



The geophysical signature of impact craters

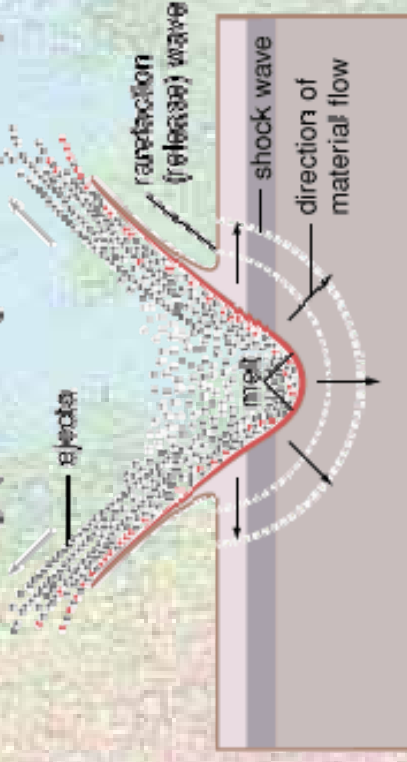
Stephanie Werner

Observation, geophysical techniques and implications

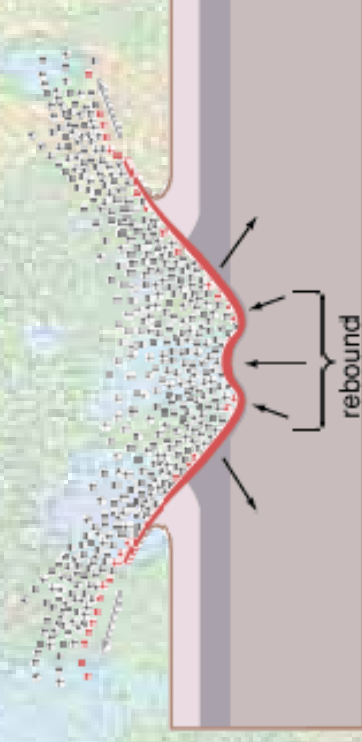
- Gravity Anomalies
- Magnetic Anomalies
- Seismic profiling
- Electrical /Electromagnetic
- Magnetotelluric
- Ground Penetrating Radar
- Radiometric
- ...

Crater Formation Process

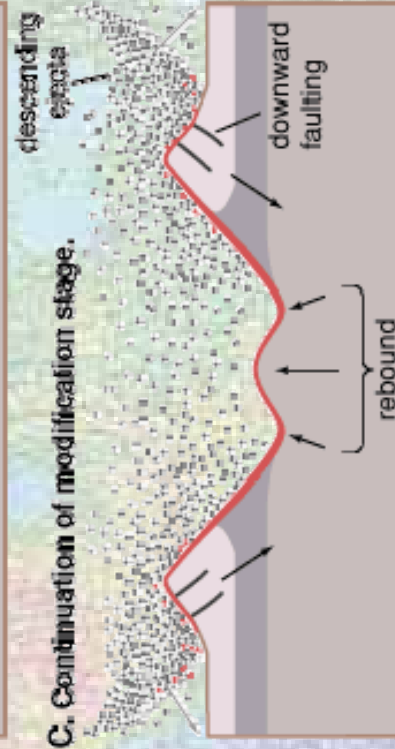
A. Excavation stage (the sole stage for a simple crater).



B. End of excavation stage; start of modification stage.



C. Continuation of modification stage.

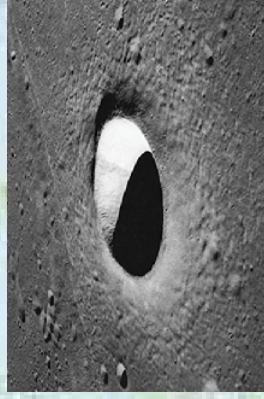


D. Final structure.

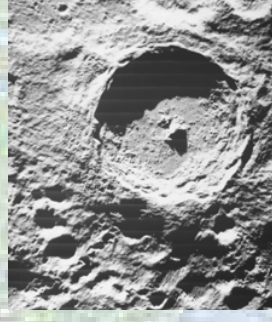


Observations: Physical

Shape:
circular features



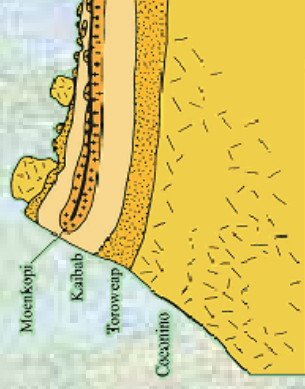
Moltke
(2.7 mi)



Tycho
(53 mi)

Inverted Stratigraphy:

first recognized by Barringer
(only for well preserved craters)



Meteor Crater



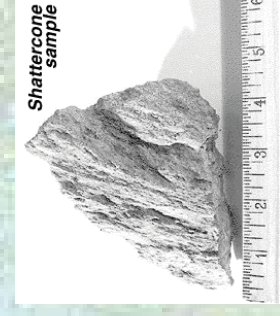
Slide from B. Pierazzo

Material displacement:

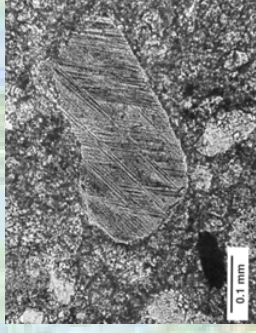
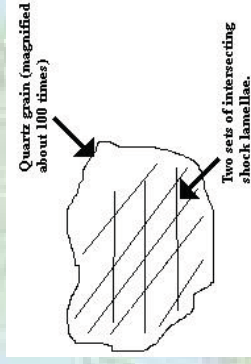
Solid material broken up and ejected
outside the crater: breccia, tektites

Observations: Shock Evidence

Shatter cones:
conical fractures with typical markings produced by shock waves



Shocked Material:
shocked quartz
high pressure minerals



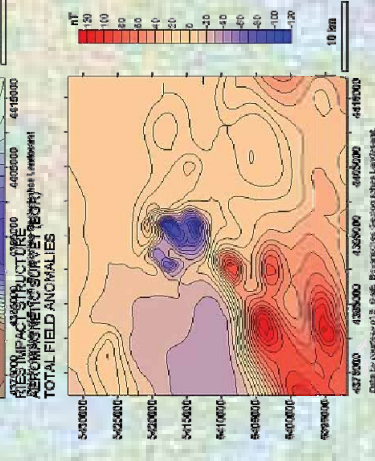
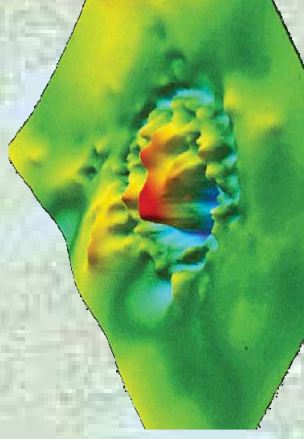
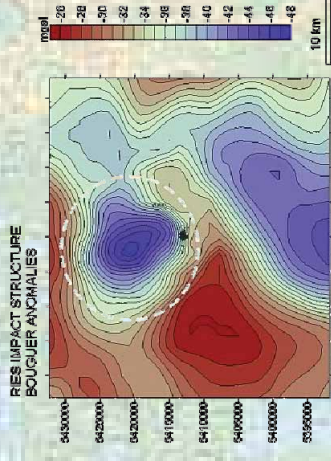
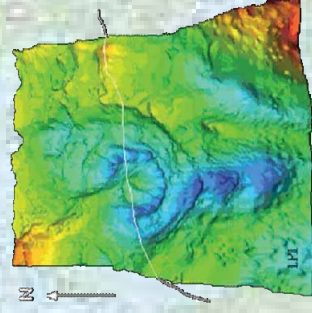
Melt Rocks:
may result from shock and friction



Observations: Geophysical data

Gravity anomaly:

based on density variations of materials

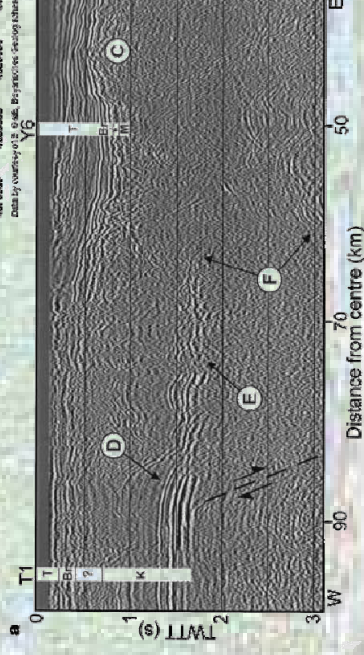


Magnetic:

based on *general* variation of magnetic properties of materials

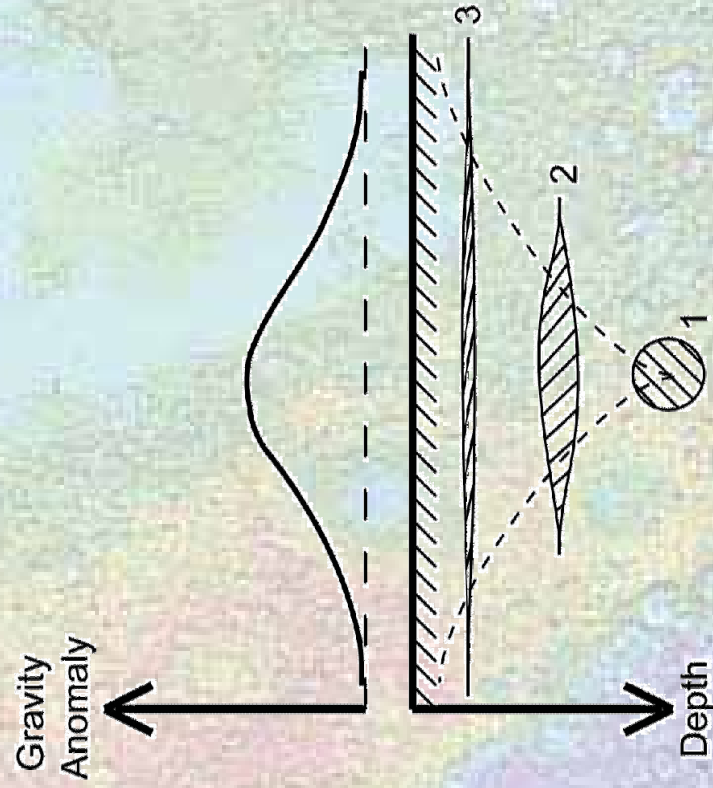
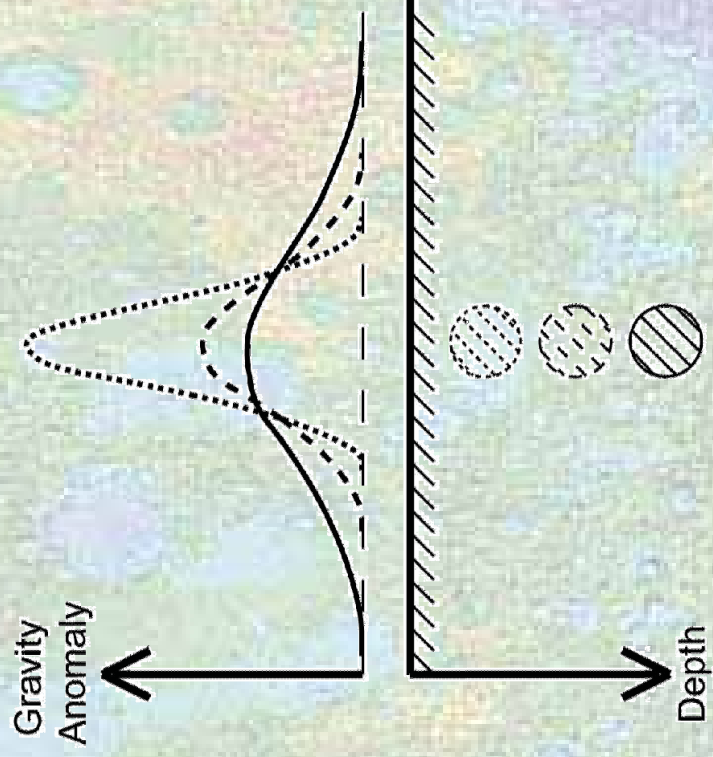
Seismic:

sound waves reflection and refraction from subsurface layers with different characteristics



Slide from B. Pierazzo

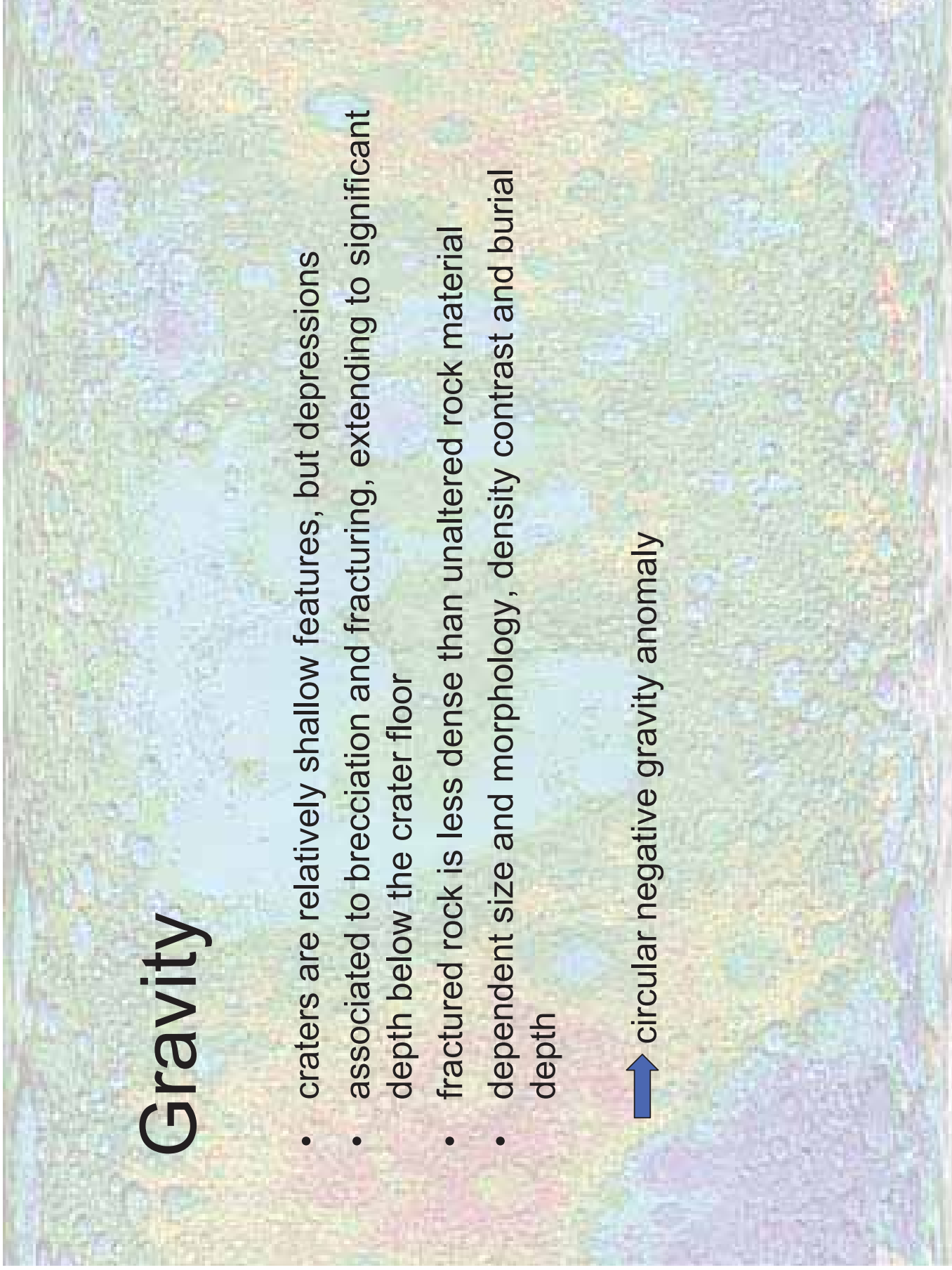
Gravity Anomalies



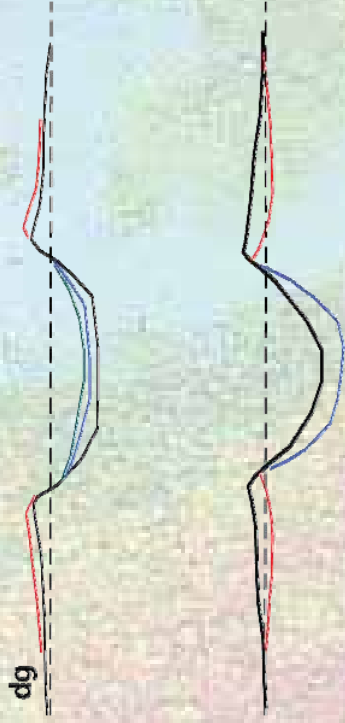
Gravity

- craters are relatively shallow features, but depressions
- associated to brecciation and fracturing, extending to significant depth below the crater floor
- fractured rock is less dense than unaltered rock material
- dependent size and morphology, density contrast and burial depth

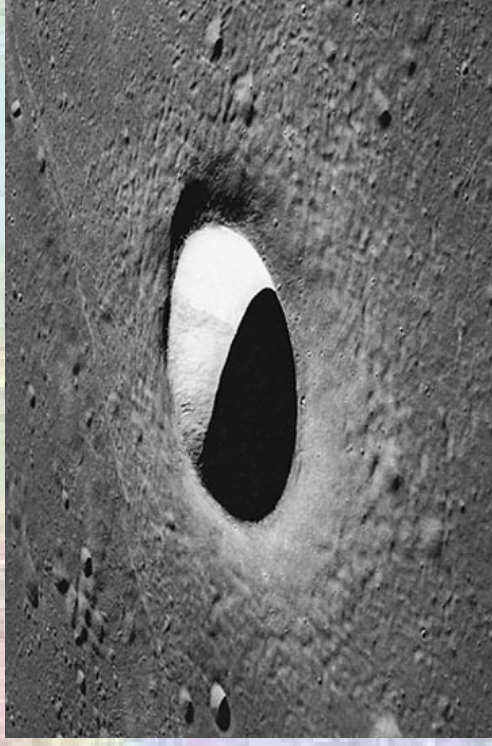
↑ circular negative gravity anomaly



Gravity anomaly: Simple crater

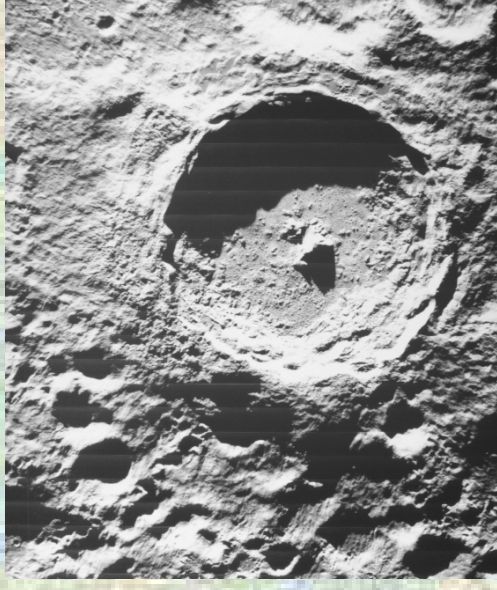
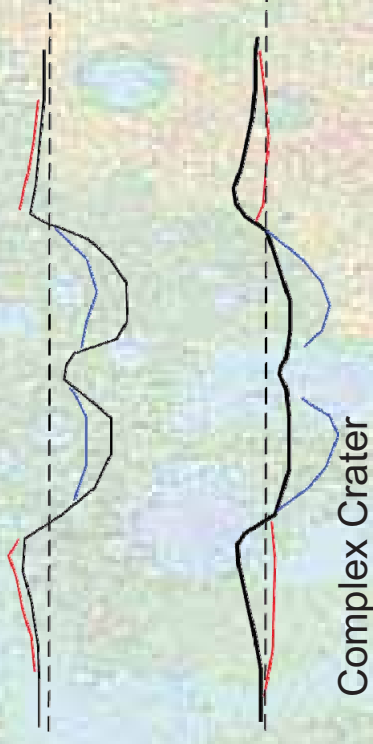


- Negative; unless filled or eroded



Gravity anomaly: Complex Crater

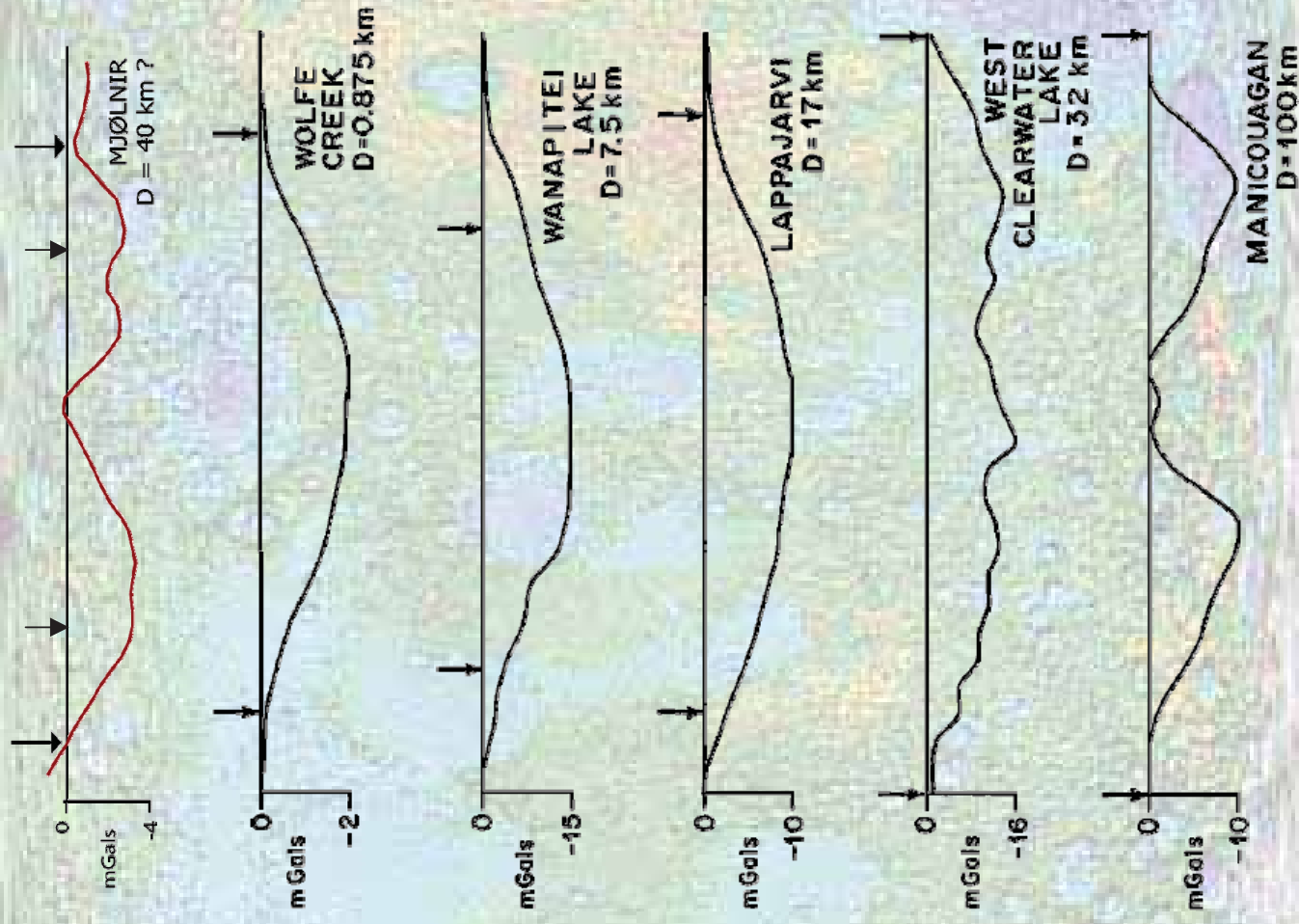
- **fracturing and brecciation of target rock and minerals**
- **formation of high-pressure polymorphs**
- **mineralogic diaplectic change**
- **uplift of the rim wall**
- **formation of allochthonous breccia and melt sheets**



Density Contrast

	kg/m ³	km	Ma
Gosses Bluff, Australia	150	22	142.5 ± 0.5
Mjølnir, Norway	150	40 (25)	142±6
Brent, Canada	170 - 340	3.8	450± 30
Nicholson Lake, Canada	70 – 140	12.5	~400
Clearwater West, Canada	170	32	290± 20
Holleford, Canada	240	2.35	550 ± 100
Manicouagan, Canada	130	100	212 ± 1
Söderfjärden, Finland	160	6	~550
Average (Crystalline)	177.5		

Gravity Signatures



Statistics on gravity anomalies

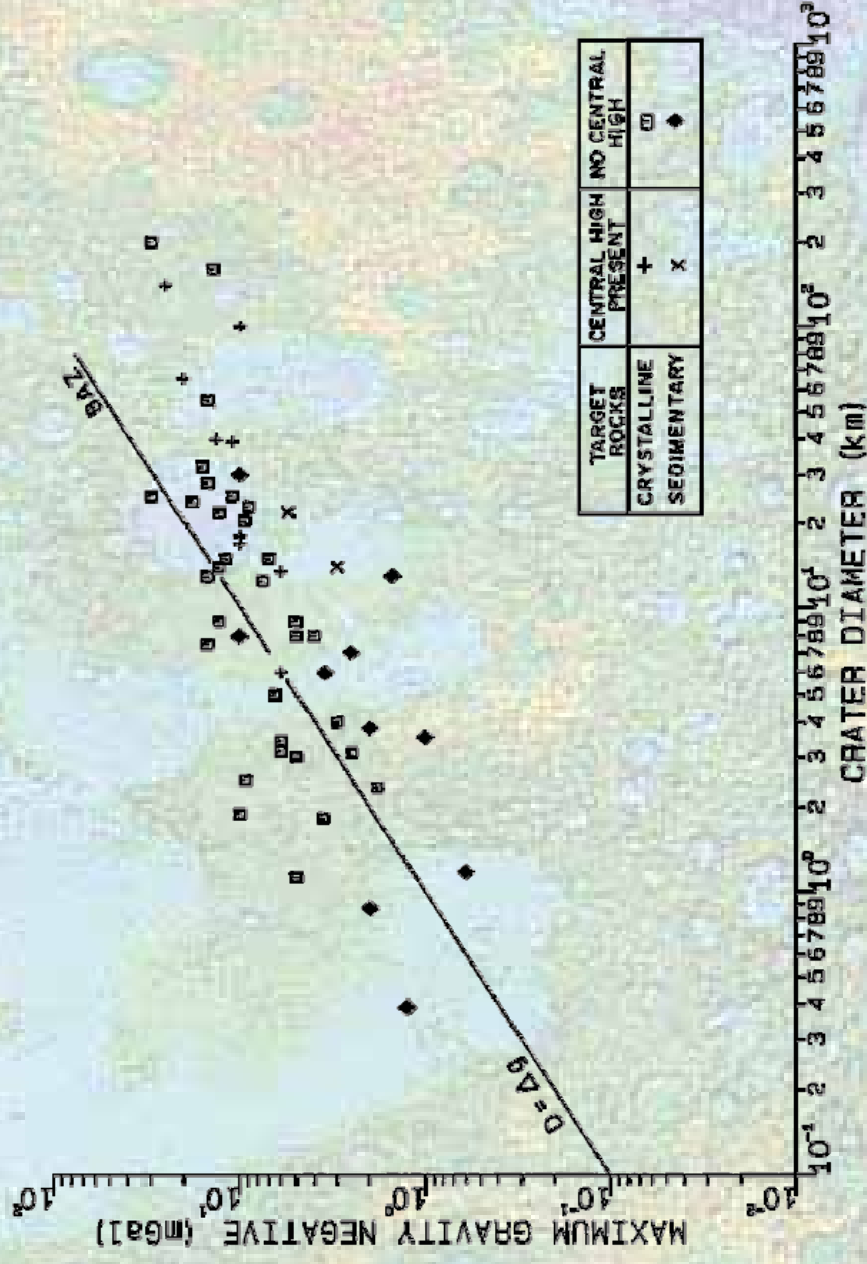
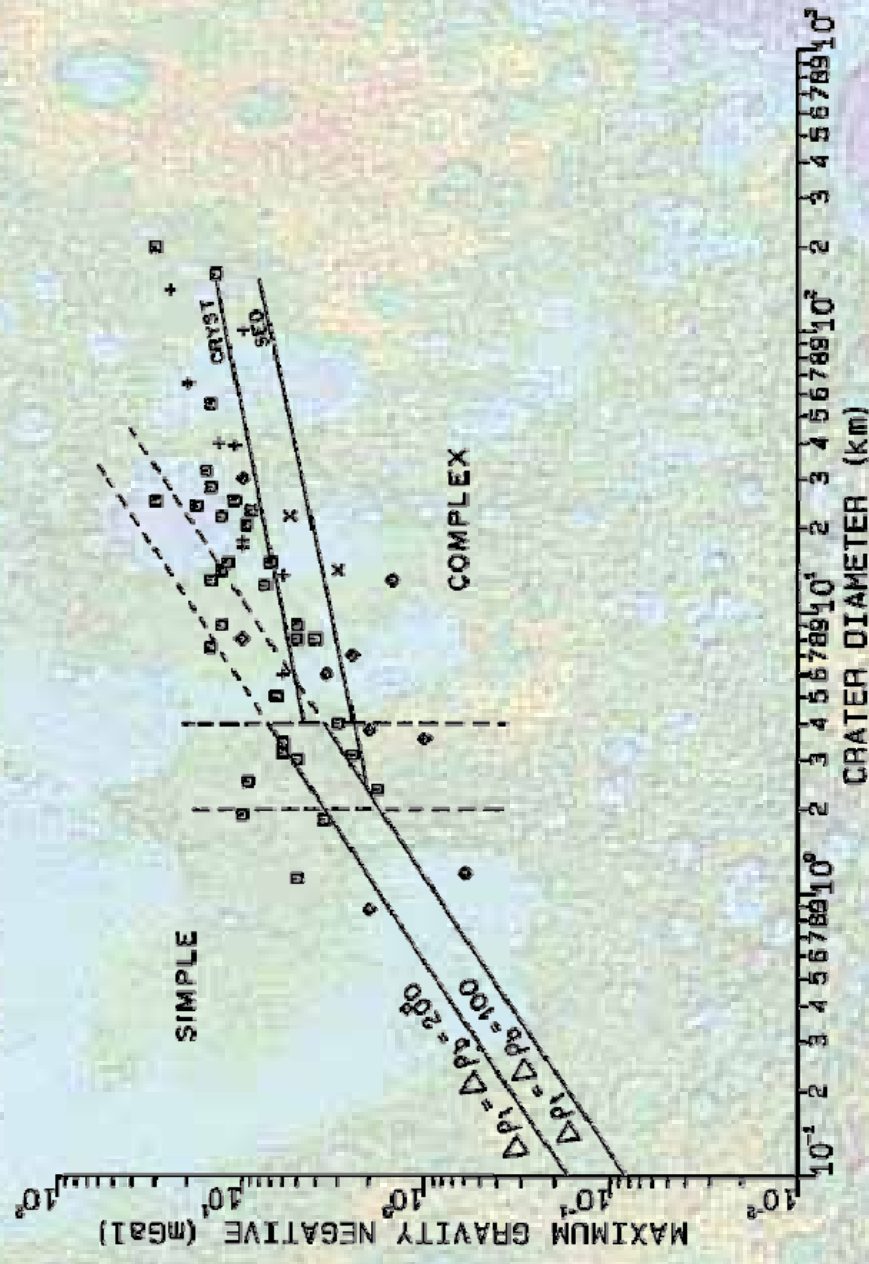


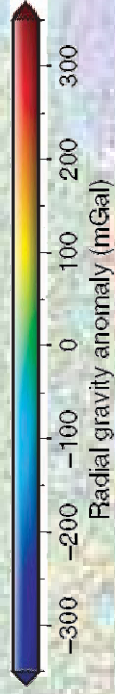
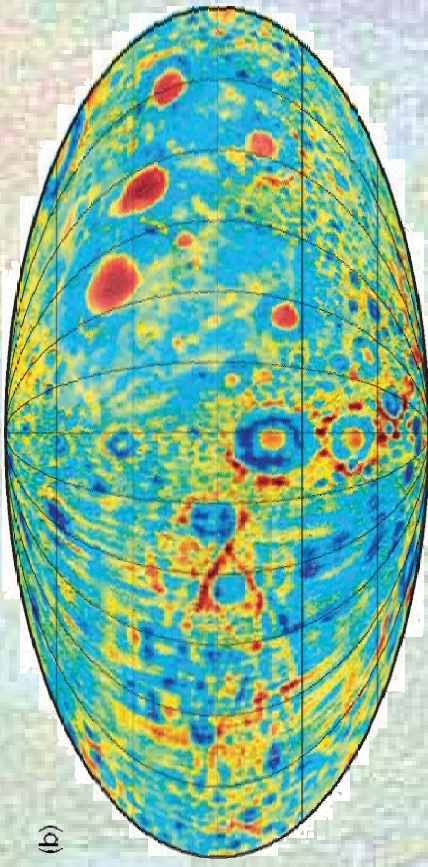
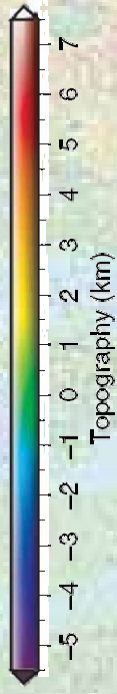
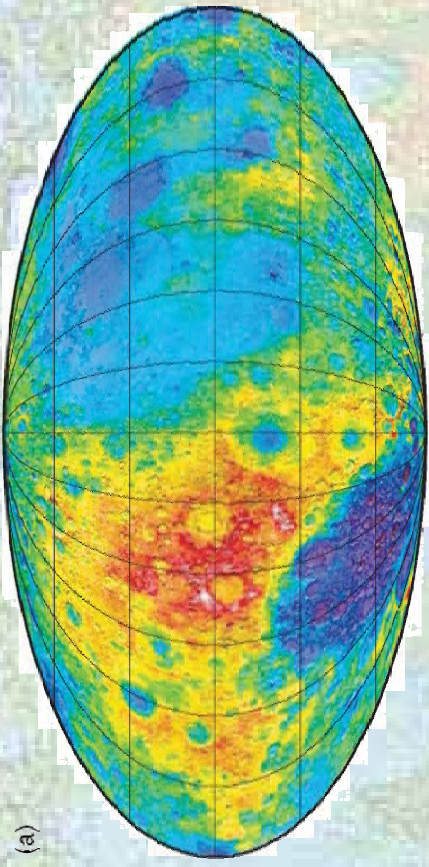
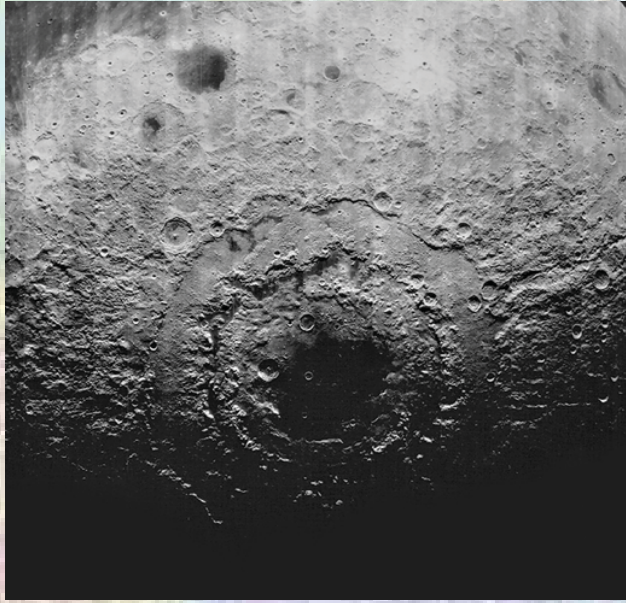
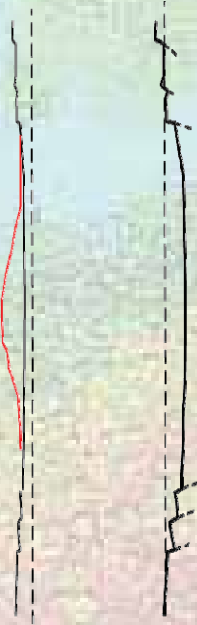
Figure 4. Variation in the maximum negative gravity anomaly with crater diameter. Line marked BAZ shows the variation predicted by the simple hemispherical crater model of Basilevsky *et al.* [1983].

Modelled . . .



Pilkington & Grieve 1992

Mascon



Wieczorek 2007

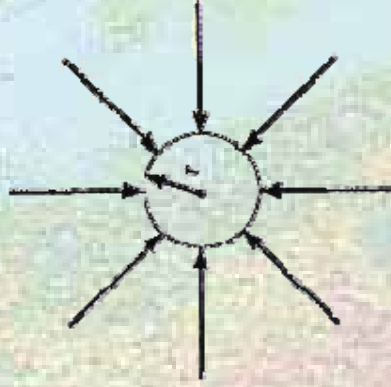
Typical Densities

Table 8.1 Densities of some rock types

	Density (Mg/m ³)
<i>Unconsolidated</i>	
clay	1.5–2.6*
sand, dry	1.4–1.65
sand, saturated	1.9–2.1
<i>Sediments</i>	
chalk	1.9–2.5
coal, anthracite	1.3–1.8
coal, lignite	1.1–1.5
dolomite	2.3–2.9
limestone	2.0–2.7
salt	2.1–2.6
sandstone	2.0–2.6
shale	2.0–2.7
<i>Igneous and metamorphic</i>	
andesite	2.4–2.8
basalt	2.7–3.0
gneiss	2.6–3.0
granite	2.5–2.8
peridotite	2.8–3.2
quartzite	2.6–2.7
slate	2.6–2.8
<i>Minerals and ores</i>	
barite	4.3–4.7
chalcopyrite	4.1–4.3
galena	7.4–7.6
haematite ore	4.9–5.3
magnetite ore	4.9–5.3
pyrite	4.9–5.2
sphalerite	3.5–4.0
<i>Other</i>	
oil	0.6–0.9
water	1.0–1.05

*The ranges of values (taken from a variety of sources) are approximate. Densities depend partly on whether the rock is weathered and the degree of its porosity.

Potential fields

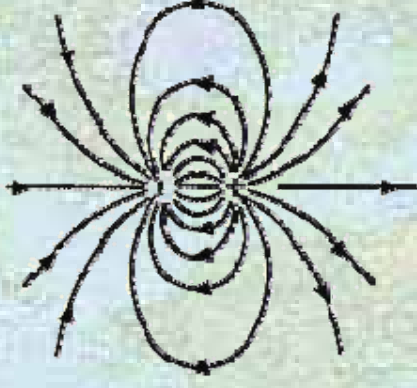


monopole

$$g = \frac{Gm}{r^2}$$

$$m = \text{mass}$$

Gravity Field



dipole

$$B_r = \frac{2m \cos \theta}{r^3}$$

$$B_\theta = \frac{m \sin \theta}{r^3}$$

m = dipole moment

Magnetic Field



Gravity – Magnetic Field

monopolar

attractive (+ or -)

$1/R^2$ dependency

density variations

reflects a bulk rock property

Source depth is very non-unique

stable

dipolar

attractive and repulsive (+ AND -)

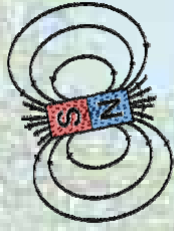
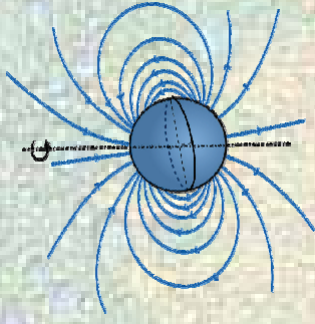
$1/R^3$ dependency

variation of magnetic susceptibility

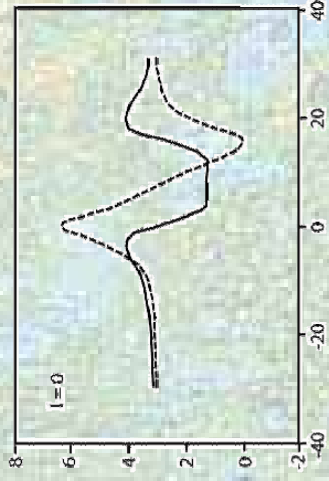
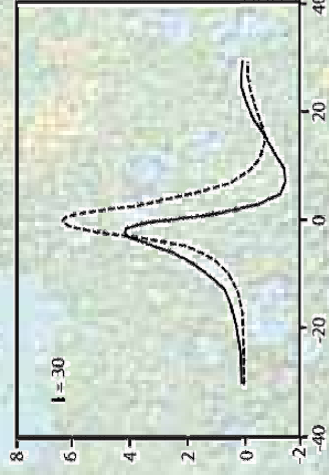
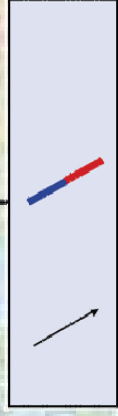
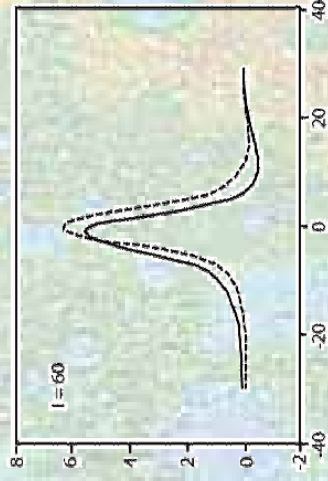
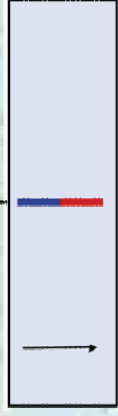
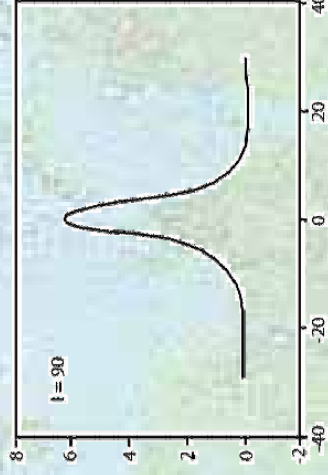
reflects presence of trace elements

source depth within 10-20% accuracy

time dependent



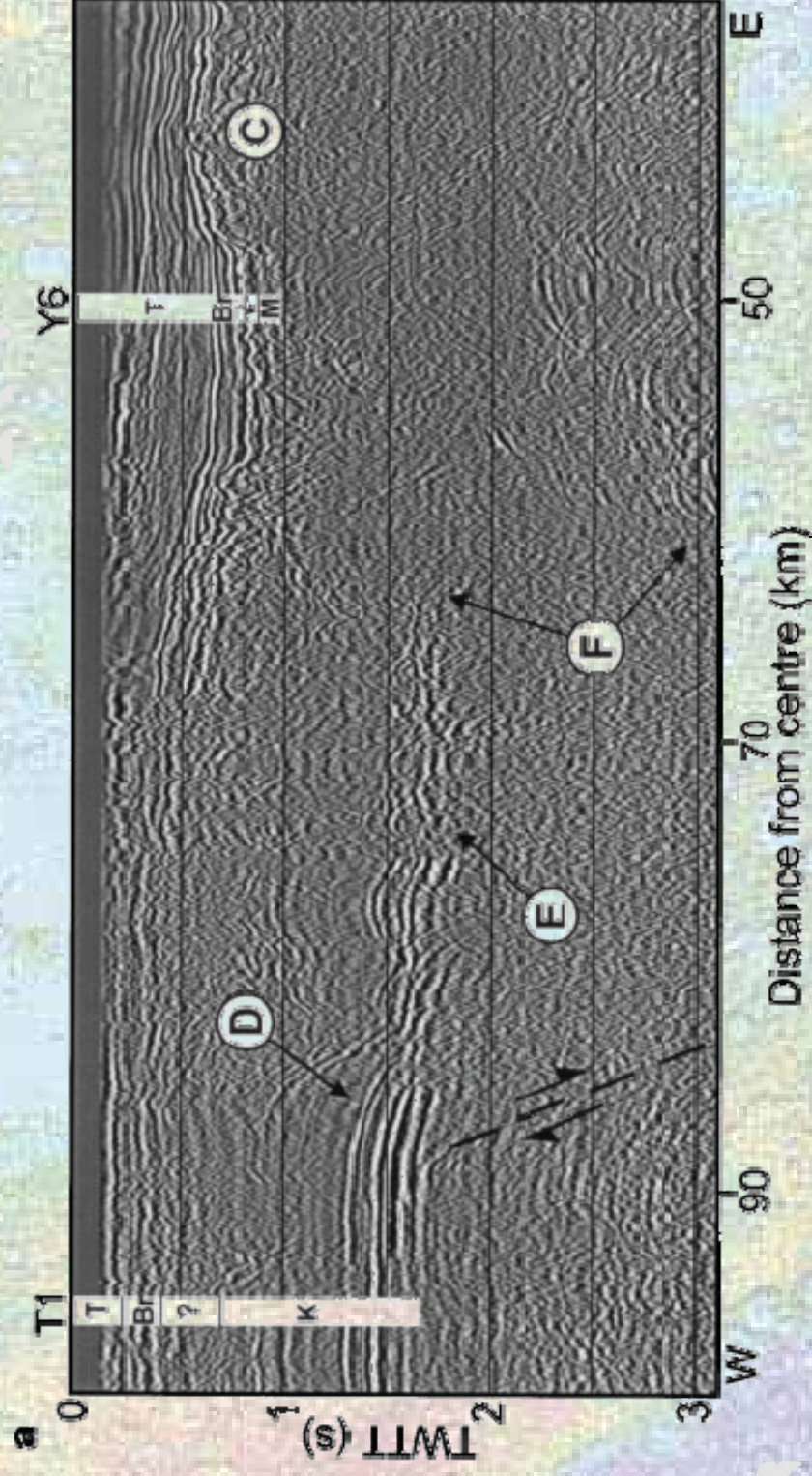
Magnetic Anomalies



Magnetic anomalies

- more complicated than gravity
- no specific signature, but
 - anomalous low or random signature interrupt the regional magnetic pattern (fragmentation and mixing of target rock)
 - strong local magnetic anomaly (melt or uplift)
 - circular
 - if associated to melt, it could be used for palaeomagnetic dating

Seismic Reflection or Refraction



Sound waves (pulses) are sent downward. They are reflected or refracted by layers with different properties in the crust. Different materials have very different sound speeds

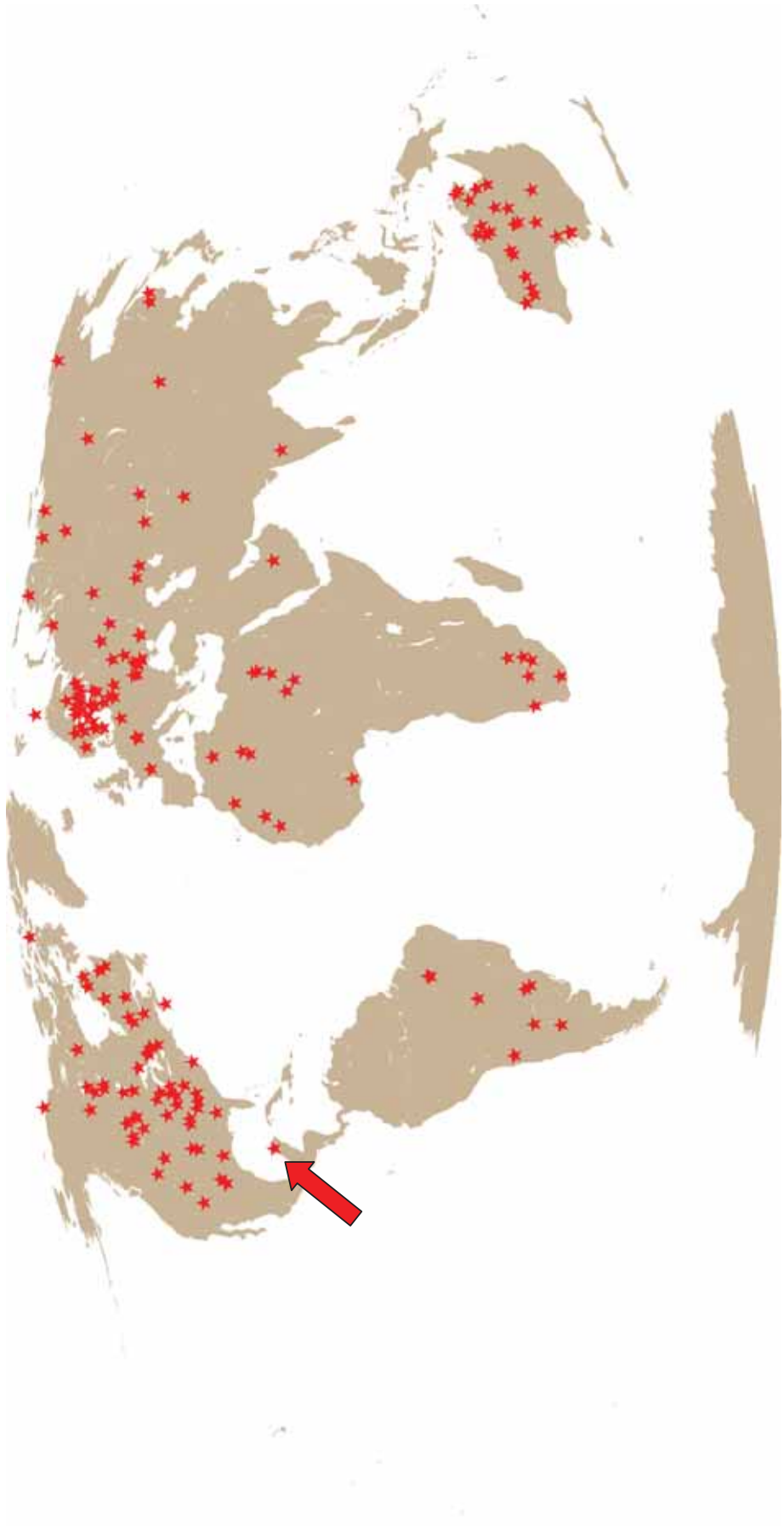
Seismic studies

- refraction and reflection
- reveal subsurface deformation, distinctive for impact structures
 - modest downward and inward displacement of rocks along the edges
 - structural disruption with no coherent seismic reflectors, (roughly corresponding to the transient cavity)
 - presence of reflectors in the central zone at greater depth, is evidence that the structure is shallow and not connected with roots the lower crust or mantle
- essential when buried under younger sediments or water

Electrical Methods

- resistivity variations due impact related target modification
 - impact induced increased porosity and fluid content (low)
 - uplifted material (high)
 - extends farther than the rim

Global Crater Distribution



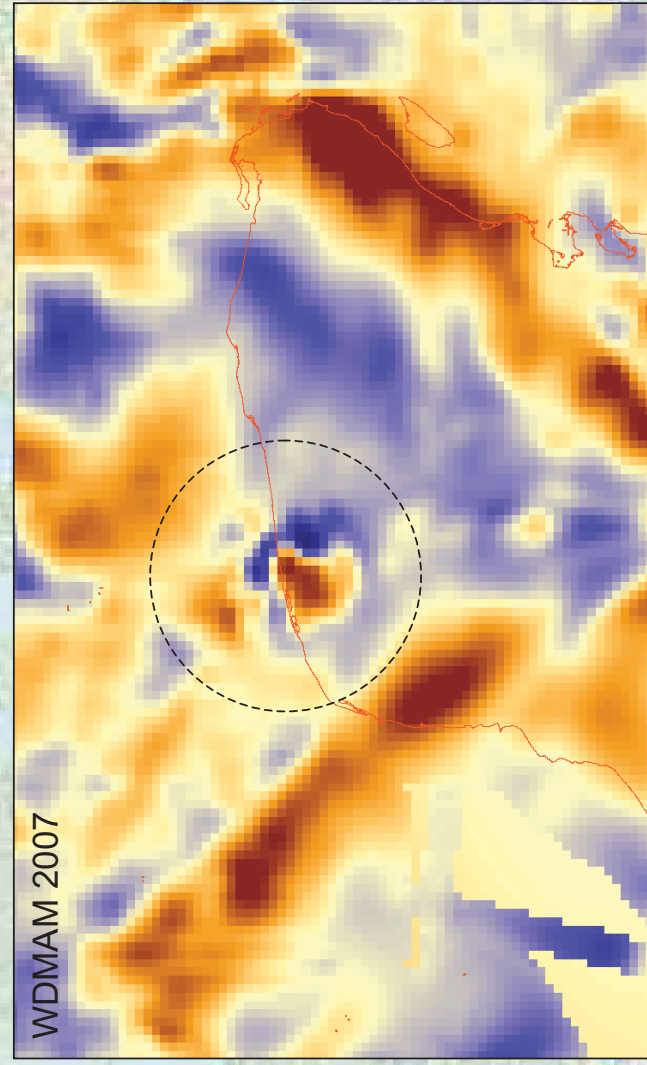
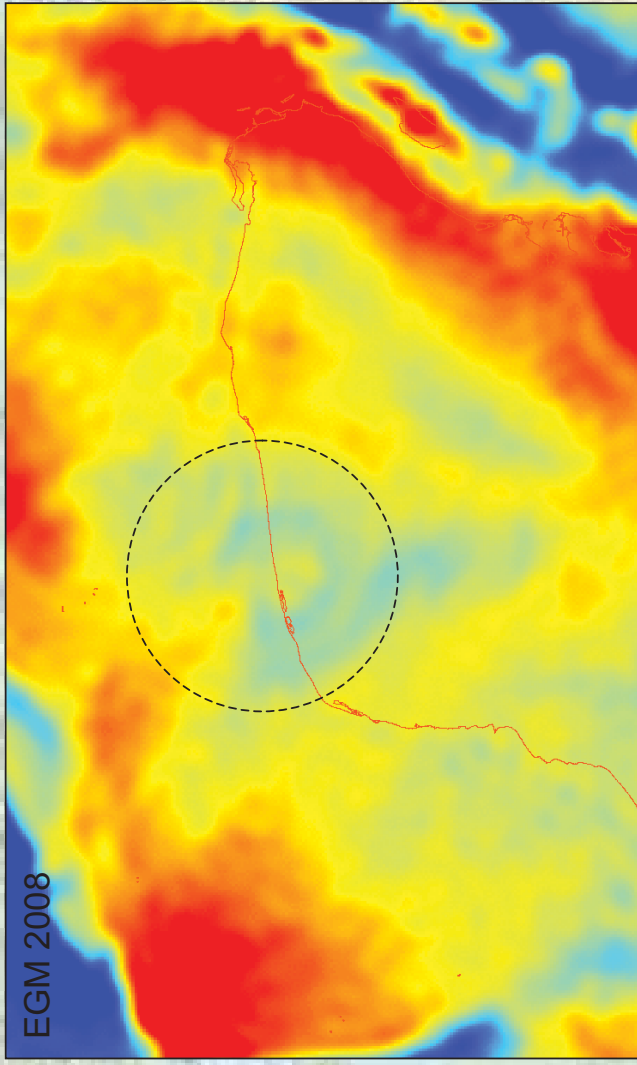
Chicxulub

21°20'N 89°30'W

170 km 64.98 ± 0.05 Ma

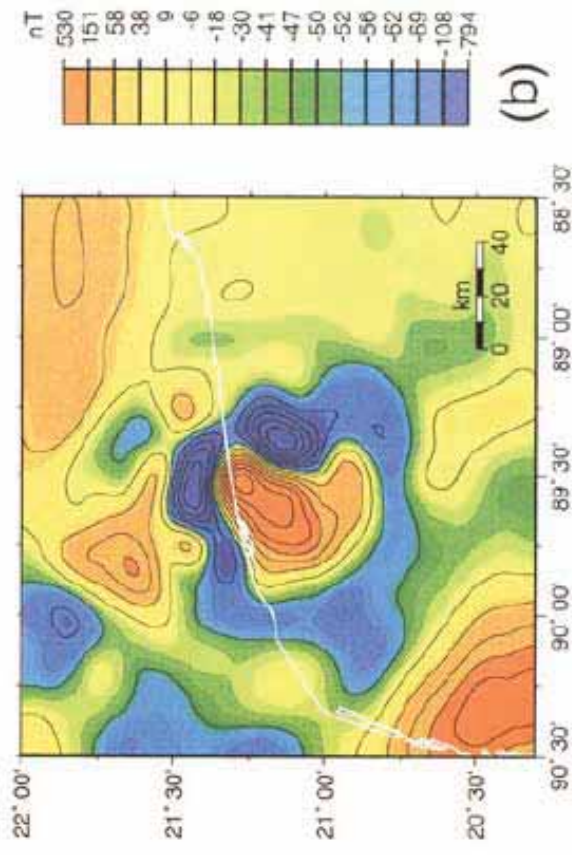
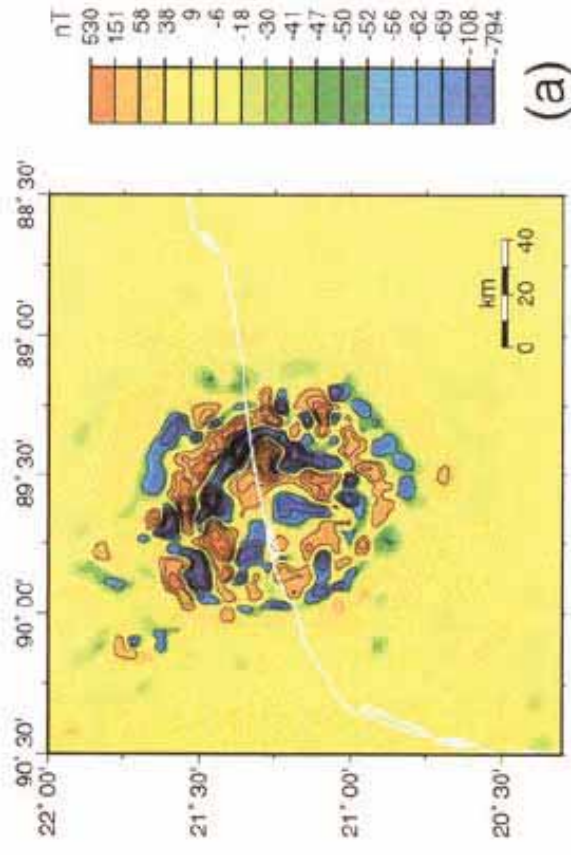


Chicxulub



Chicxulub

filtered magnetic anomalies

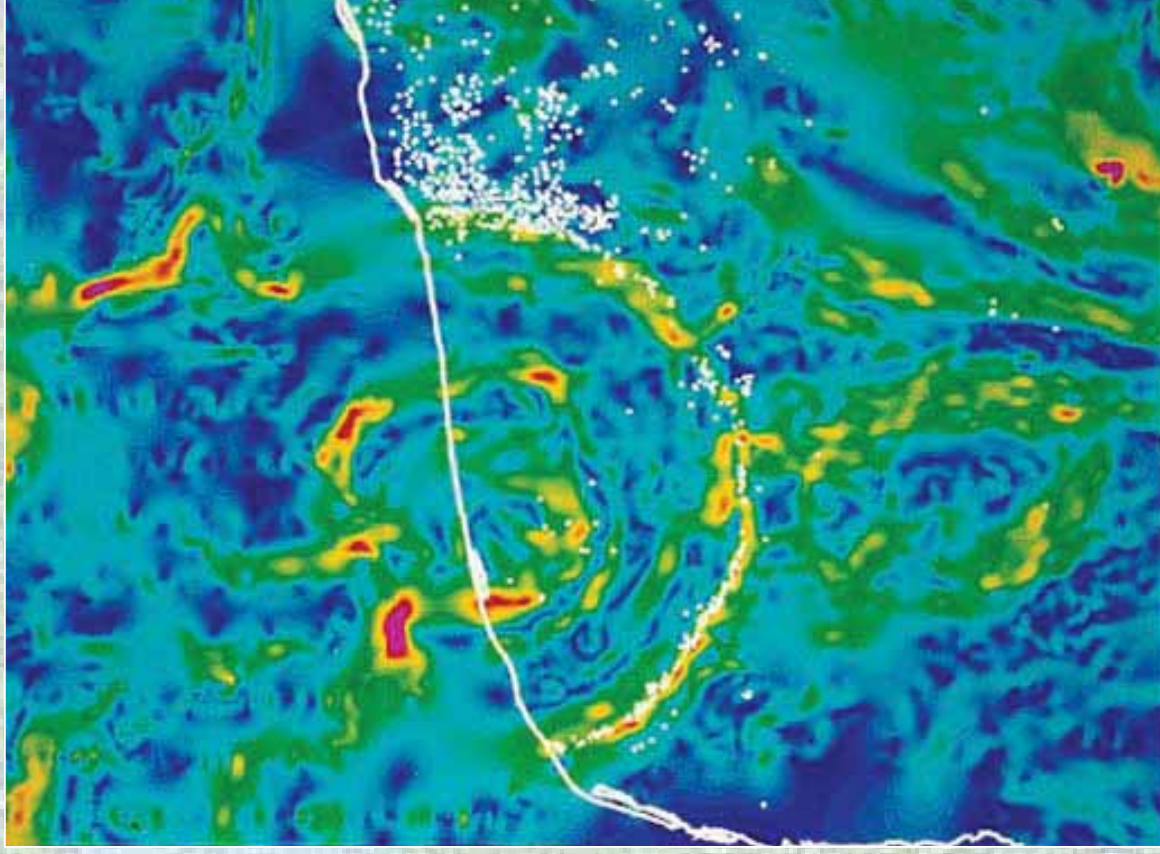


Pilkington & Hildebrand 2000

Chicxulub

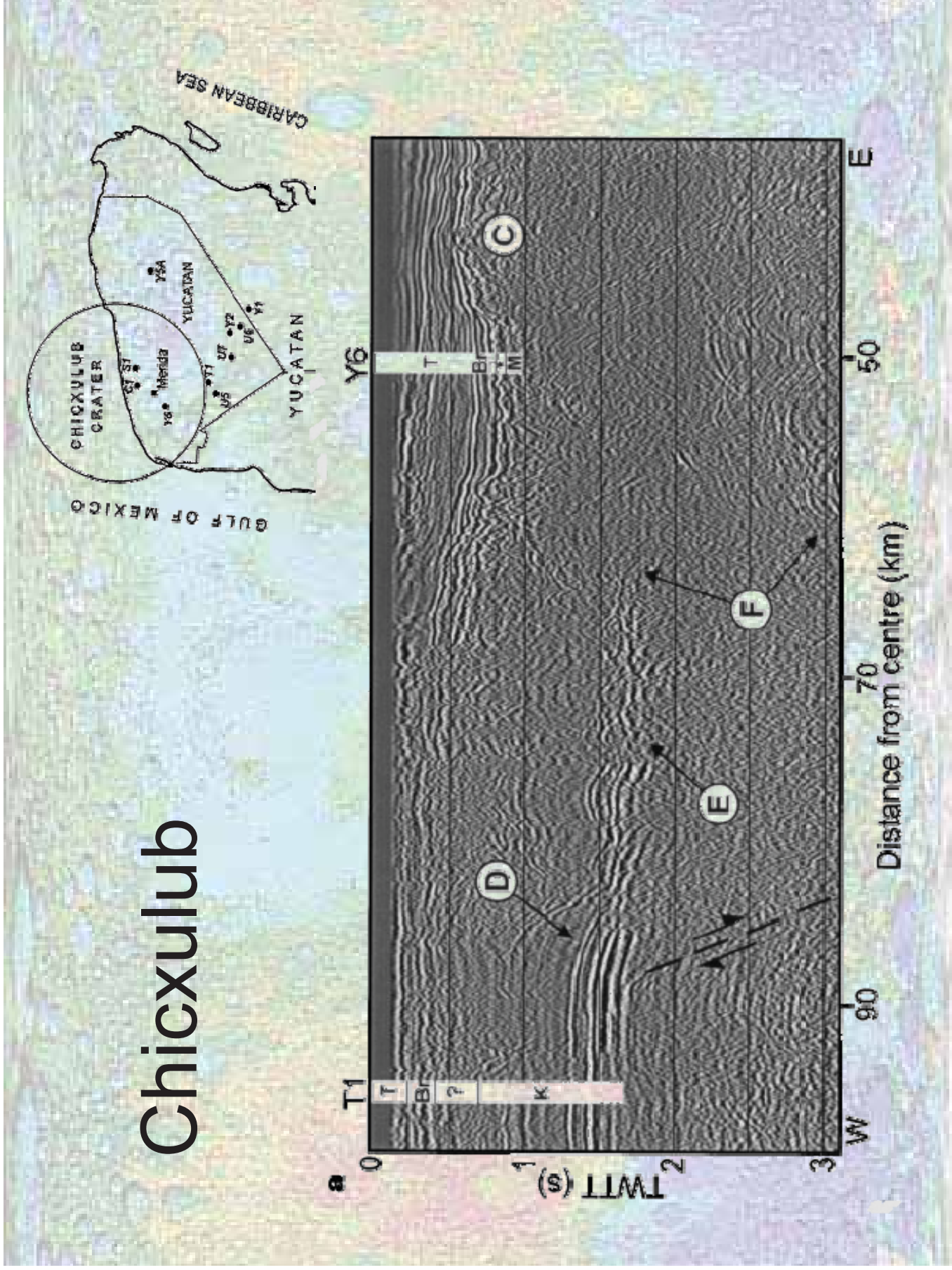
Bouguer gravity horizontal gradients:
highlights lateral density contrasts
and suppresses regional anomalies

Cenote distribution

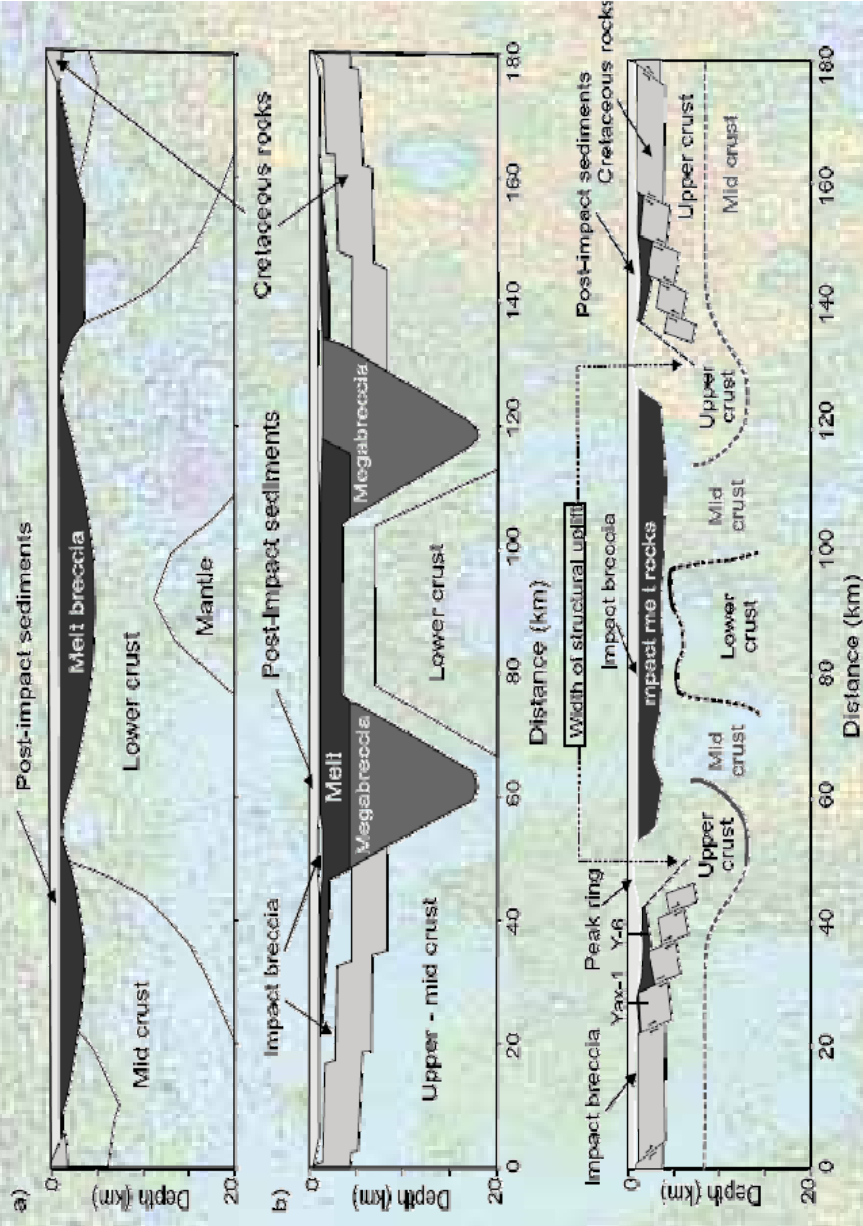


Hildebrand et al. 1995

Chicxulub

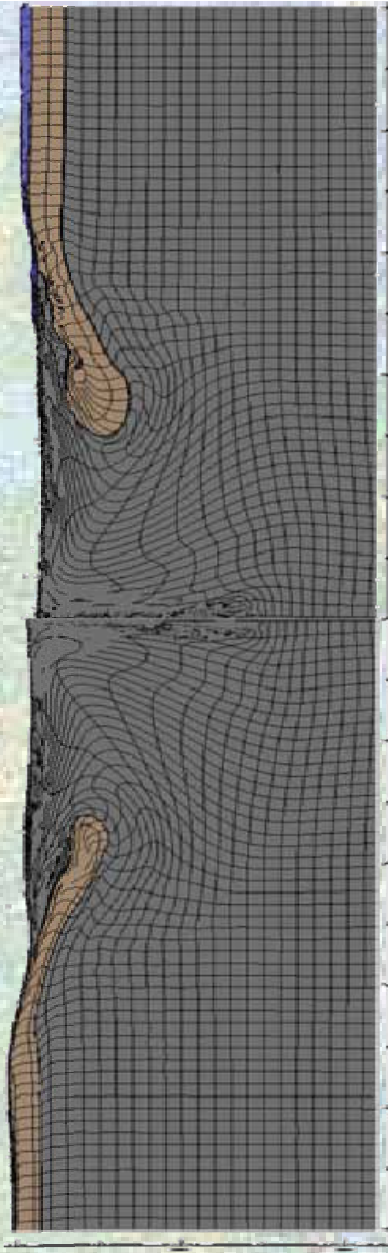


Sharpton et al. 1996



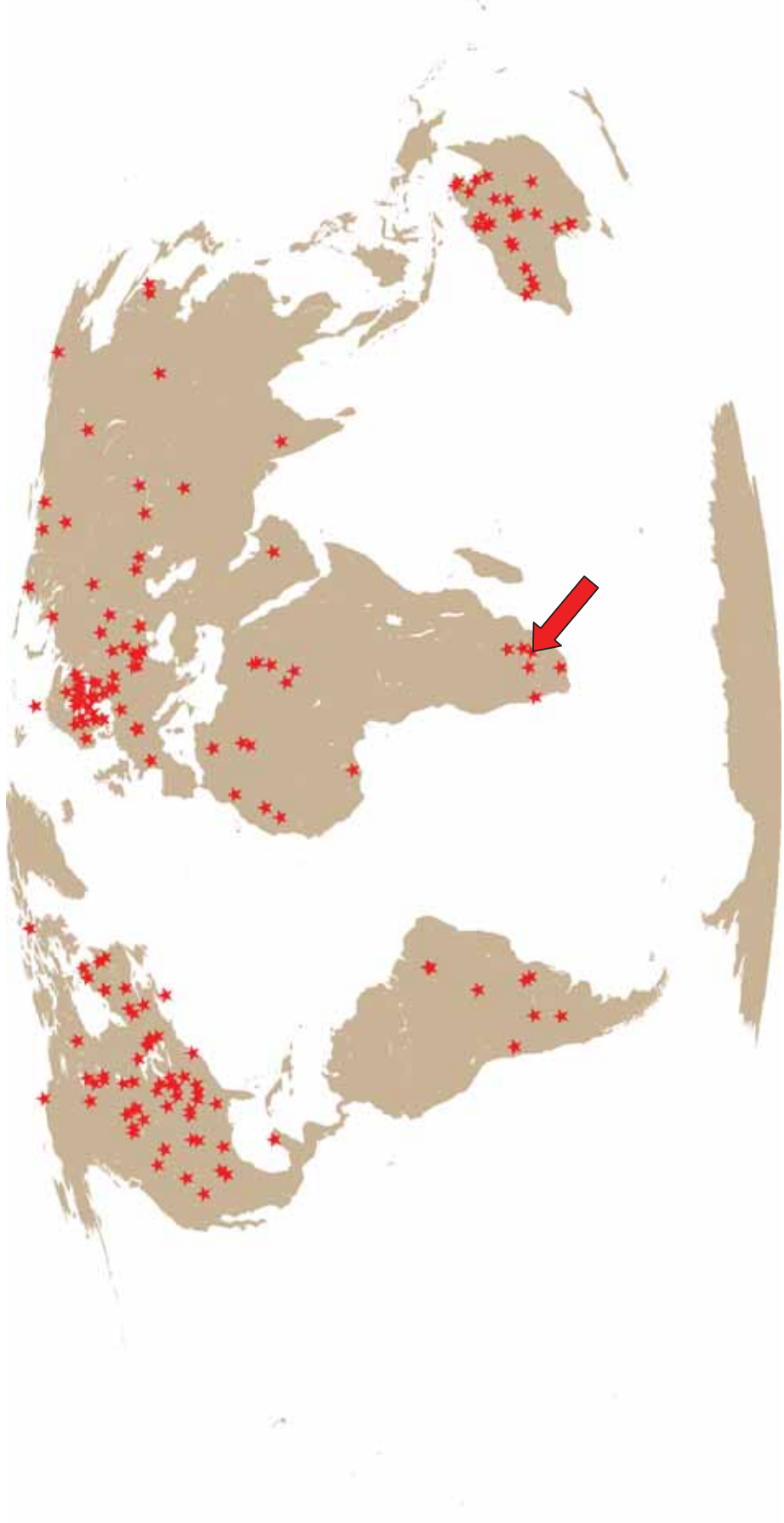
Hildebrand et al. 1998

Vermeesch & Morgan 2008



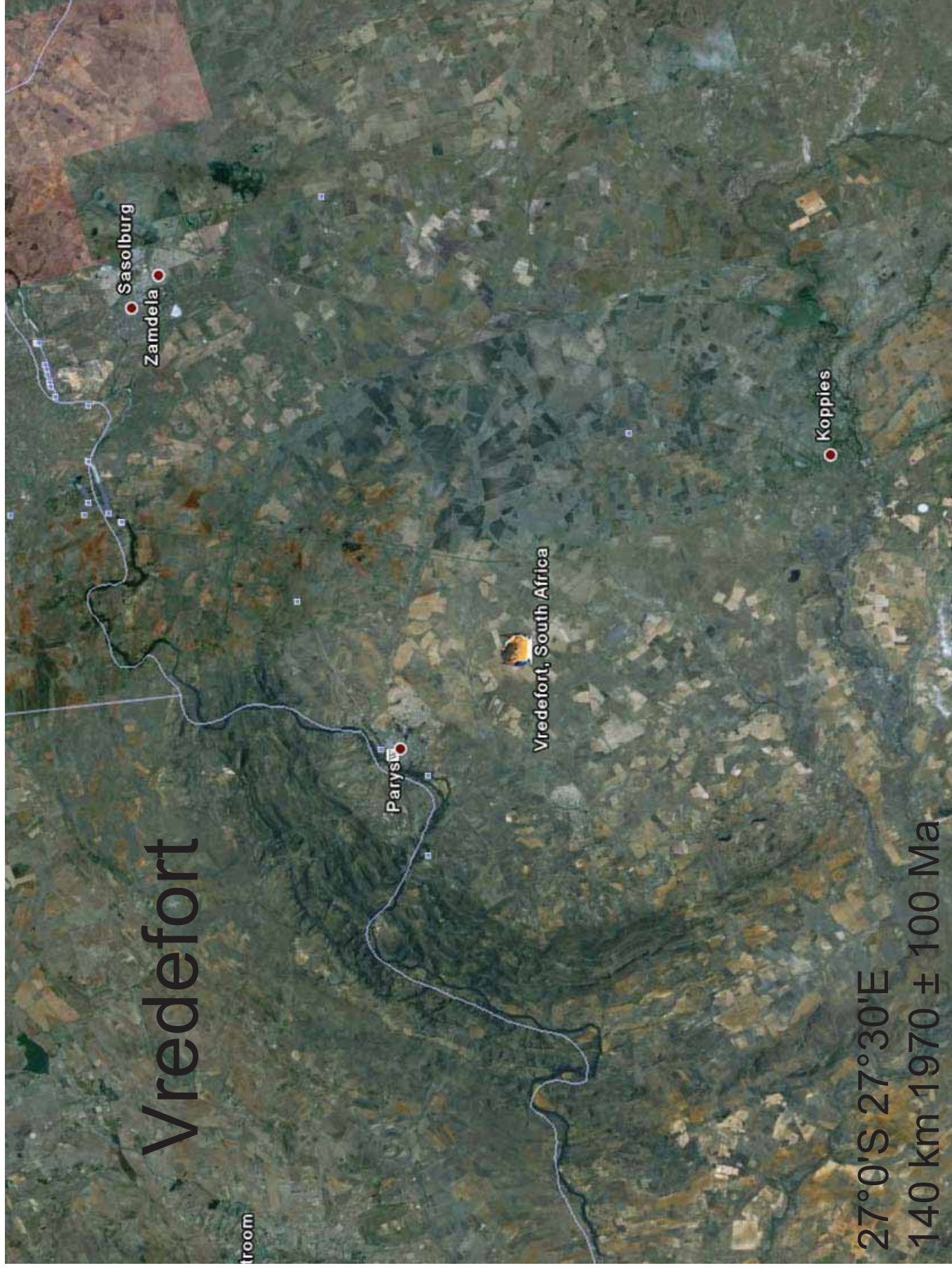
Collins et al. 2008

Global Crater Distribution



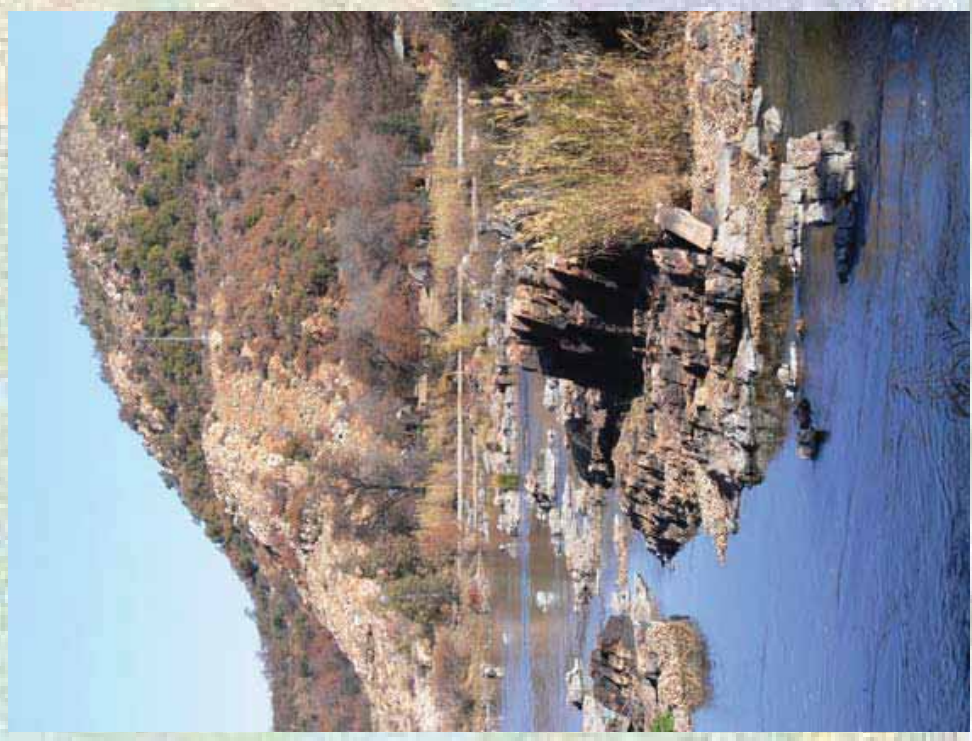
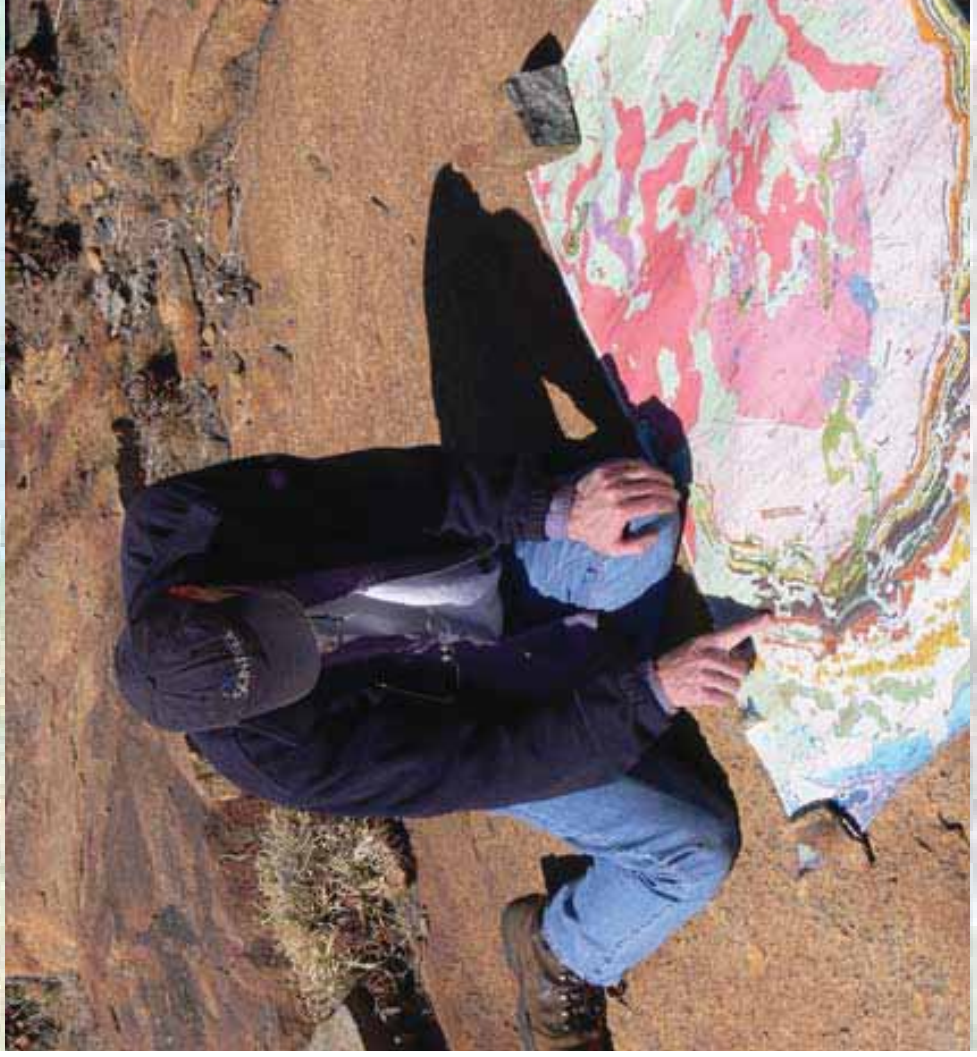
Vredefort

room

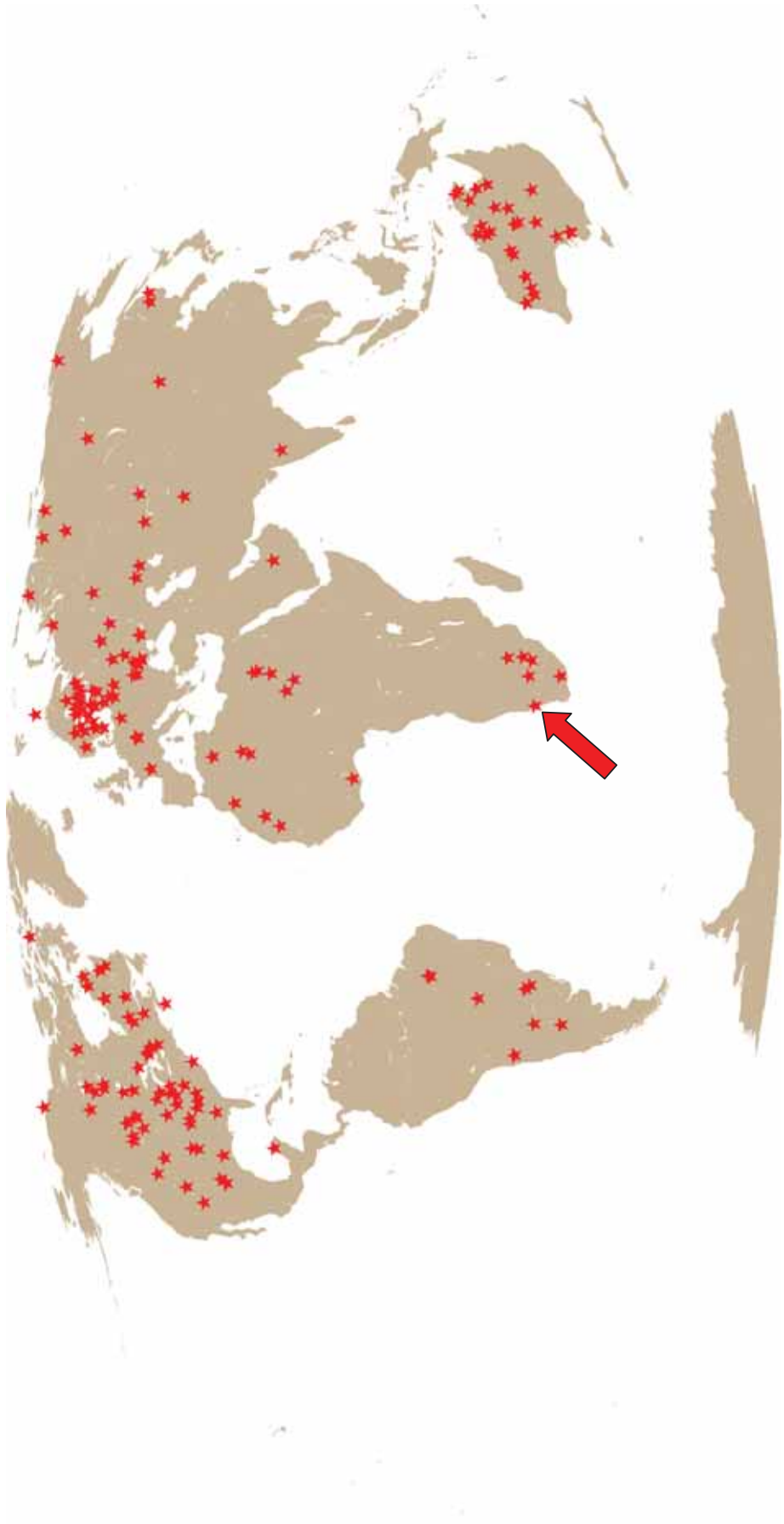


27°0'S 27°30'E
140 km 1970 ± 100 Ma

Geological / Structural Interpretation



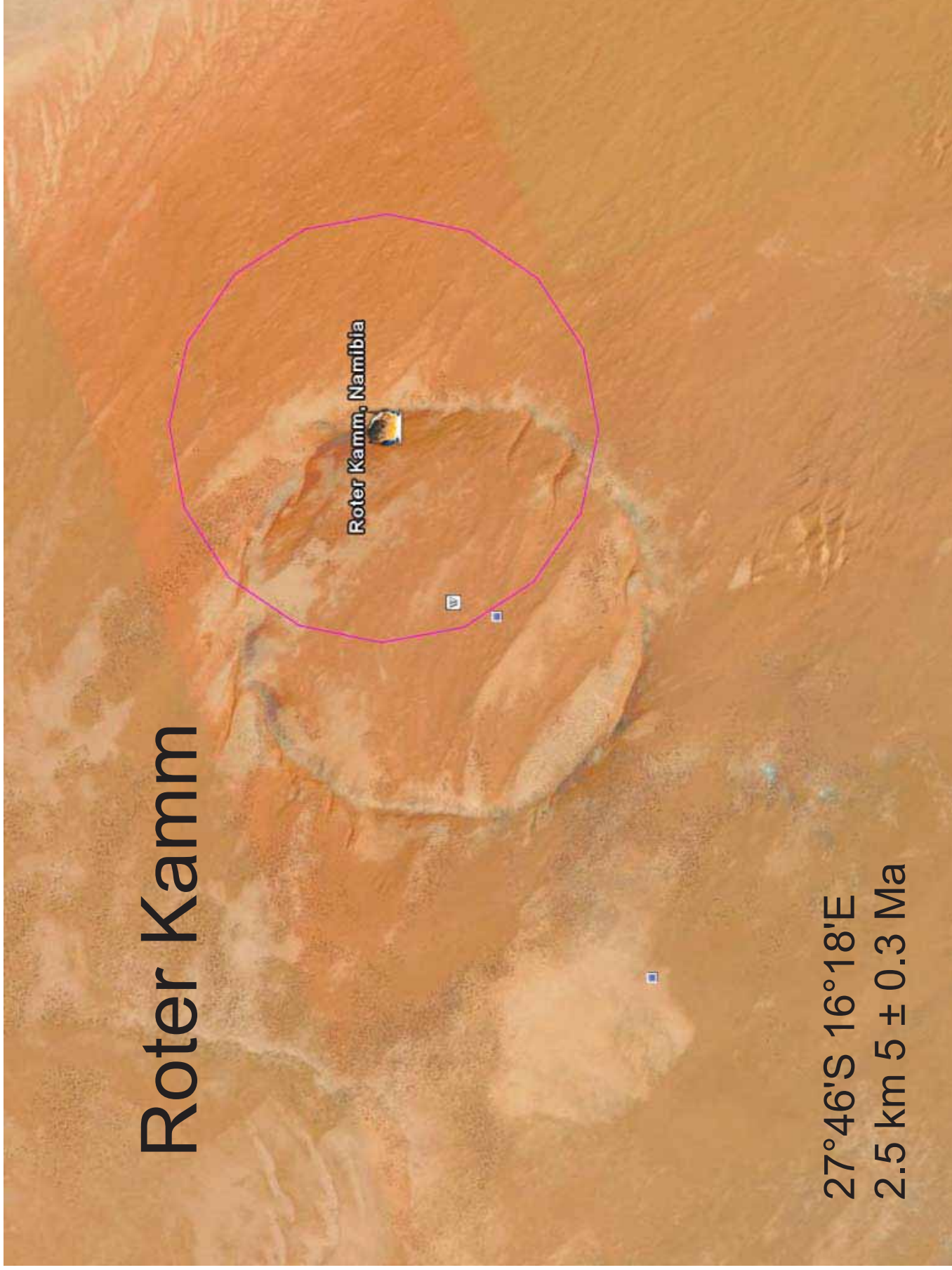
Global Crater Distribution



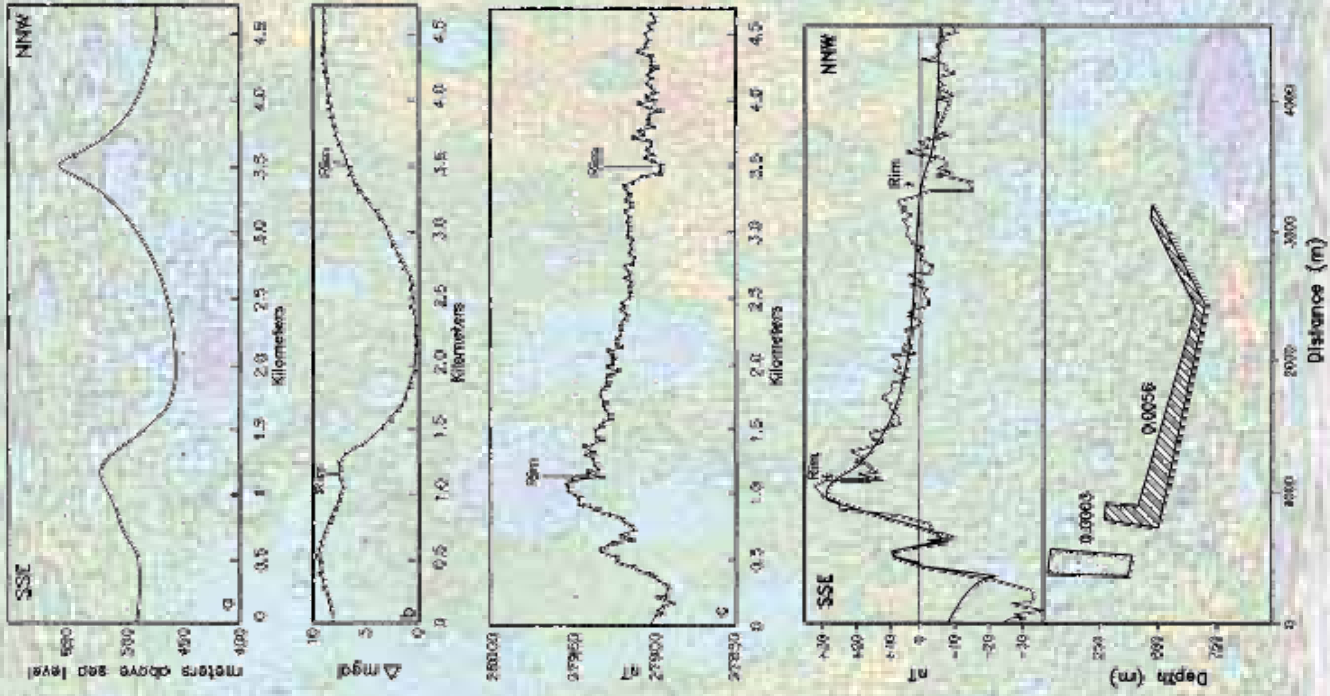
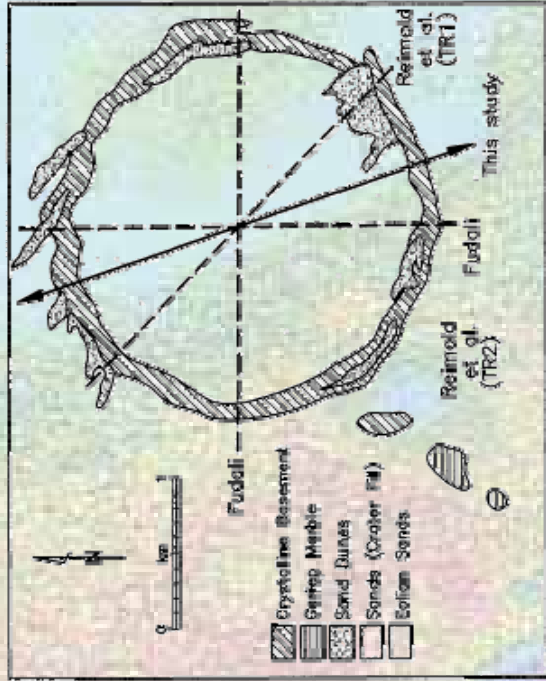
Roter Kamm

Roter Kamm, Namibia

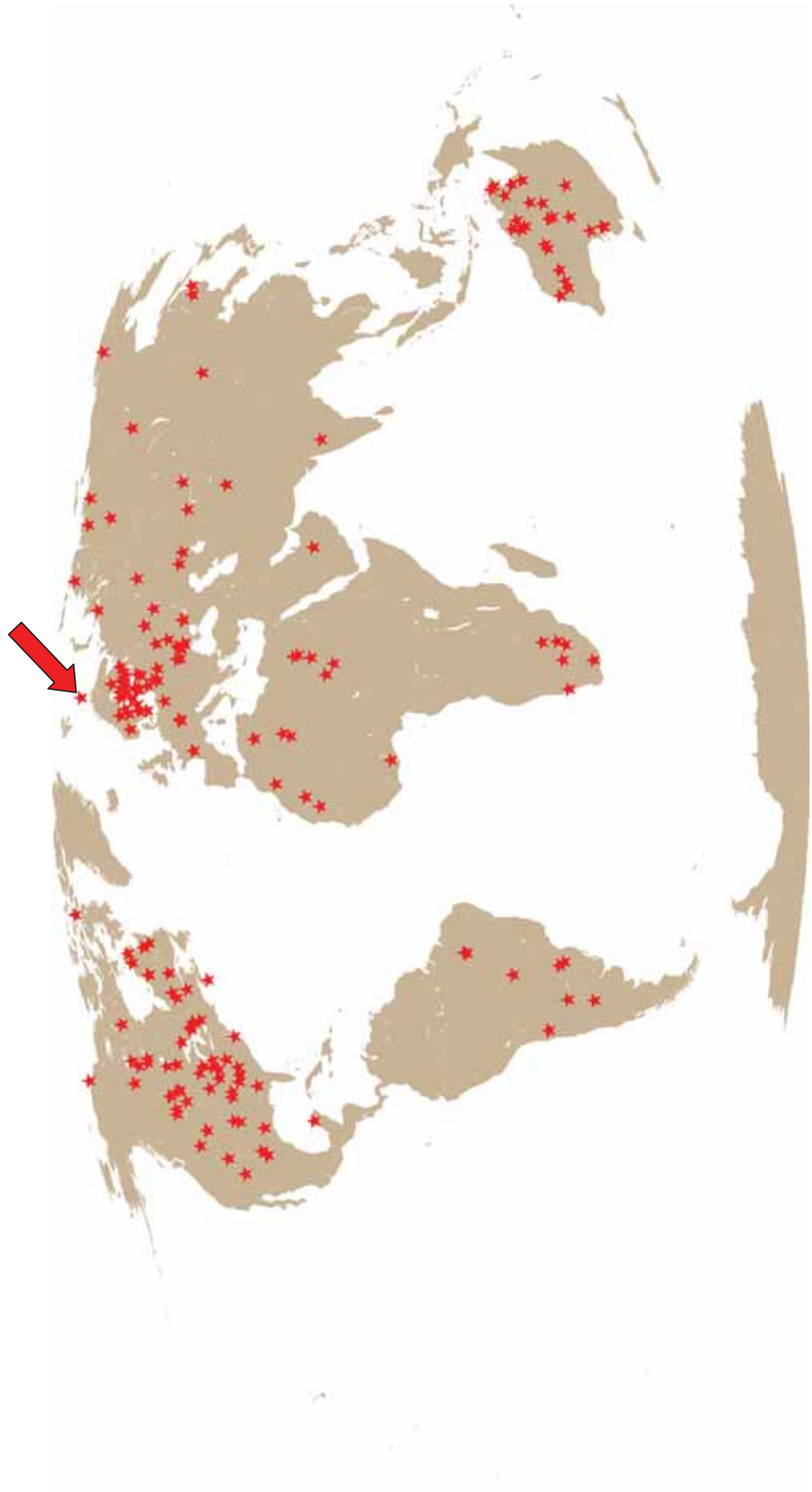
27°46'S 16°18'E
2.5 km 5 ± 0.3 Ma



Roter Kamm



Global Crater Distribution



Mjølnir



Morphology

- 40 km in diameter
- 8-km wide central uplift,
- 4-km wide annular trough
- and a 12-km wide outer zone

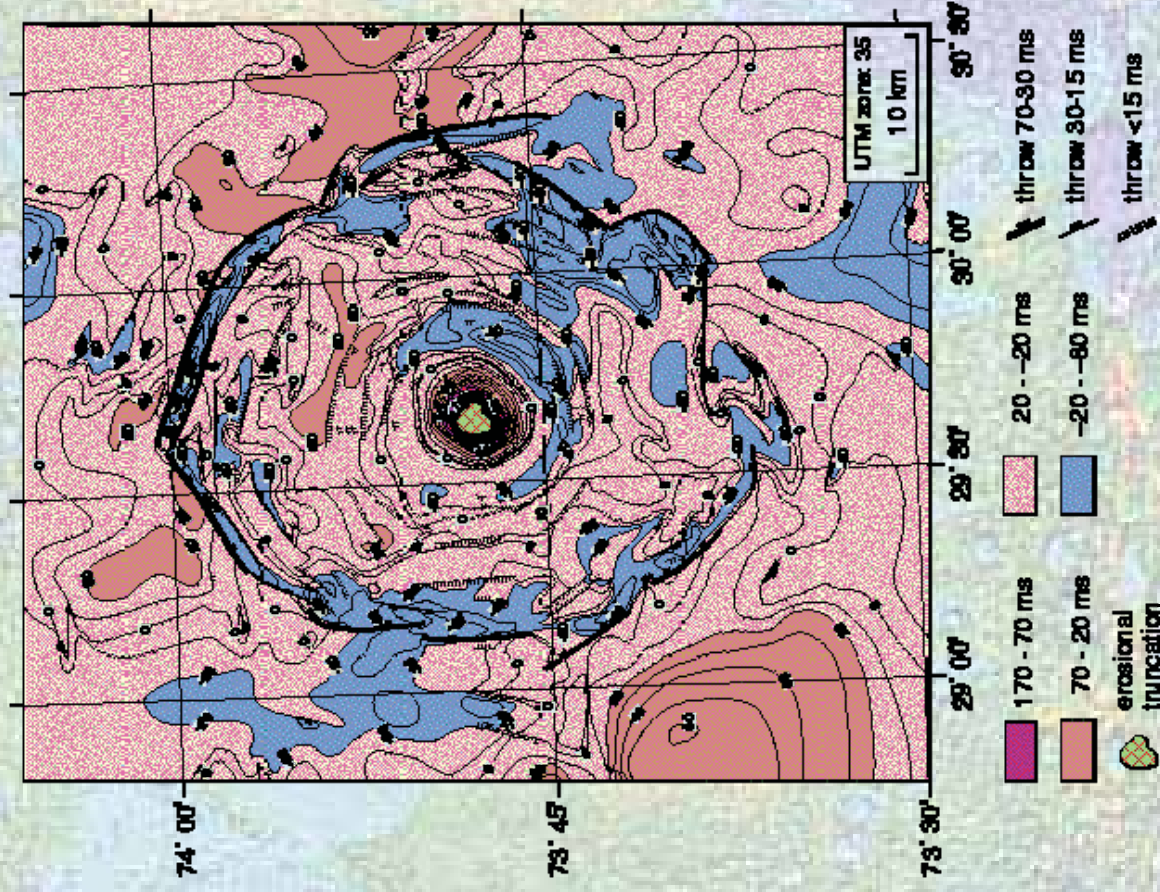
Gudlaugsson, 1993

- apparent depth 30-70 m

Tsikalas et al. 1998

- Volgian-Ryazanian; 142 ± 6 Ma

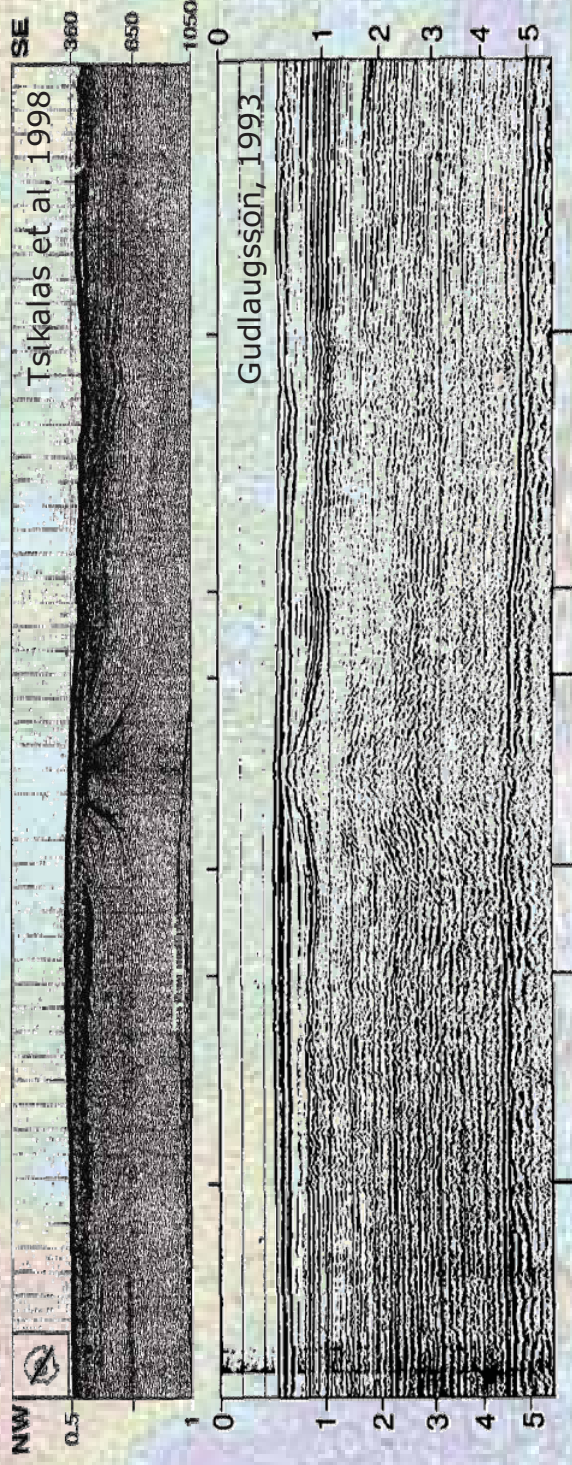
Dypvik et al. 1996; Smelror et al. 2001



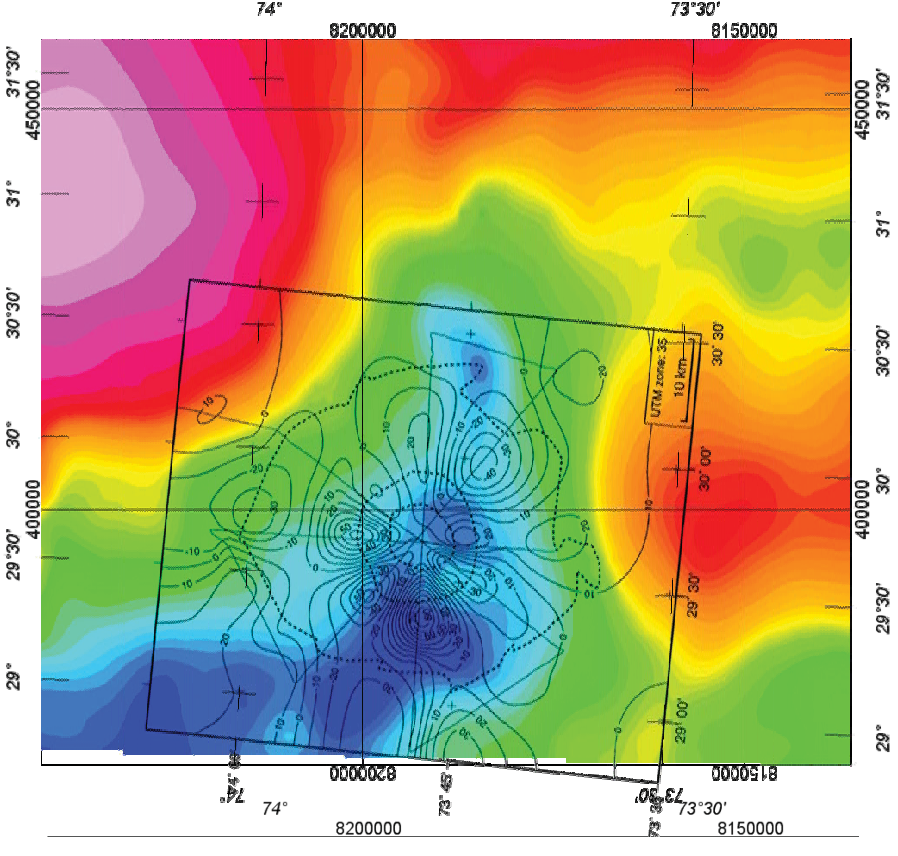
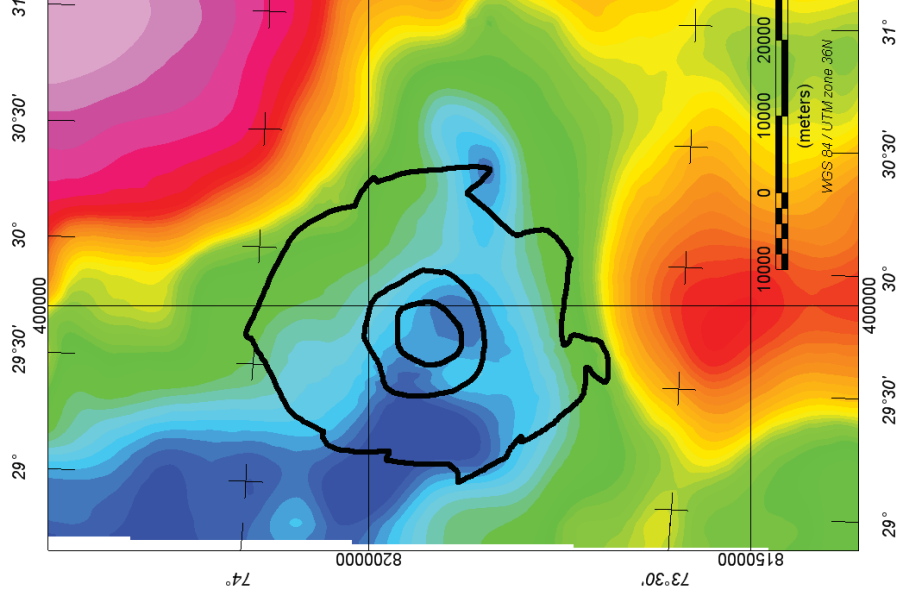
Seismics

Bjarmeland Platform:

- horizontally layered sedimentary sequences of Devonian to Jurassic
- pronounced horizontal reflector marking the top of the Permian succession (Barents-Sea wide)
- Top Permian reflector remains undisturbed by the impact



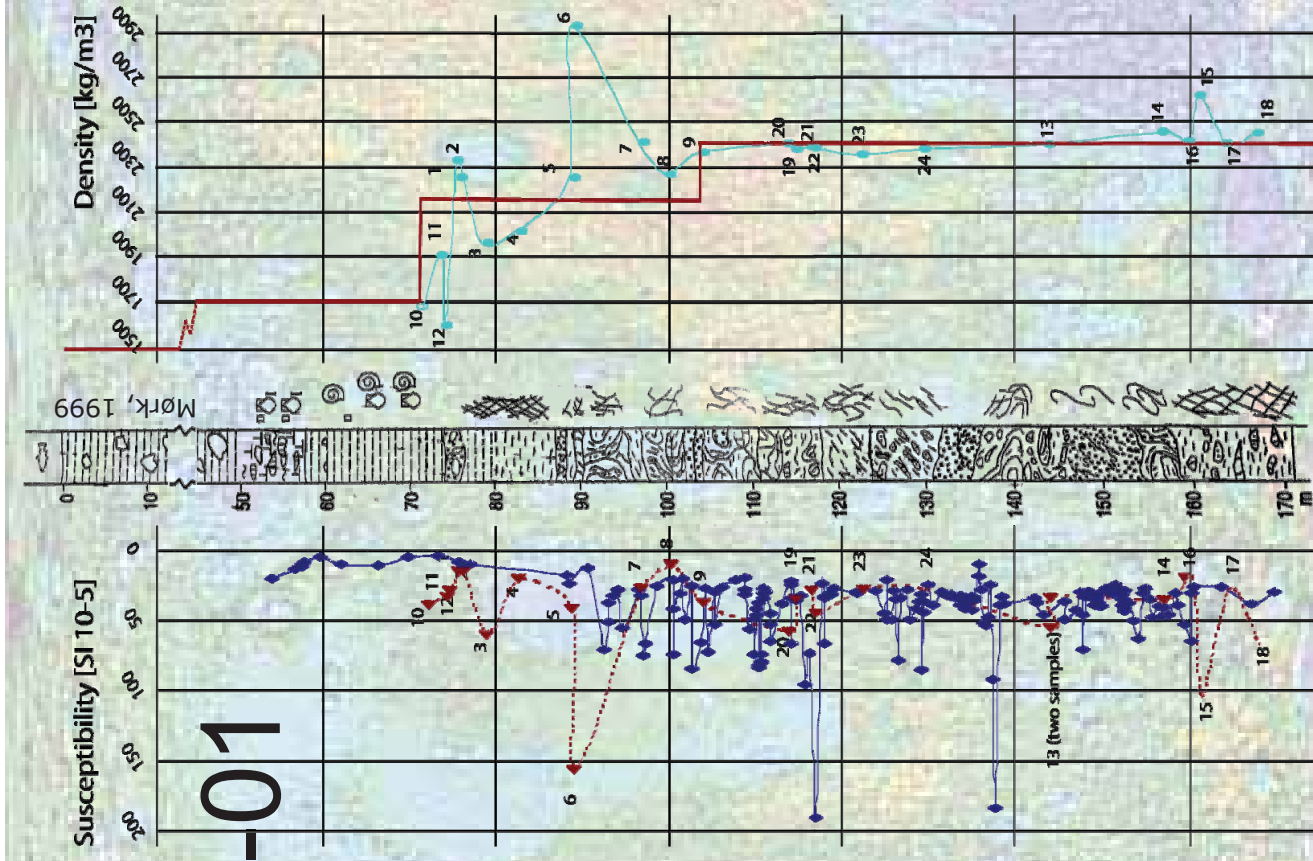
Aeromagnetic Survey (BAS06)



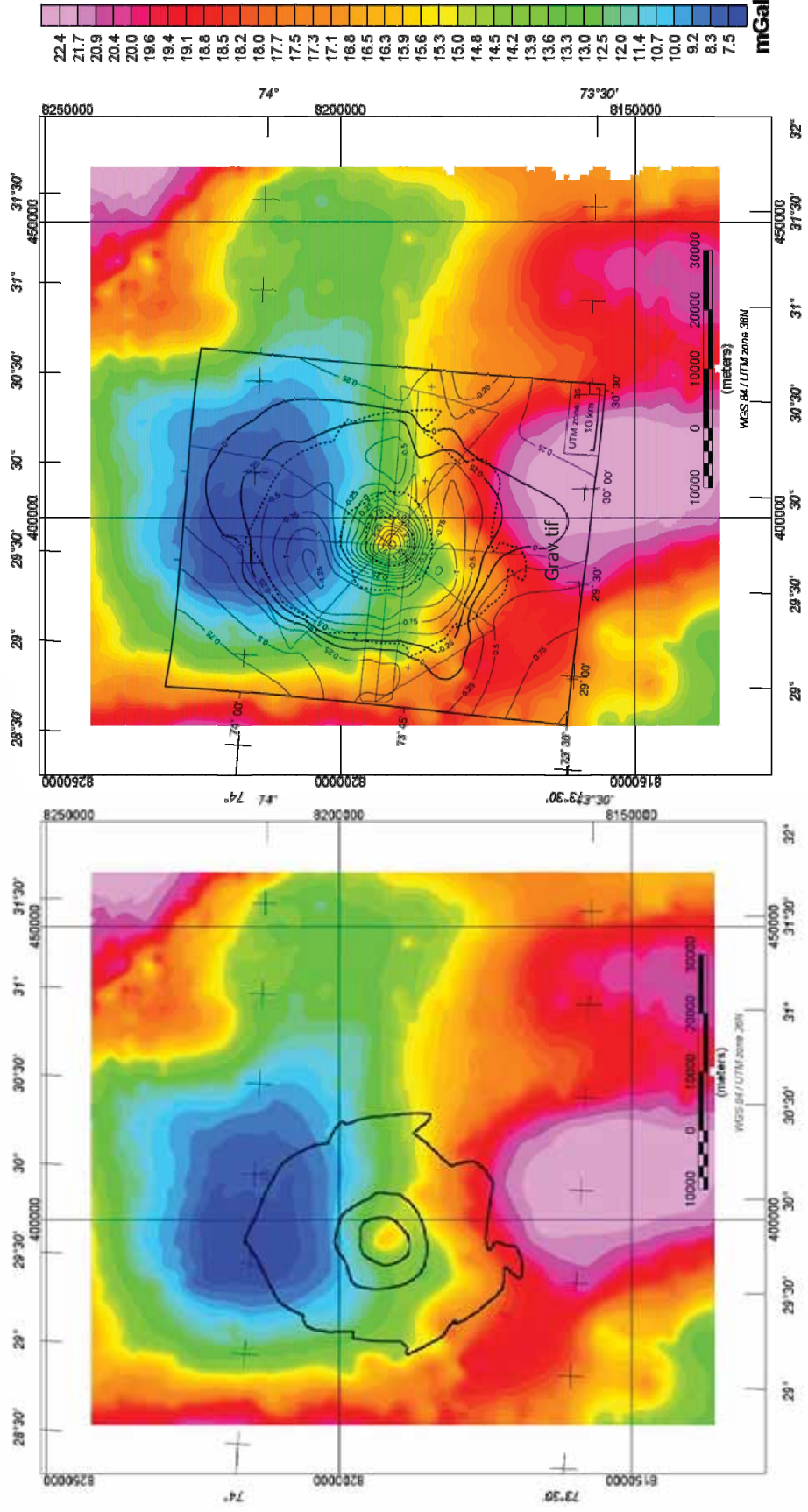
Sponsored by Chevron, EniNorge, Norwegian Petroleum Directorate, RWE-Dea Norge and StatoilHydro.
L. Gernigon and J.O. Mogaard collected and processed the survey data.

Core 7329/03-U-01

- close to structure centre, depth 171m
- total susceptibilities: Bartington MS2c sensor in core-scanning set-up
- volume-specific susceptibilities
- range: 0-200 SI10⁻⁵; typical values for marine sedimentary rocks
- peak values found for siderite-cemented beds or nodules
- other sources are detrital magnetic minerals



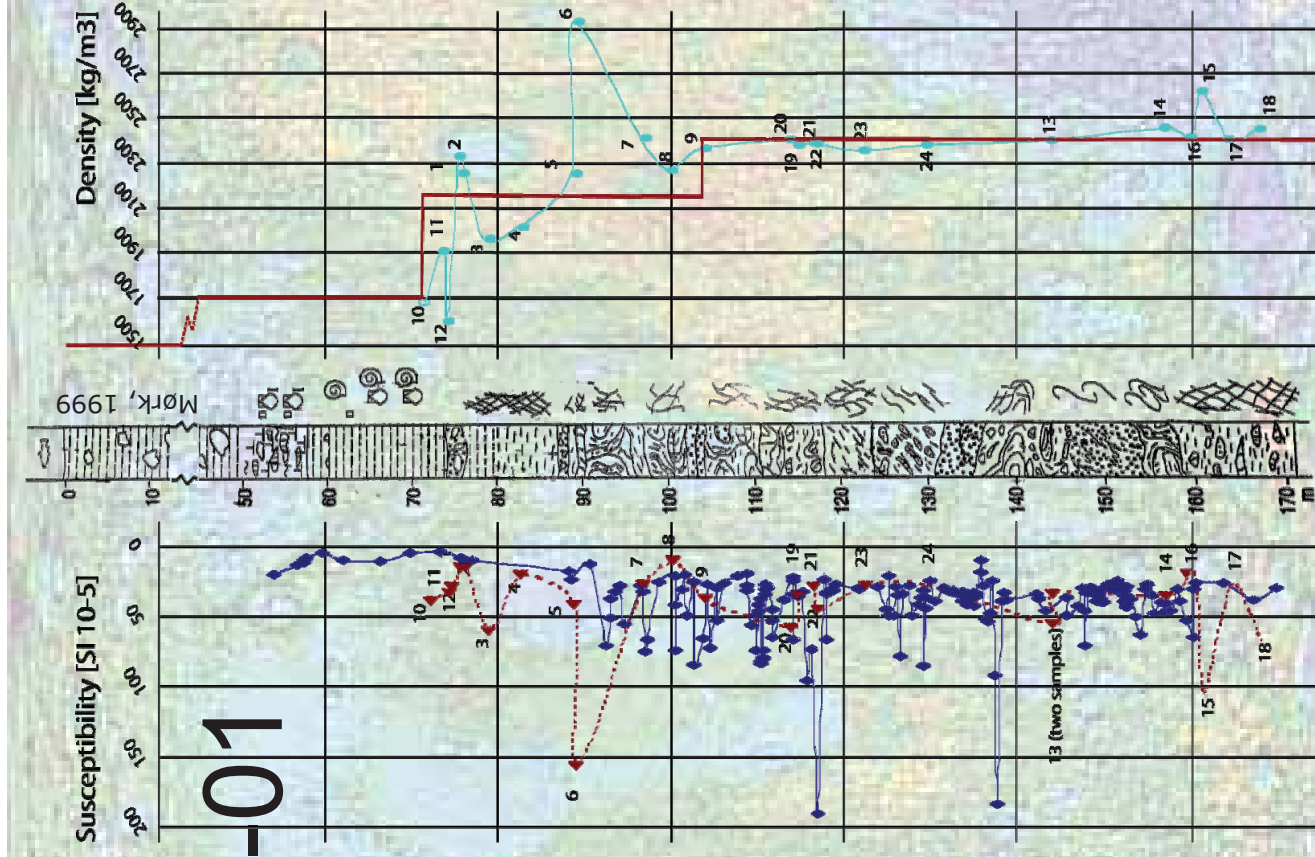
Gravity



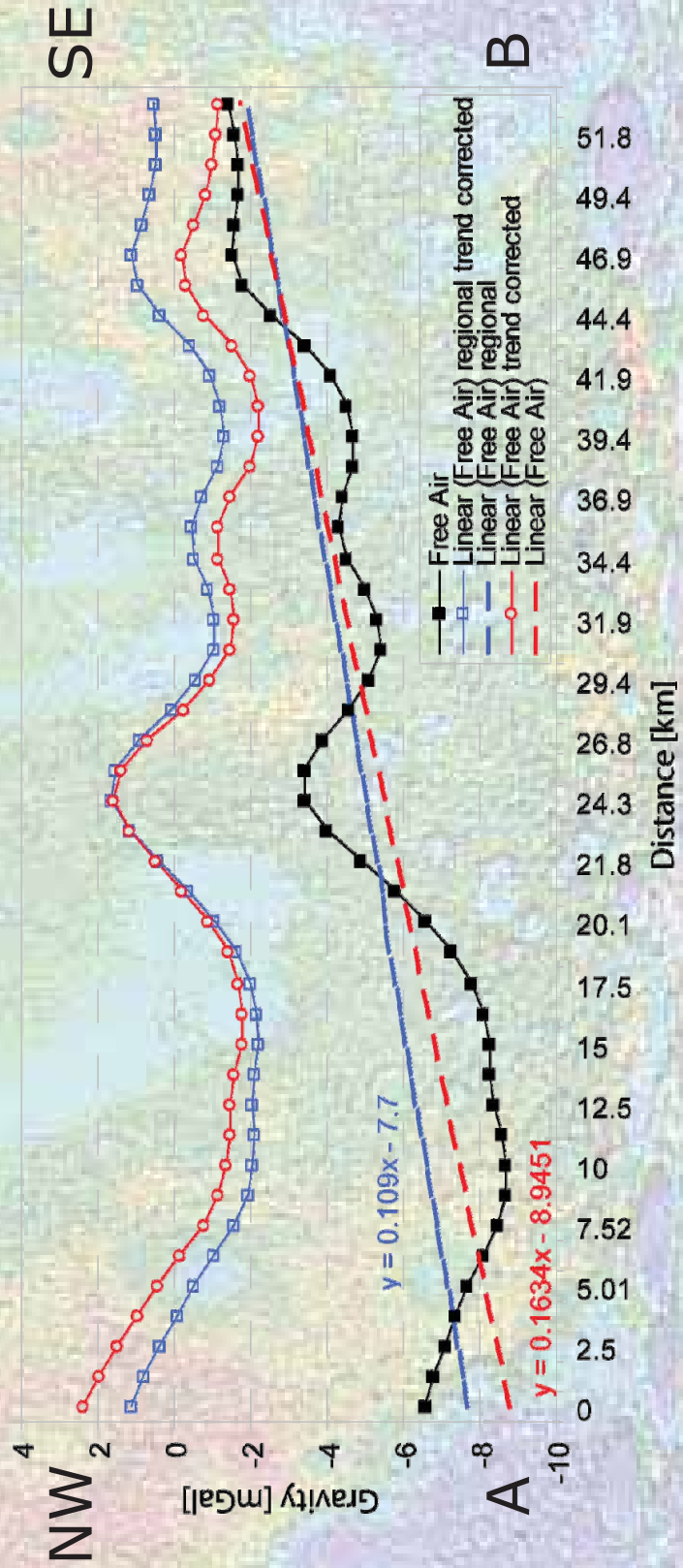
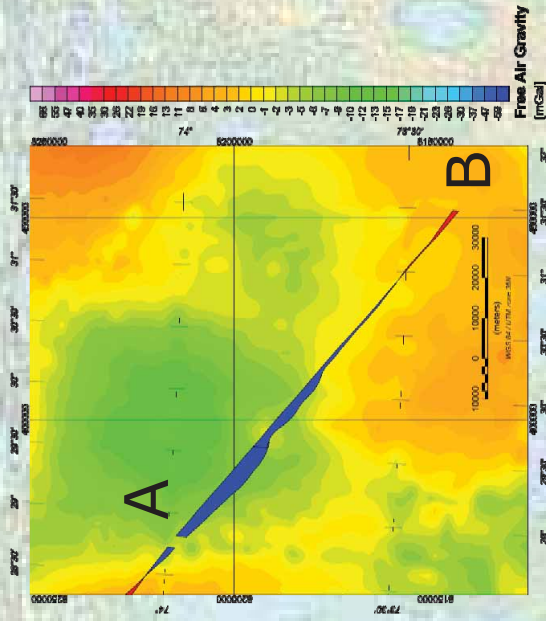
NGU database

Core 7329/03-U-01

- close to structure centre, depth 171m
- dry bulk densities (most samples dissolve in water)
- range 1500 –2900 kg/m³
- were used to constrain the uppermost layers of the gravity model



Regional Trend

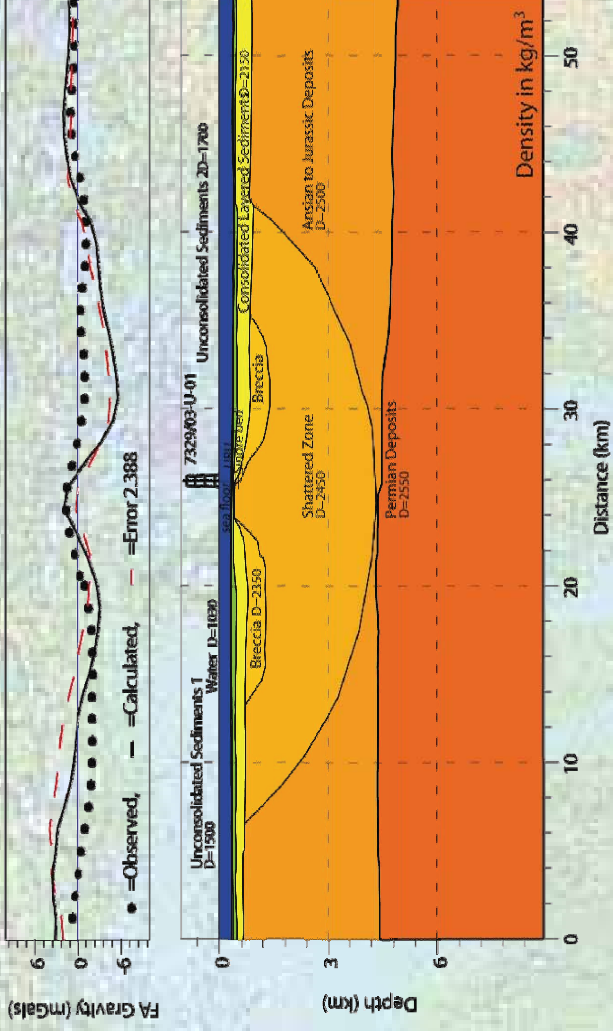


Modelling Results

- Water: 1030 kg/m³
- Unconsolidated Sed. 1: 1500 kg/m³
- Unconsolidated Sed. 2: 1700 kg/m³
- Consolidated Layered Sed.: 2150 kg/m³
- Breccia: 2350 kg/m³
- Shattered Zone: 2450 kg/m³
- Anisian to Jurassic Deposits: 2500 kg/m³
- Permian Deposits: 2550 kg/m³
-
- Crust: 2670 kg/m³

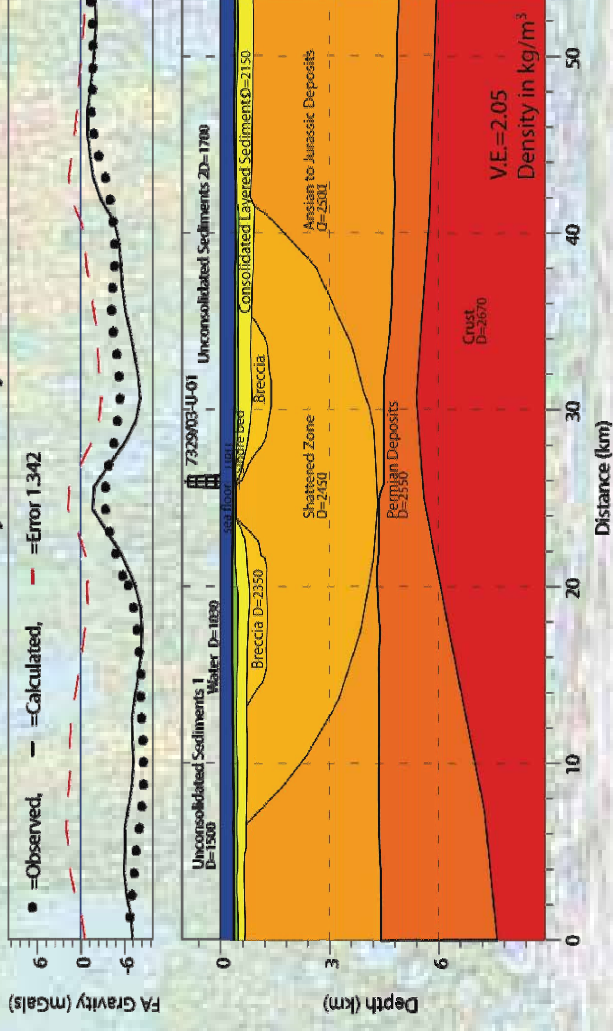
A

Trend-corrected Free Air Gravity Anomaly

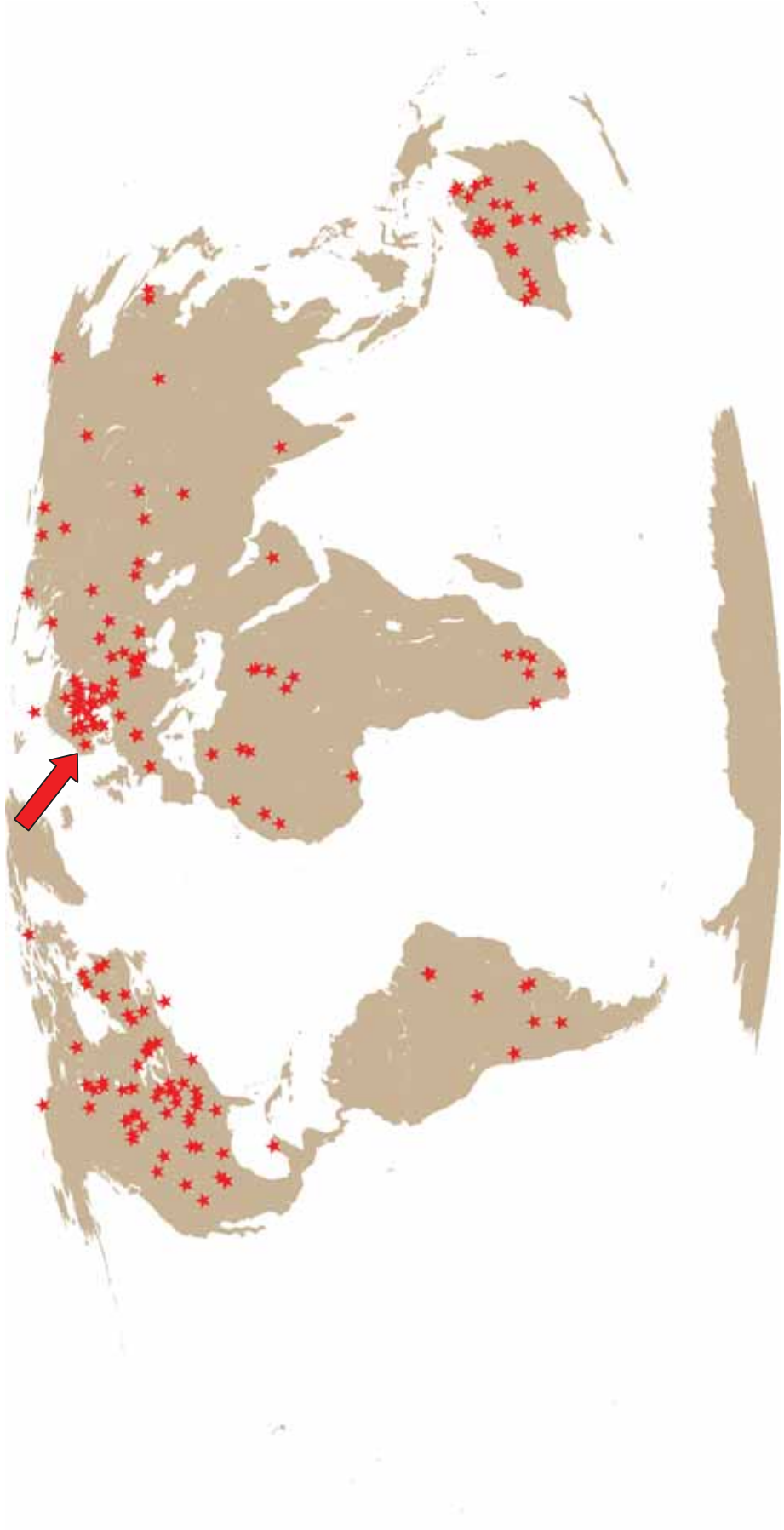


B

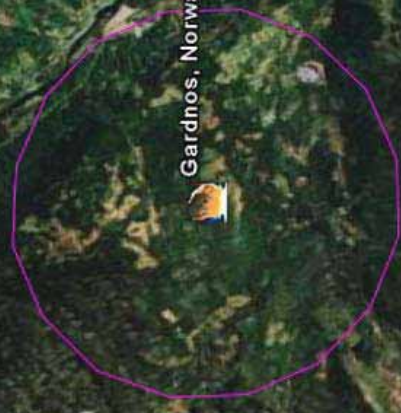
Uncorrected Free Air Gravity Anomaly



Global Crater Distribution



Gardnos



Gardnos, Norway

60°39' N 9°0'E
5 km 500 ± 10. Ma

Gol

Nesbyen

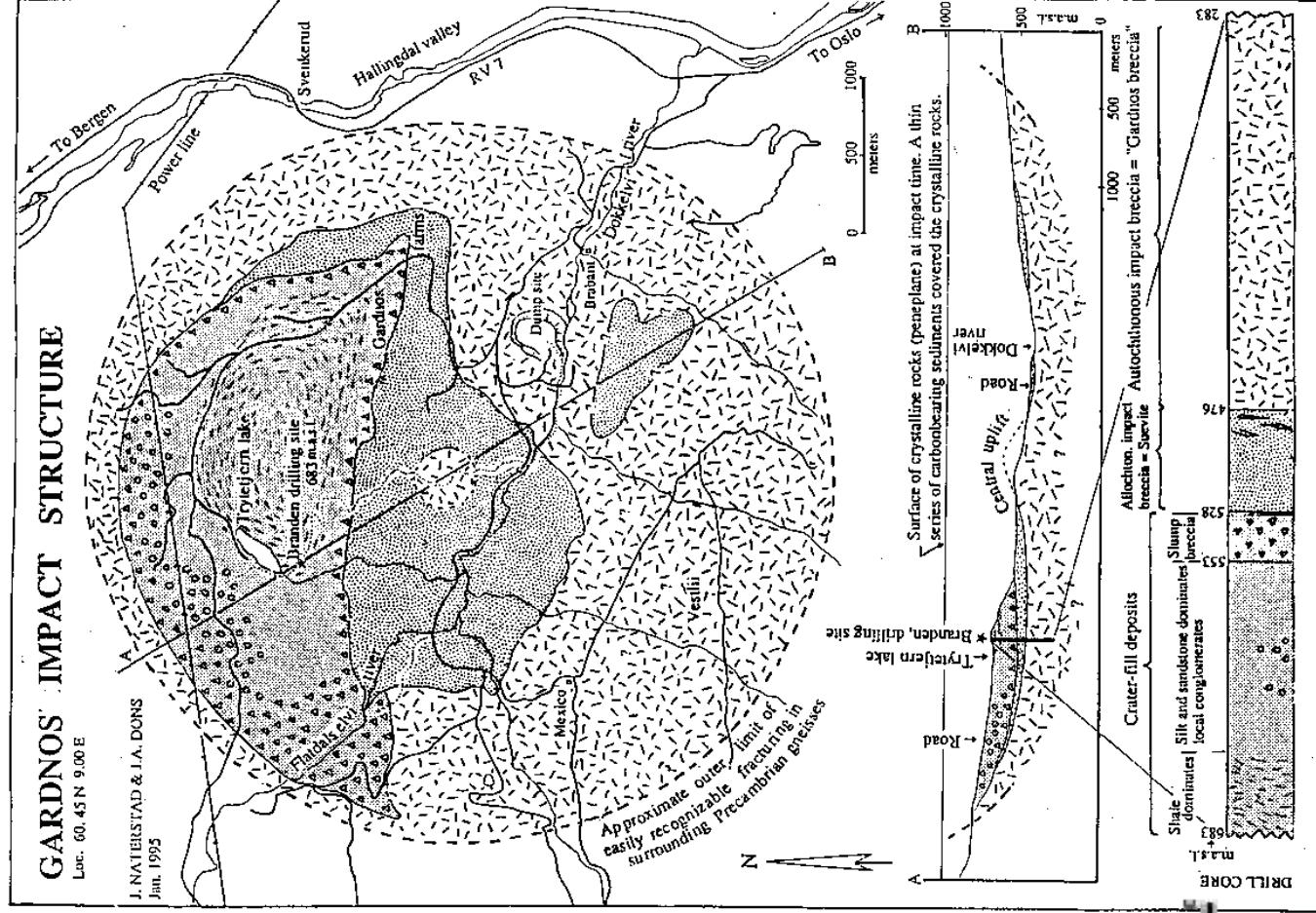
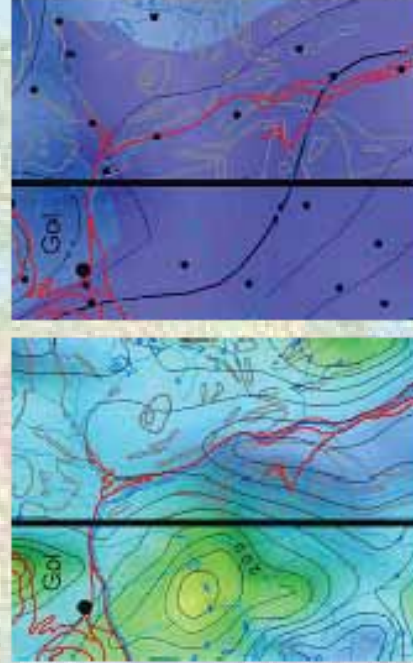
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Pointer 60°42'27.82"N 8°57'32.97"E elev 367 m

Streaming 100%

Gardnos

Gravity and Magnetic Anomaly Maps



Summary

- main attribute: *circularity*
- useful to describe the structural features
- though no proof for impact origin, good for *crater detection*
- *Gravity*: generally negative anomaly
- *Magnetics*: complex and dependent on the presence of magnetic mineral assemblages
- *Seismics*: structural information
- *Other Methods*: general mapping
- *Modelling* is needed for more information subsurface structures
- Calibration utilising properties measured in drill cores