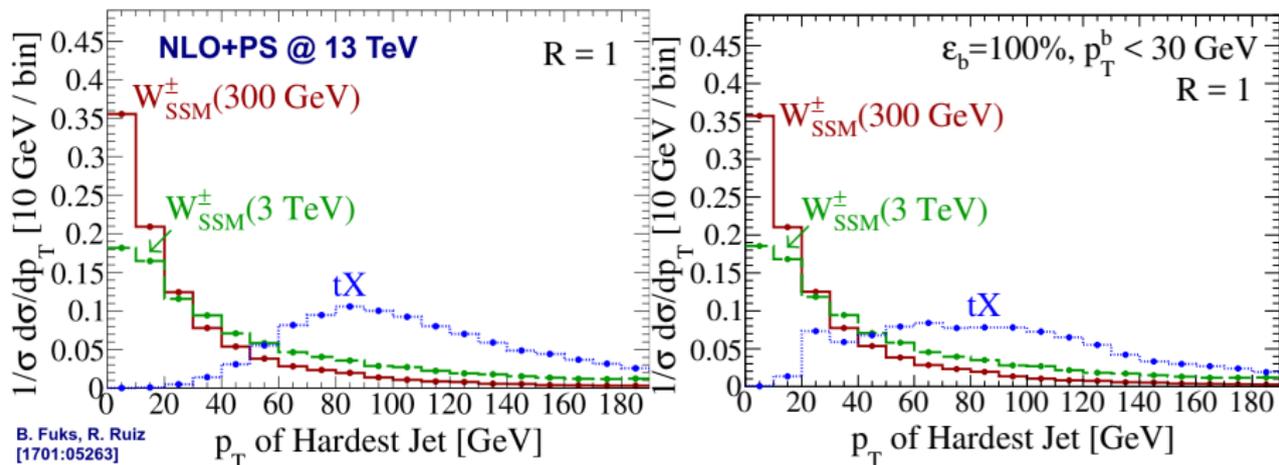
- N is light and "long"-lived

- $pp \rightarrow W_R \rightarrow Ne$ looks like SSM W'



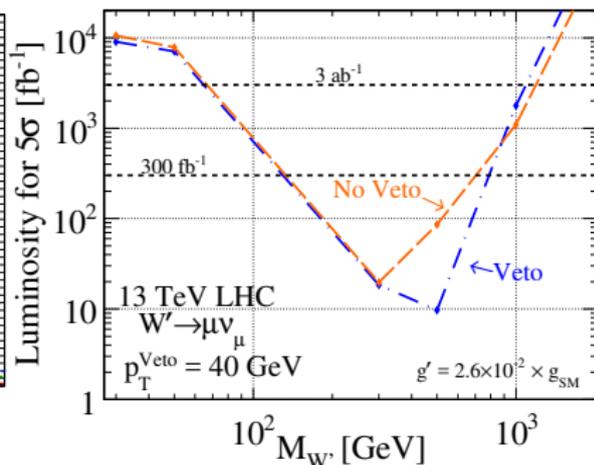
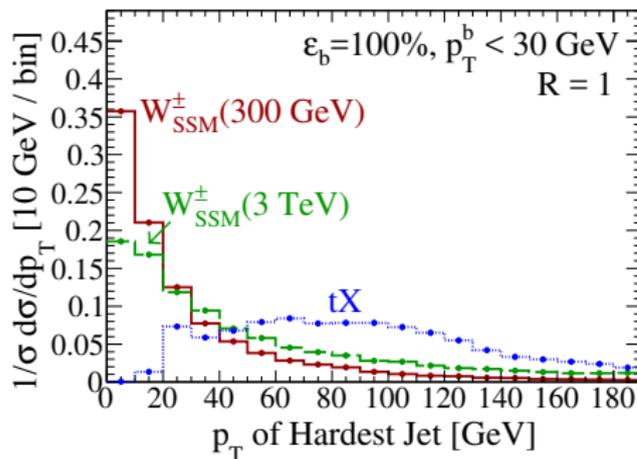
Monte Carlos: modeling jet observables *correctly now possible*

- b -jet vetoes do not remove all QCD radiation ($tX = t\bar{t}, tW, tq$)



B. Fuks, R. Ruiz
[1701.05263]

BSM @ NLO+PS/LL(q_T)+Veto



Monte Carlos: modeling jet observables as *bkg discriminants* **now possible**

- **Why?** QCD/ $t\bar{t}$ have different radiation patterns than color-singlets
- E.g., veto $R = 1$ anti- k_T jets with $p_T^j > 40$ GeV eliminates **top quarks**
- E.g., improve $\tilde{\ell}\tilde{\nu}_\ell$ discovery potential [Tackman, et al, 1603.03052]

NLO+PS in **agreement** w/ NLO+NNLL [Fuks, **RR**, 1701.05263]

- Nontrivial but not total surprise \implies NLO+PS sufficient for discovery

Resummation in Modern Event Generators

NLO+PS and NLO+NNLL(Veto)¹¹ automated in MG5_amc@NLO

- All one needs NLO-accurate FeynRules input model file

¹¹Veto possible for color-singlet processes only. Becher, et. al [[1412.8408](#)]

Resummation in Modern Event Generators

NLO+PS and NLO+NNLL(Veto)¹¹ automated in MG5_amc@NLO

- All one needs NLO-accurate FeynRules input model file

Explosion past two years: [feynrules.irmp.ucl.ac.be/wiki/NLOModels]

| Description | Contact | Reference | FeynRules model files | UFO libraries | Validation material |
|---|------------------|--|-----------------------|--|--|
| Dark matter simplified models (more details) | K. Mawatari | arXiv:1508.00564 , arXiv:1508.05327 , arXiv:1509.05785 | - | DMsimp_UFO.2.zip | - |
| Dark Matter Gauge invariant simplified model (scalar s-channel mediator) (more details) | G. Busoni | arXiv:1612.03475 , arXiv:1710.10764 | - | - | - |
| Effective LR symmetric model (more details) | R. Ruiz | arXiv:1610.08985 | effLRSM.fr | EFFLRSM UFO | - |
| GM (more details) | A. Peterson | arXiv:1512.01243 | - | GM_NLO UFO | - |
| Heavy Neutrino (more details) | R. Ruiz | arXiv:1602.06957 | heavyN.fr | HeavyN_NLO UFO | - |
| Higgs characterisation (more details) | K. Mawatari | arXiv:1311.1829 , arXiv:1407.5089 , arXiv:1504.00611 | - | HC_NLO_X0_UFO.zip | - |
| Inclusive sgluon pair production | B. Fuks | arXiv:1412.5589 | sgluons.fr | sgluons_ufo.tgz | sgluons_validation.pdf ; sgluons_validation_root.tgz |
| Pseudoscalar top-phillic resonance (more details) | D.B. Franzosi | http://arxiv.org/abs/1707.06760 | - | AHTTbar_NLO UFO | - |
| Spin-2 (more details) | C. Degrande | http://arxiv.org/abs/1605.09359 | dm_s_spin2.fr | SMspin2_NLO UFO | - |
| Stop pair -> t bbar + missing energy | B. Fuks | arXiv:1412.5589 | stop_ttmet.fr | stop_ttmet_ufo.tgz | stop_ttmet_validation.pdf ; stop_ttmet_validation_root.tgz |
| SUSY-QCD | B. Fuks | arXiv:1510.00391 | - | susyqcd_ufo.tgz | All figures available from the arxiv |
| Two-Higgs-Doublet Model (more details) | C. Degrande | arXiv:1406.3030 | - | 2HDM_NLO | - |
| Top FCNC Model (more details) | C. Zhang | arXiv:1412.5594 | TopEFFCNC.fr | TopFCNC UFO | - |
| Vector like quarks | B. Fuks | arXiv:1610.04622 | VLQ_v3.fr | UFO in the 5FN5, UFO in the 4FN5, event generation scripts | All figures available from the arxiv |
| W/Z' model (more details) | R. Ruiz, B. Fuks | arXiv:1701.05263 | vPrimeNLO.fr | vPrimeNLO UFO | - |

As of Dec 2017,
updated regularly

Modern general purpose MC packages are very sophisticated

"With great power there must also come - great responsibility" - S. Lee ('62)

¹¹Veto possible for color-singlet processes only. Becher, et al [[1412.8408](#)]

Threshold (or Soft Gluon) Resummation for Gluon Fusion¹²

¹²**Non-experts:** Roughly speaking, resummation is a procedure for collecting most (or next-to-most or next-to-next-...) divergent radiation terms at each order of perturbation theory to obtain a finite result. Useful since FO results breakdown near poles. ▶

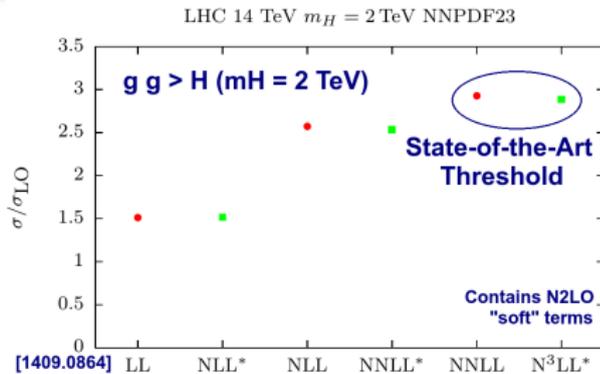
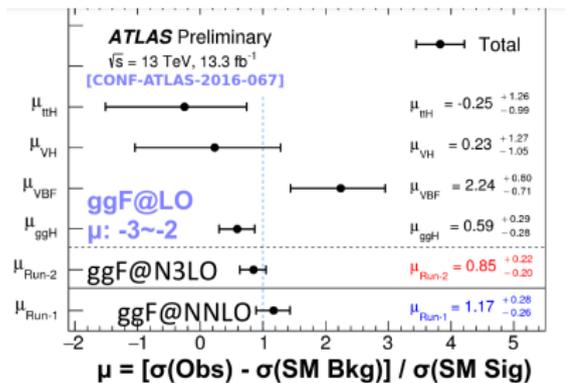
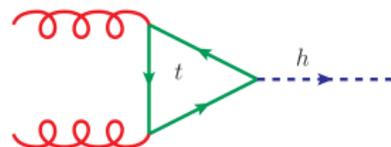
Threshold (or Soft Gluon) Resummation for Gluon Fusion¹²

Myth that QCD is Unimportant for BSM

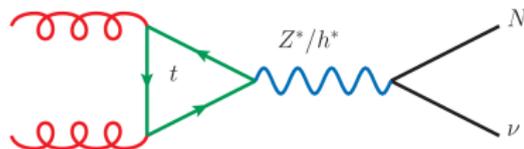
¹²**Non-experts:** Roughly speaking, resummation is a procedure for collecting most (or next-to-most or next-to-next-...) divergent radiation terms at each order of perturbation theory to obtain a finite result. Useful since FO results breakdown near poles. ▶

Threshold Resummation for GF

- QCD corrections to $gg \rightarrow h_{\text{SM}}$ are *large*
- GF@LO is *excluded* by LHC!



- Corrections also *large* for heavy H^0, A^0 .
- Resum. captures leading FO corrections.
- Bonvini, et al, [1409.0864]



- What about heavy N production?

Common Statement: “QCD is unimportant for colorless BSM”

More correct: “Away from phase space boundaries, totally inclusive fixed order QCD corrections are $\sim +20 - 40\%$ for colorless s -channel BSM processes initiated by quarks for non-hierarchical scale choices”

These are the assumptions for the Collinear Factorization Thm

$$\sigma(pp \rightarrow A + \text{anything}) = \sum_{i,j} f_{i/p} \otimes f_{j/p} \otimes \Delta_{ij} \otimes \hat{\sigma}(ij \rightarrow A)$$

Common Statement: “QCD is unimportant for colorless BSM”

More correct: “Away from phase space boundaries, totally inclusive fixed order QCD corrections are $\sim +20 - 40\%$ for colorless s -channel BSM processes initiated by quarks for non-hierarchical scale choices”

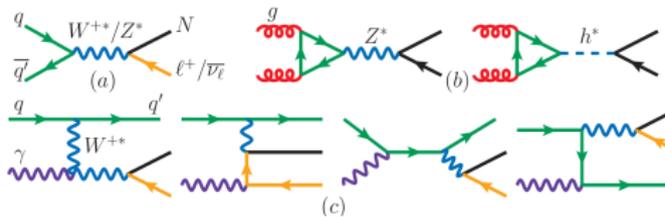
These are the assumptions for the Collinear Factorization Thm

$$\sigma(pp \rightarrow A + \text{anything}) = \sum_{i,j} f_{i/p} \otimes f_{j/p} \otimes \Delta_{ij} \otimes \hat{\sigma}(ij \rightarrow A)$$

Relaxing these assumptions has consequences:

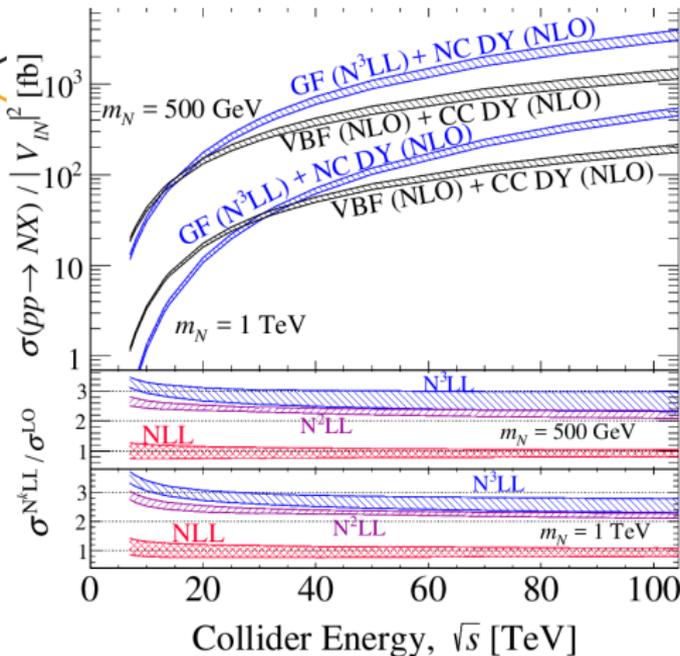
- For $M_{W'/Z'} \sim \sqrt{s}$, $\sigma_{DY}^{NLO+N^2LL}/\sigma^{LO} \sim 2 - 2.5$
- For W'/Z' at any M_V , NLO+PS needed for jet-based/exclusive cuts
- In $gg \rightarrow H^0/A^0$ for any $m_{H/A}$, $\sigma^{N^3LX}/\sigma^{LO} \sim 2 - 3$
- How about $gg \rightarrow h^*/Z^* \rightarrow N\nu$?

Threshold Corrections to Heavy N Production¹³



Heavy N from GF is very interesting:

- ggZ^* contribution identical to pseudoscalar (Equiv Thm)
- Large corrections:
 $\sigma_{GF}^{N^3LL} / \sigma_{GF}^{LO} \sim 2 - 3$
- Largest channel for $\sqrt{s} \gtrsim 20 - 30$ TeV



¹³Willenbrock, Dicus ('85); Dicus, Roy ('91); Hessler, et al [1408.0983]; Degrande, Mattelaer, RR, Turner(*) [1602.06957]; RR, Spannowsky, Waite(*) [1706.02298]

Summary

Over the past decade, a revolution in FO and Resummed calculations!

- New tools + new formalisms \implies new results

Maturity of $N^j\text{LO}/N^k\text{LL}$ formalism allows application to BSM

- Automated tools *and instructions* are publicly available

QCD is a useful/necessary and powerful tool for Seesaws@Colliders

- Threshold resummation $\implies \sigma^{N^k\text{LL}}/\sigma^{LO} = 2 \sim 3$
- p_T /veto resummation \implies new dimension for pheno analyses
- IRC-safety \implies more rigorous and robust collider signatures

Remember: “*The LHC is planned to run over the next 20 years, with several stops scheduled for upgrades and maintenance work.*” [press.cern]

- High-Luminosity LHC and Belle II goals: $1\text{-}3 \text{ ab}^{-1}$ and 50 ab^{-1}
- Premature to claim “nightmare scenario” (SM Higgs + nothing else)

The logo consists of a light blue oval with a wavy line extending from its left and right sides. Inside the oval, the letters 'IP' are written in a large, light blue serif font, and the number '3' is written in a smaller, light blue serif font to the right of 'P'. The text 'Thank you.' is centered over the 'IP' in a bold, black sans-serif font.

Thank you.