

# Data and new constraints from Magnus-Rex



**W. Stratford** (1), H. Thybo (1), J. Faleide (2), O. Olesen (3) and A. Tryggvason (4).

(1) Department of Geography and Geology, University of Copenhagen, (2) Department of Geosciences, University of Oslo, (3) Geological survey of Norway, (4) Department of Earth Sciences, Uppsala University.

# Overview

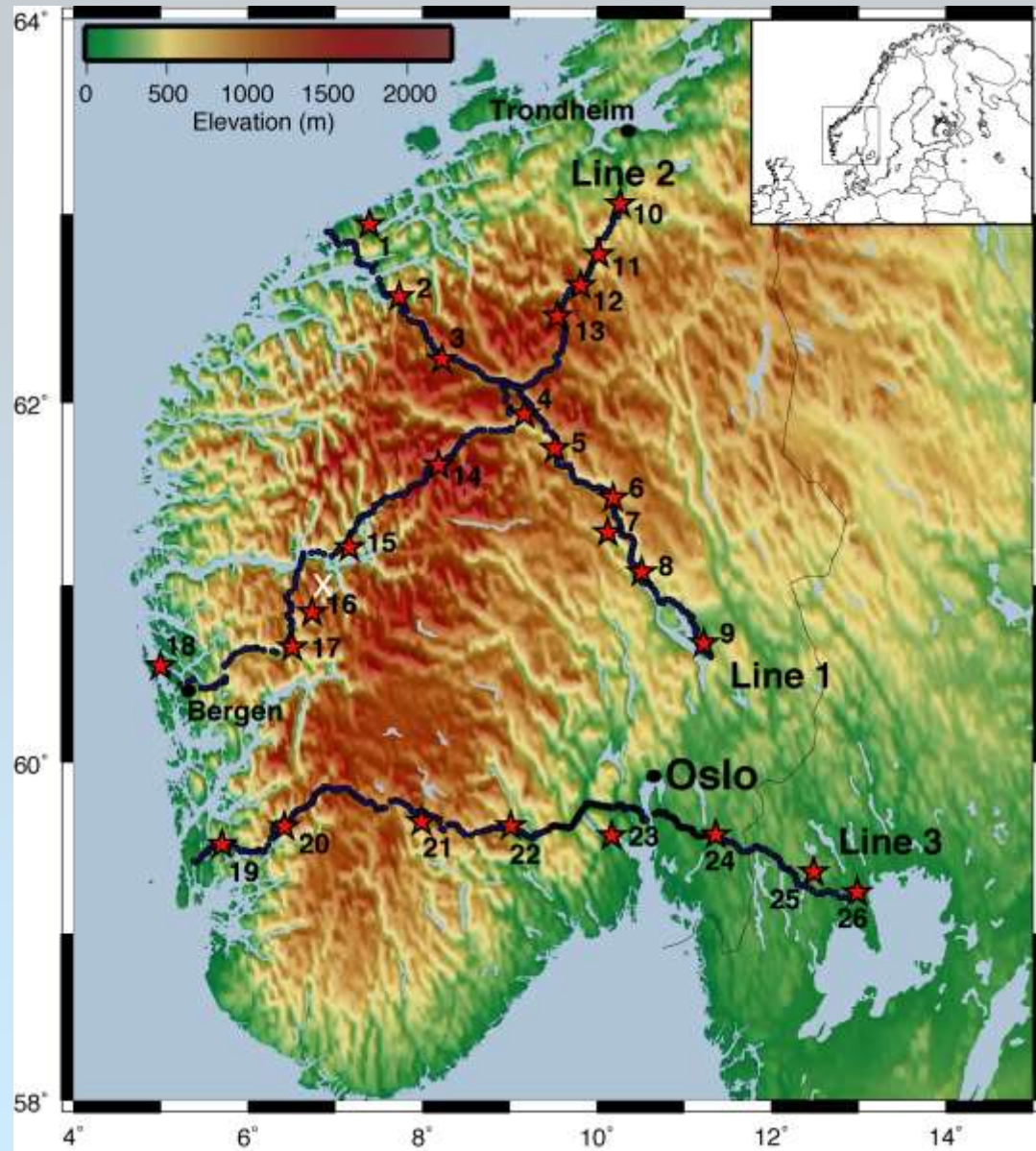
- Magnus-Rex - Mantle investigations of Norwegian uplift structure- Refraction experiment
- Project goals and preliminary findings from Magnus-Rex
- Field experiment
- Data and interpretations for the southern Scandes and Oslo Graben
- Poisson's Ratio Models for the upper crust
- Magnus-0



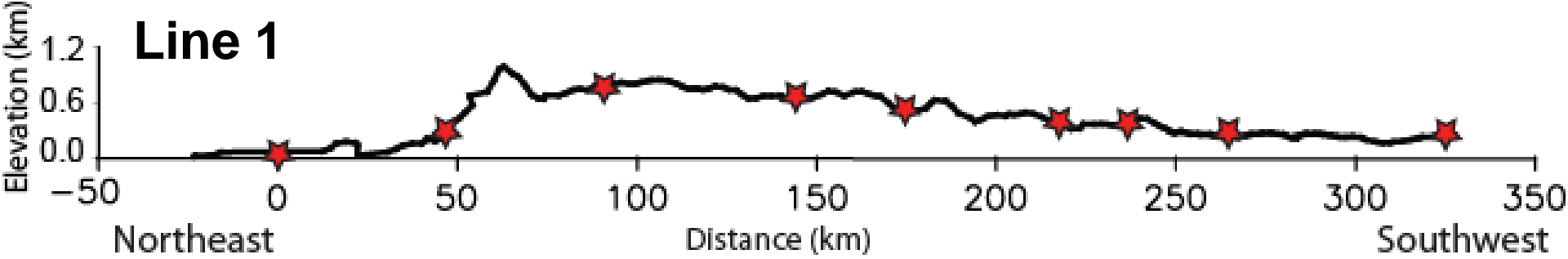


# Magnus-Rex

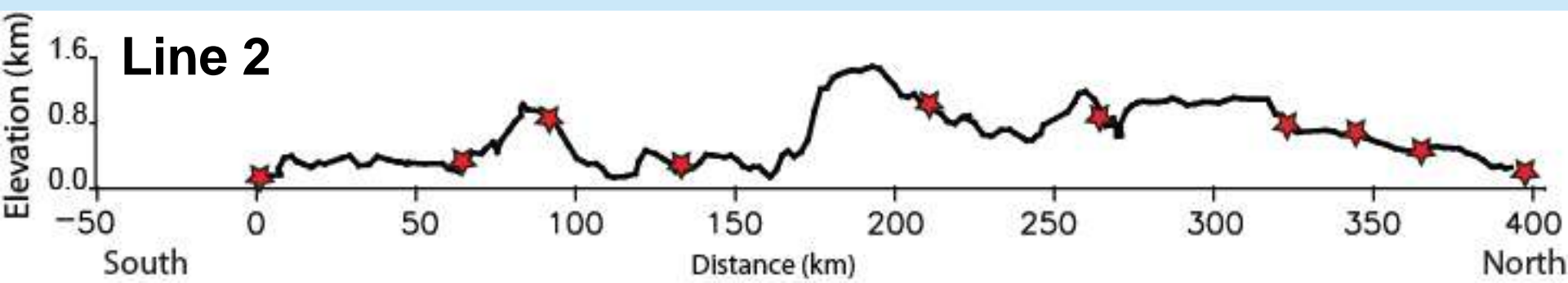
- Refractions profiling of the lithosphere. Best method of obtaining constraints on the velocity structure and thickness of the crust.
- 22 teams, ~750 texan instruments, 26 shots.
- Top down approach to forward modelling
- Constraint on crustal thickness ->
- $\pm 1$  km where Pn and PmP are available.
- $\pm 2$  km where only PmP is available.





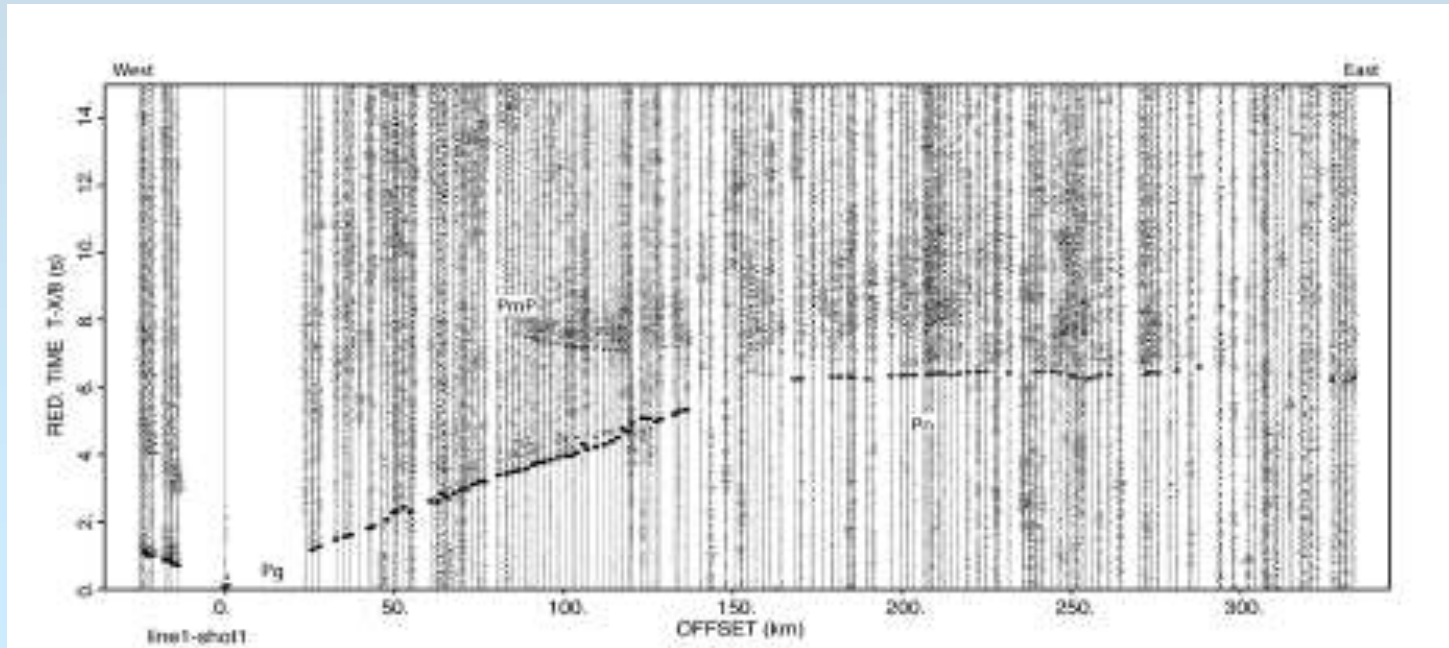


- ~360 km array
  - 9 shots (100, 200 and 400 kg charge sizes)
  - 174 instruments at 2 km spacing.
- 
- ~400 km array
  - 10 shots (100, 200 and 400 kg charge sizes)
  - 200 instruments at 2 km spacing.



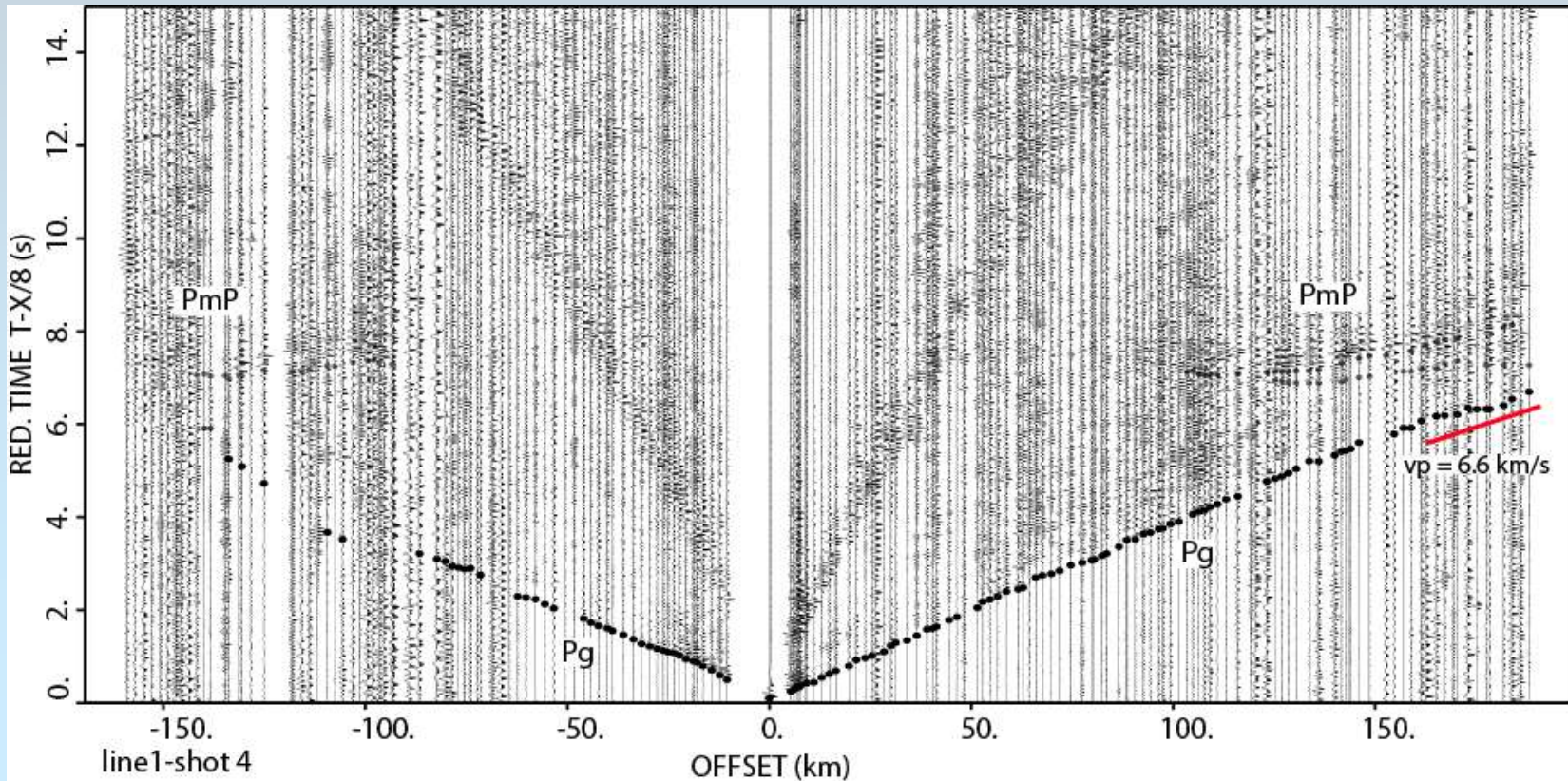
# Shot 1 - line 1

- 400 kg charge size
- $P_g V_p = 6.0 - 6.3 \text{ km/s}$
- $P_n V_p = 8.05 \pm 0.1 \text{ km/s}$  at  $x_{\text{cross}} > 160 \text{ km}$
- PmP reflection



# Shot 4 - line1

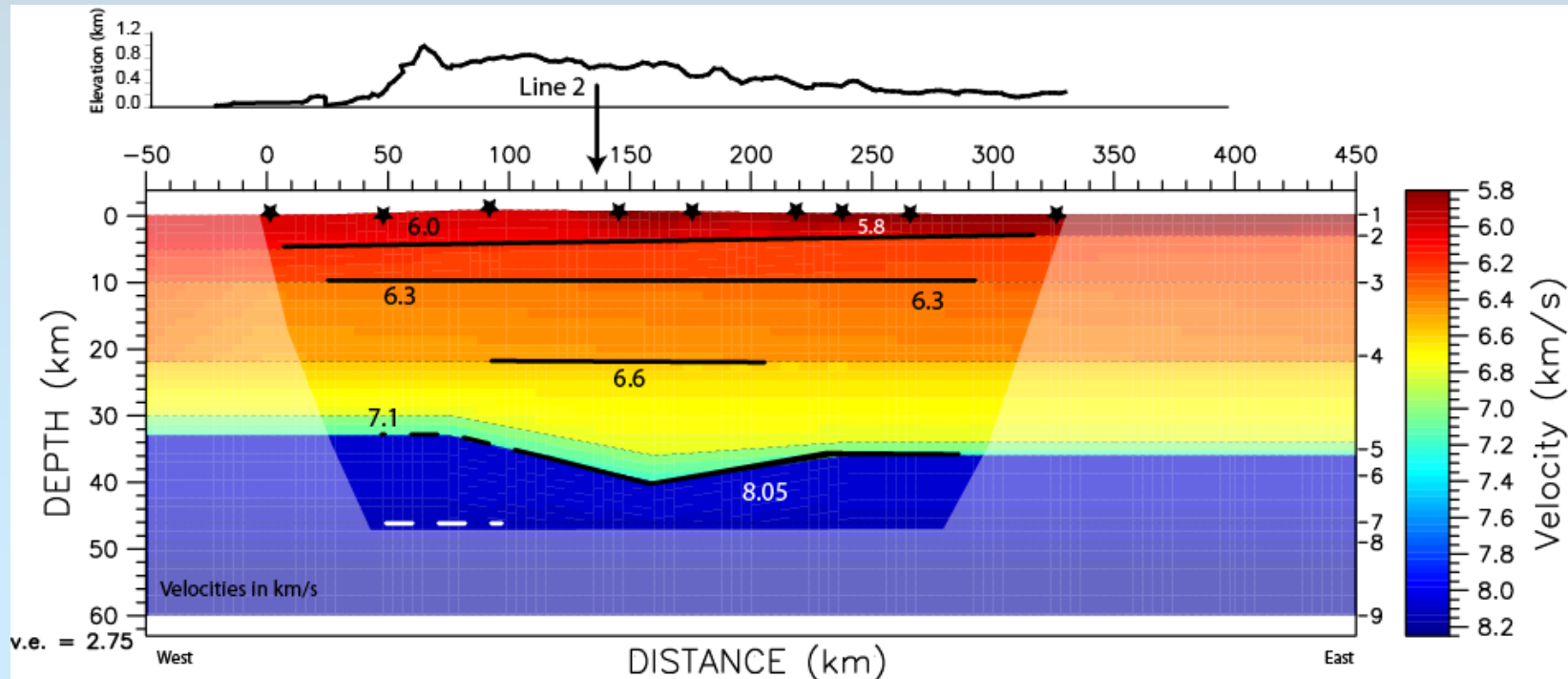
- 200 kg charge size
- $P_g$   $V_p = 5.8 - 6.6$  km/s
- $P_mP$  -> ~ flat moveout





# Velocity structure - Line 1

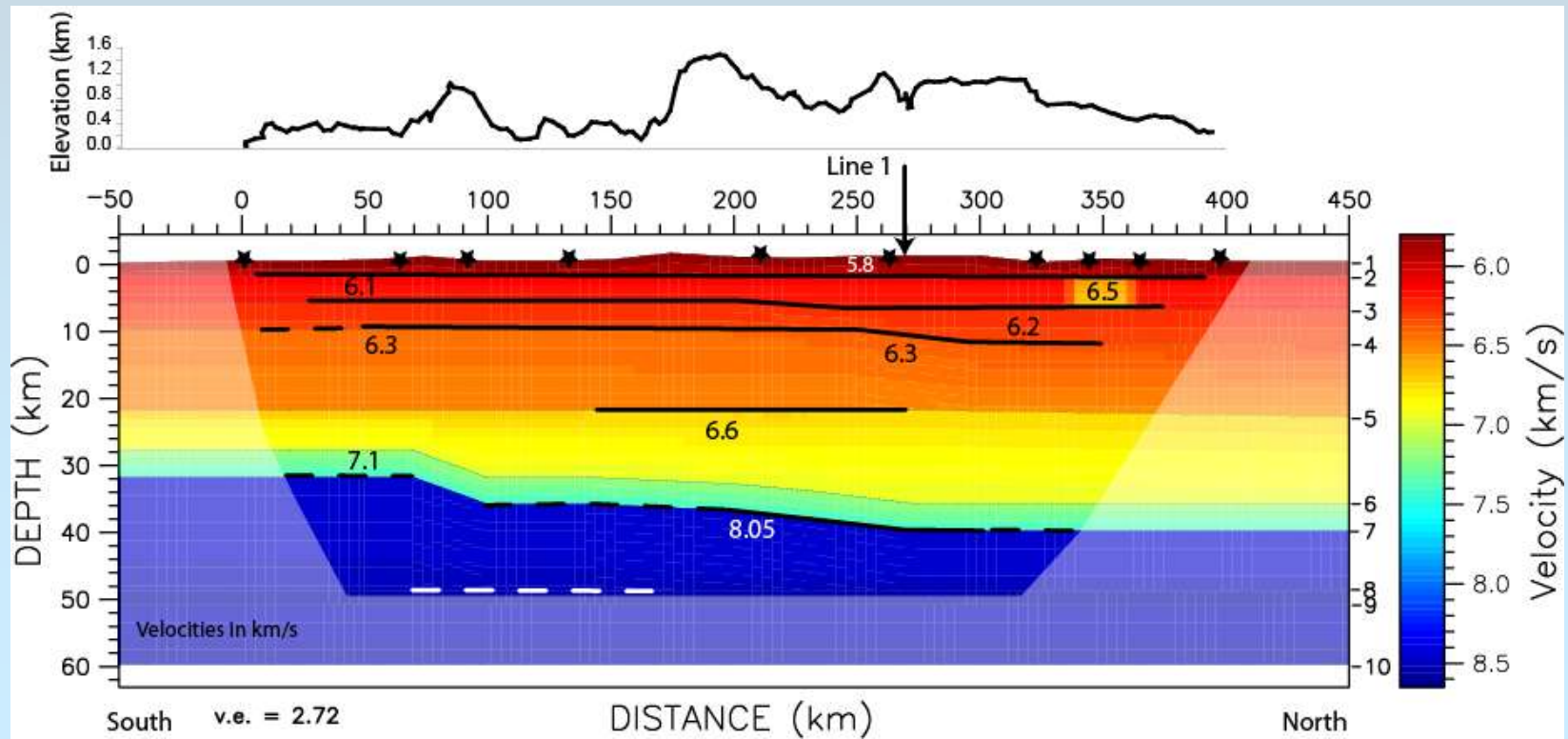
- Thick lines -> refraction velocity and depth constraint
- Thick dashed lines -> reflection structural constraint



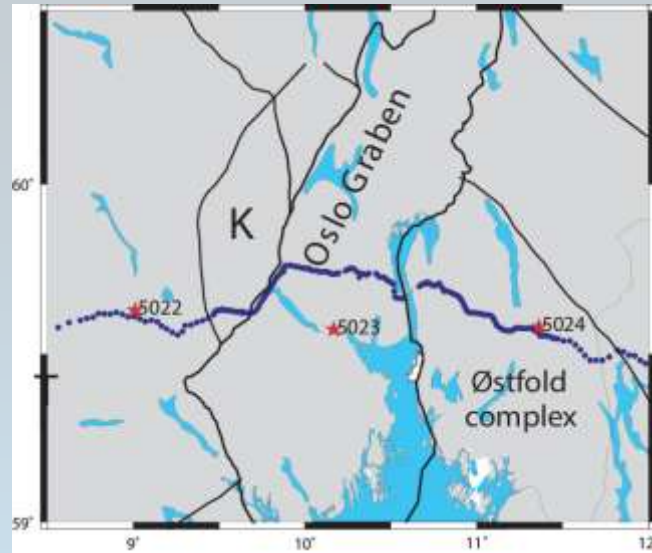


# Velocity structure - Line 2

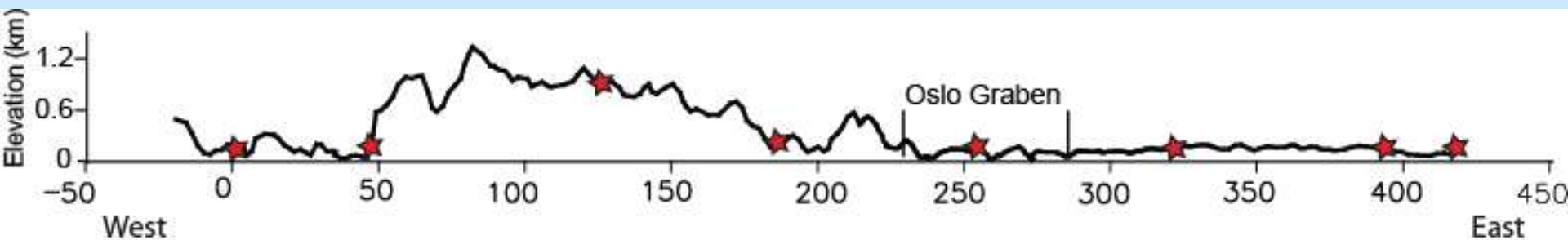
- Thick lines -> refraction velocity and depth constraint
- Thick dashed lines -> reflection structural constraint



# Line 3 - E-W across the southern Scandes and Oslo Graben



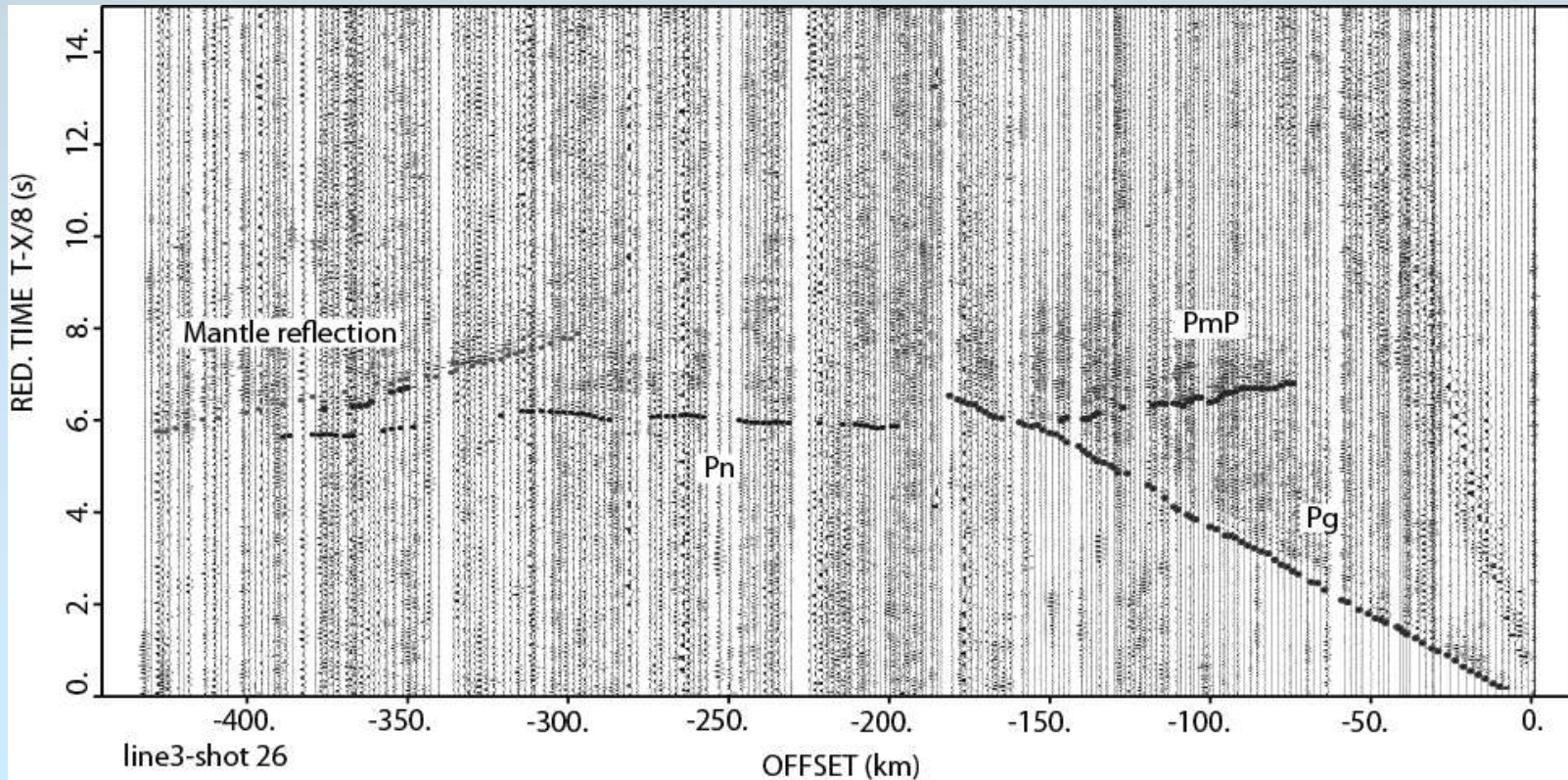
- ~450 km array
- 8 shots (100, 200 and 400 kg charge sizes)
- 330 instruments at 2 km spacing, including a 120 km section at 0.75 km spacing across the Oslo Graben





# Shot 26 - line 3

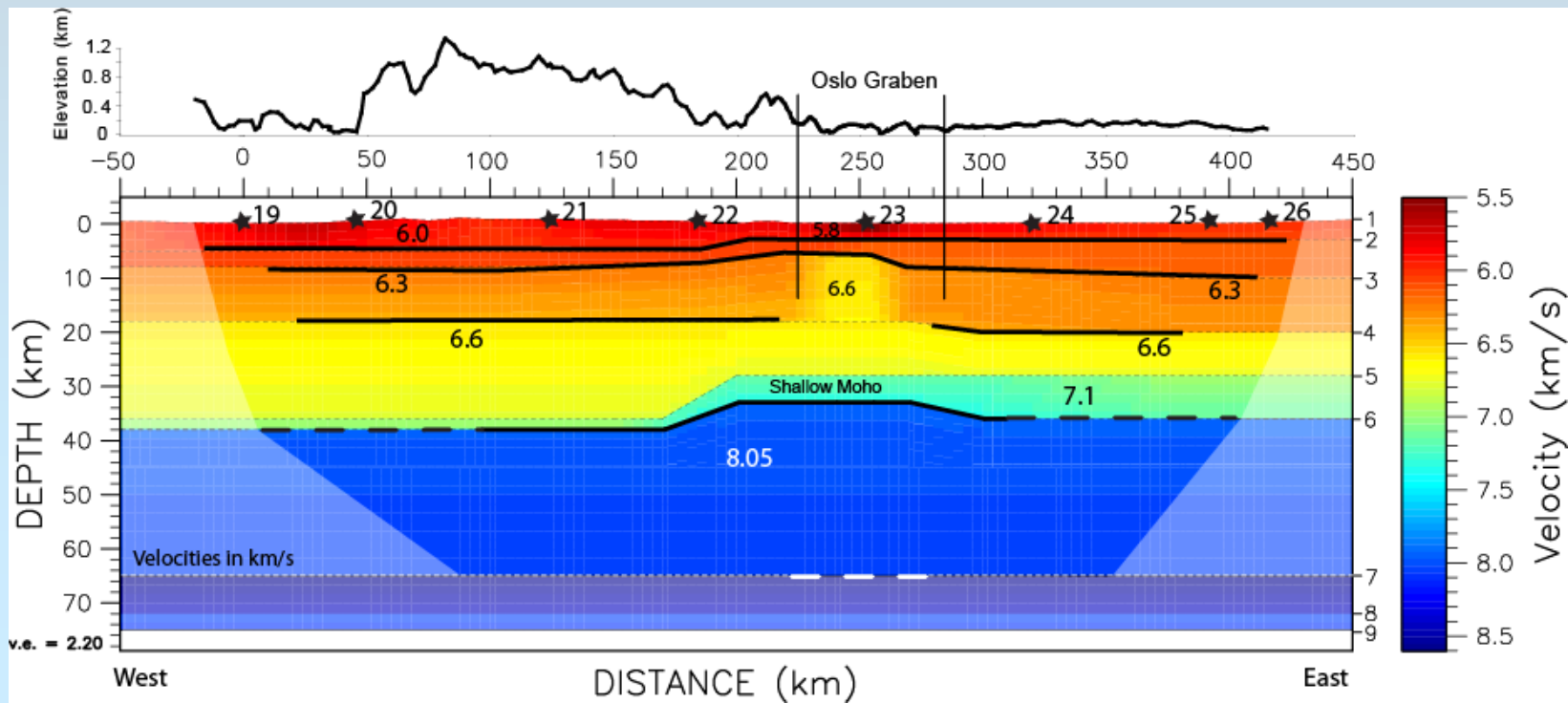
- 400 kg charge size
- Pg Vp = 5.8 - 6.3 km/s
- PmP
- Pn Vp =  $8.05 \pm 0.1$  km/s
- Mantle reflection behind Pn



# Velocity structure- Line 2 - deep Moho

Thick lines -> refraction velocity and depth constraint

Thick dashed lines -> reflection structural constraint

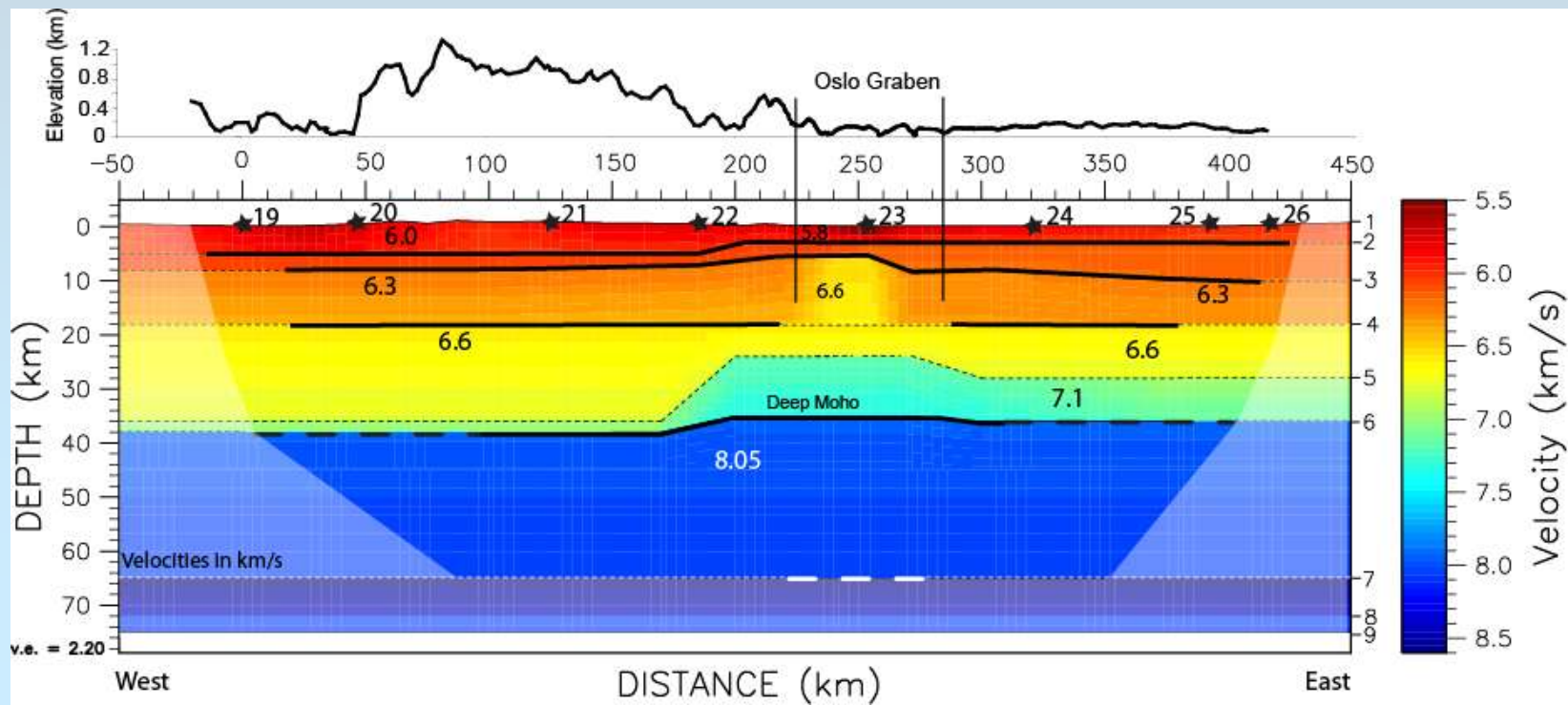


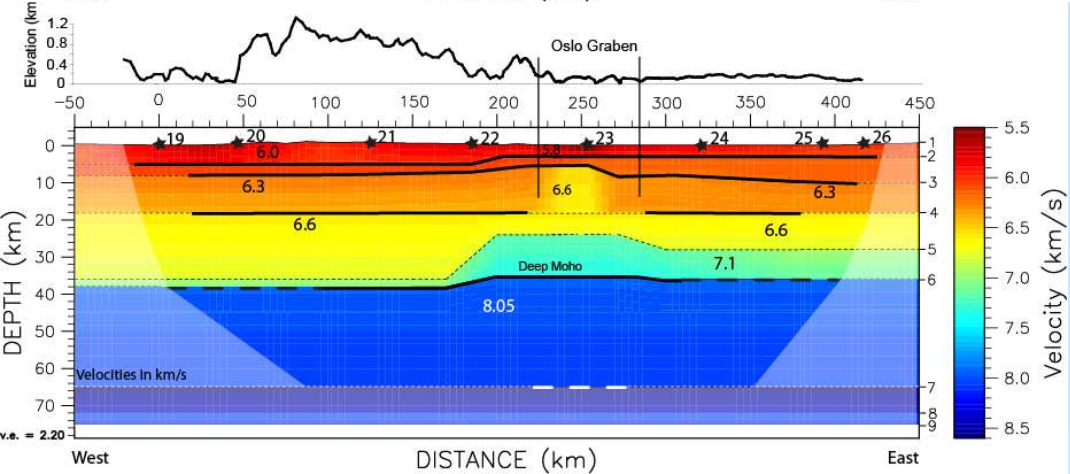
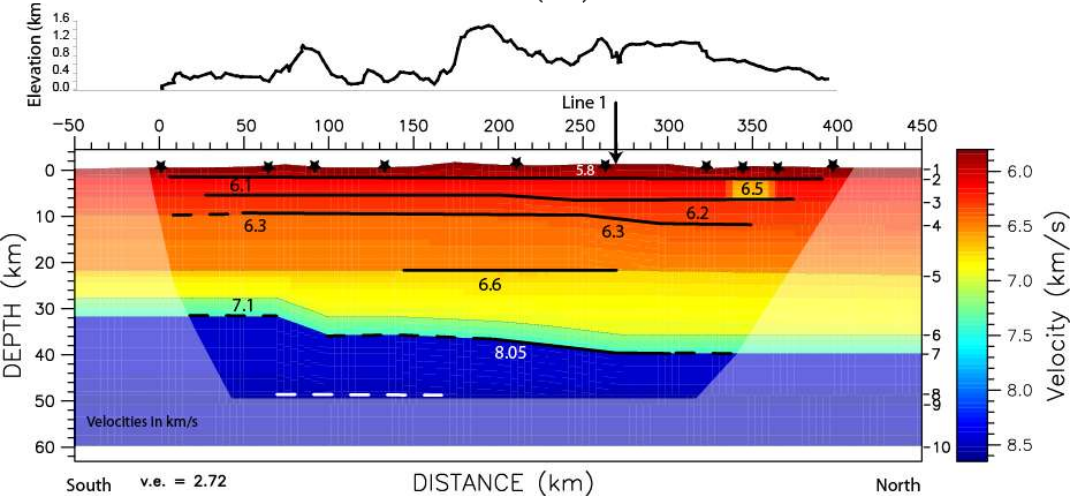
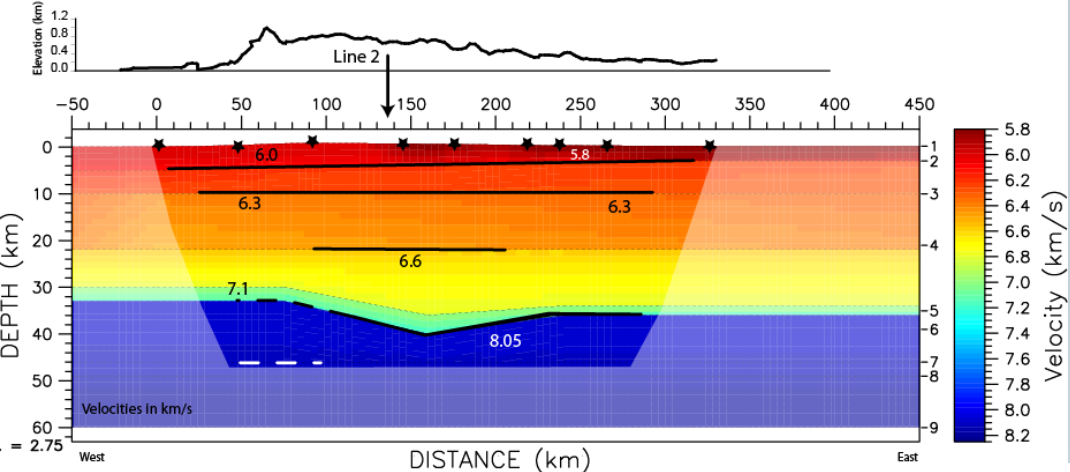


# Velocity structure- Line 2 - deep Moho

Thick lines -> refraction velocity and depth constraint

Thick dashed lines -> reflection structural constraint





Line 1



Line 2

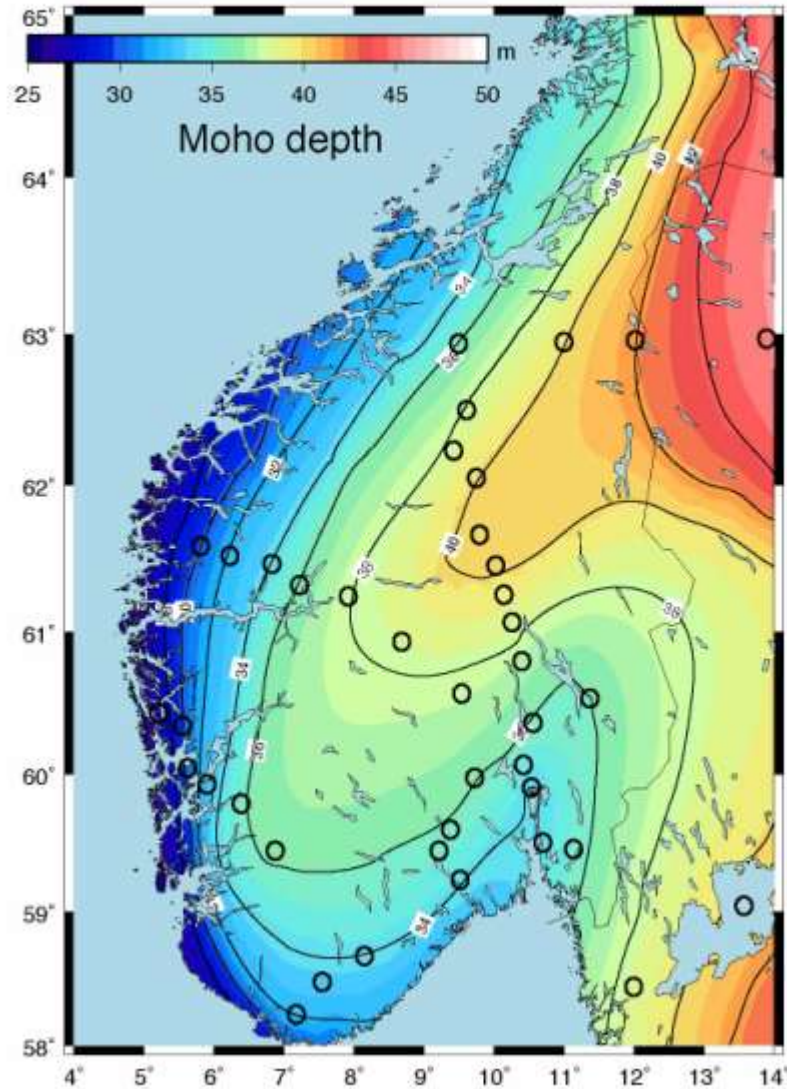
Moho depth  
beneath the  
Magnus-Rex  
seismic profiles

Line 3



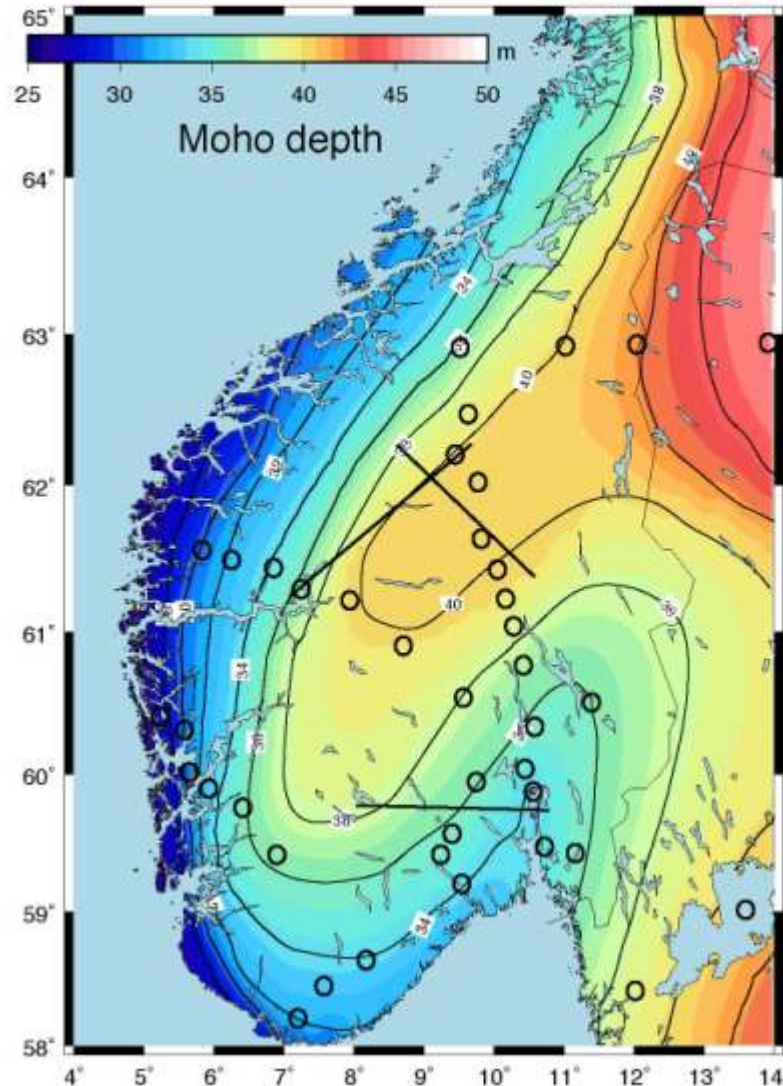
# Contour map of crustal thickness

Before



# Contour map of crustal thickness

After



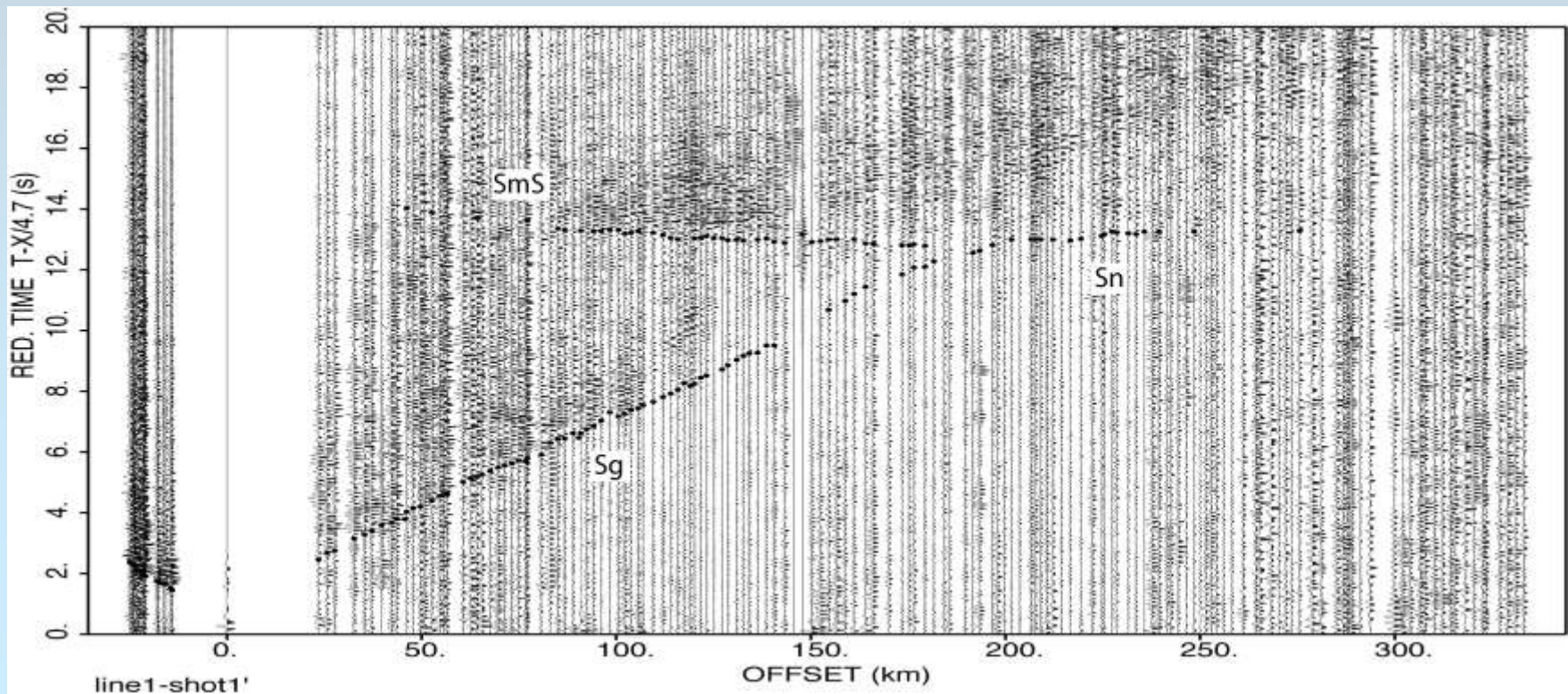
Original constraints from earlier refraction seismic studies still stand. Changes inferred are in gaps between previous surveys.



# We also recorded S-waves...

## Shot 1 - line 1

- 400 kg charge size
- Sg Vs = 3.35 – 3.65 km/s
- Sn Vs = 4.65 ± 0.1 km/s
- SmS reflection



# Line 1

Comparing Poisson's ratio ( $\sigma$ ) for caledonides to "basement"

Basement = mostly rock deformed in the Sveconorwegian Orogeny from 1250-900 Ma.

## Uncertainties in Poisson's Ratio

$$\delta\sigma/\sigma = f(\eta)[\delta V_p/V_p + \delta V_s/V_s]$$

$$f(\eta) = 2\eta^2/(\eta^2 - 1)(\eta^2 - 2)$$

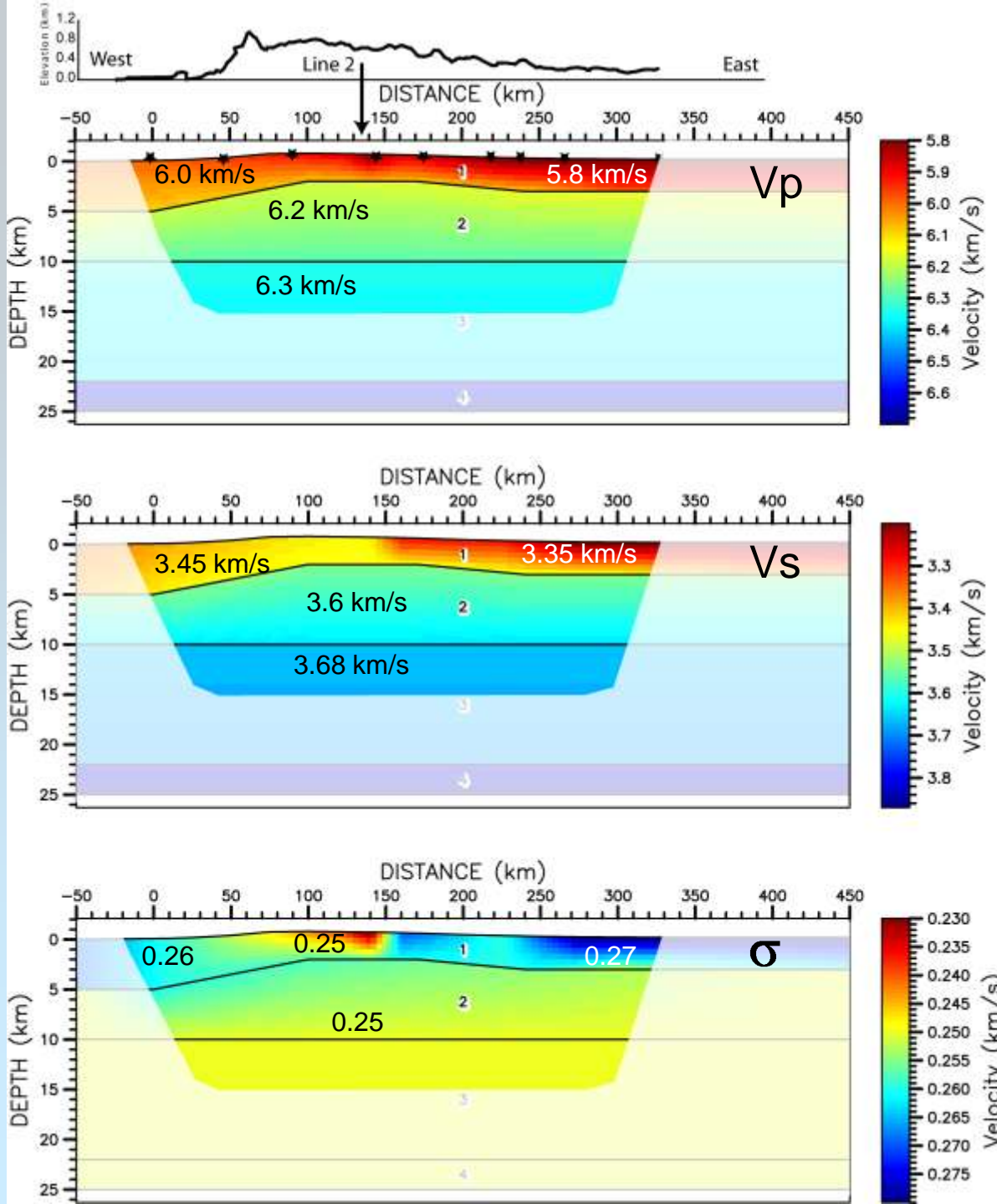
Where  $\eta = V_p/V_s$

*Christensen, (1996).*

## Uncertainties in $\sigma$

~ 3.5% or  $\pm 0.009$  for  $\sigma = 0.29$

~ 6.5% or  $\pm 0.014$  for  $\sigma = 0.20$

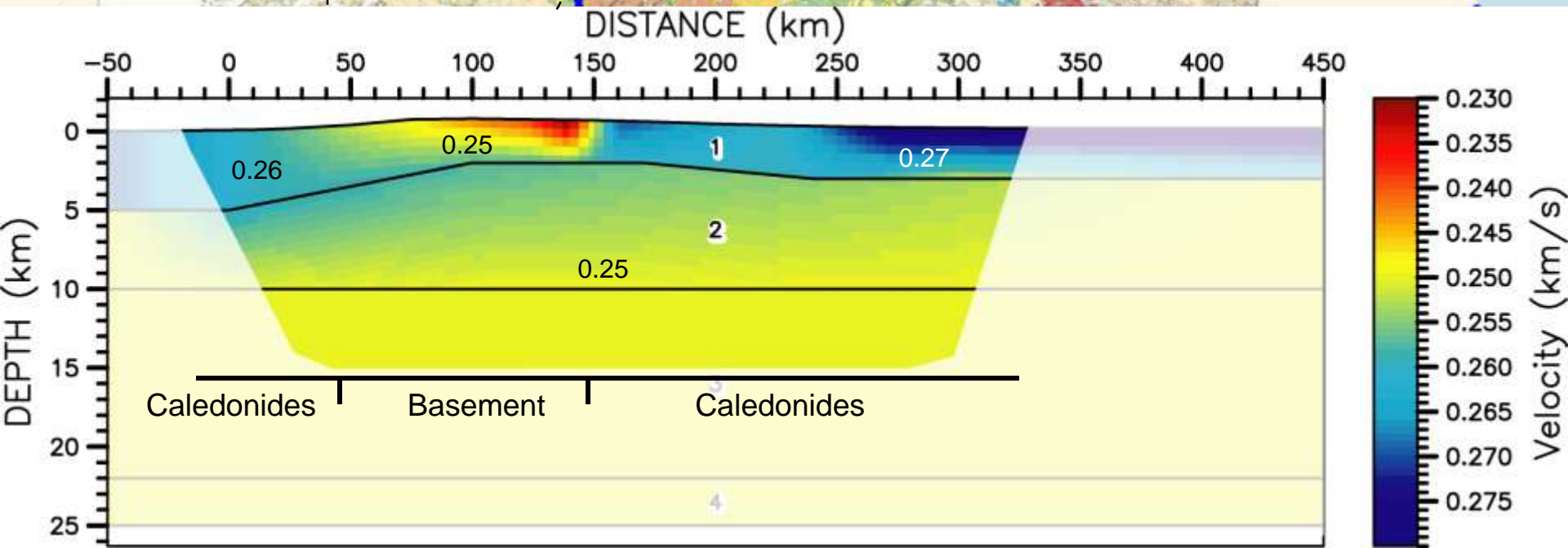
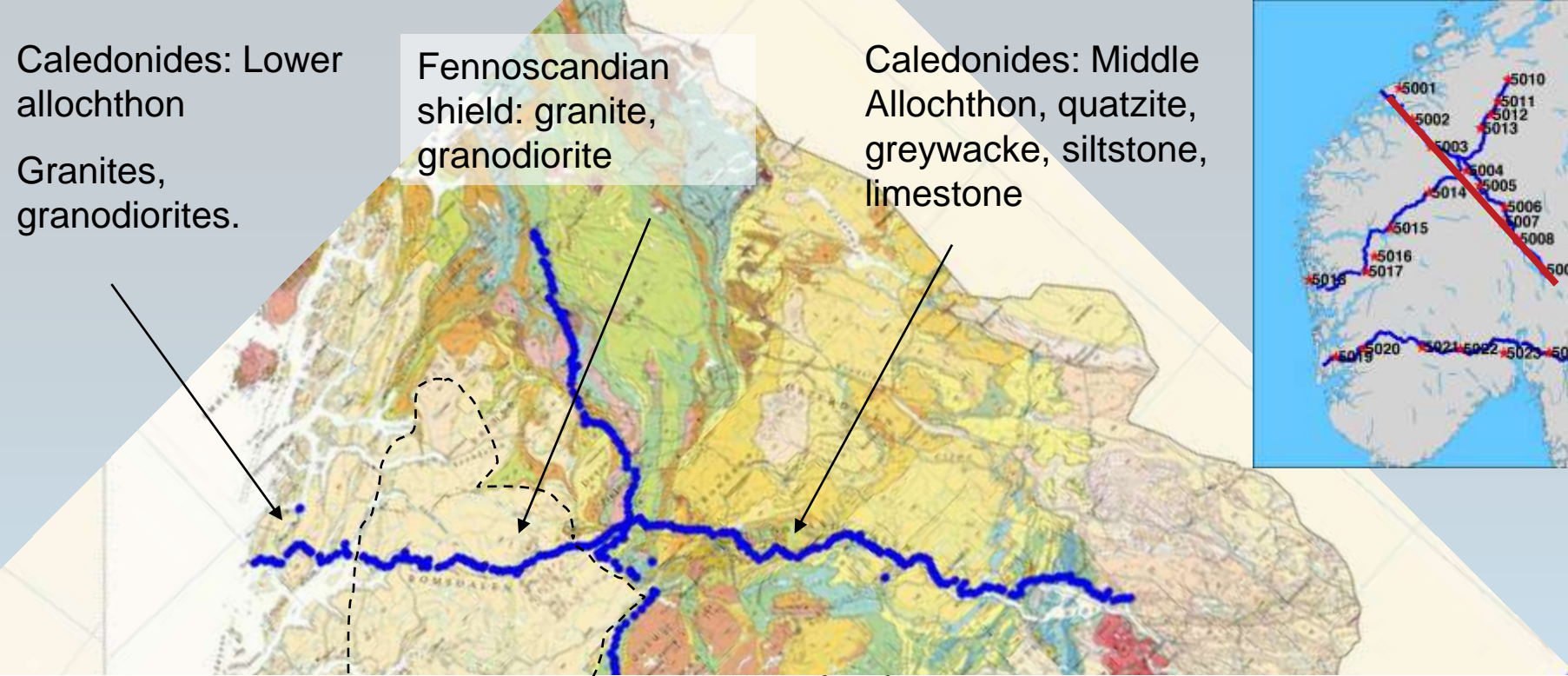




Caledonides: Lower allochthon  
Granites, granodiorites.

Fennoscandian shield: granite, granodiorite

Caledonides: Middle Allochthon, quartzite, greywacke, siltstone, limestone



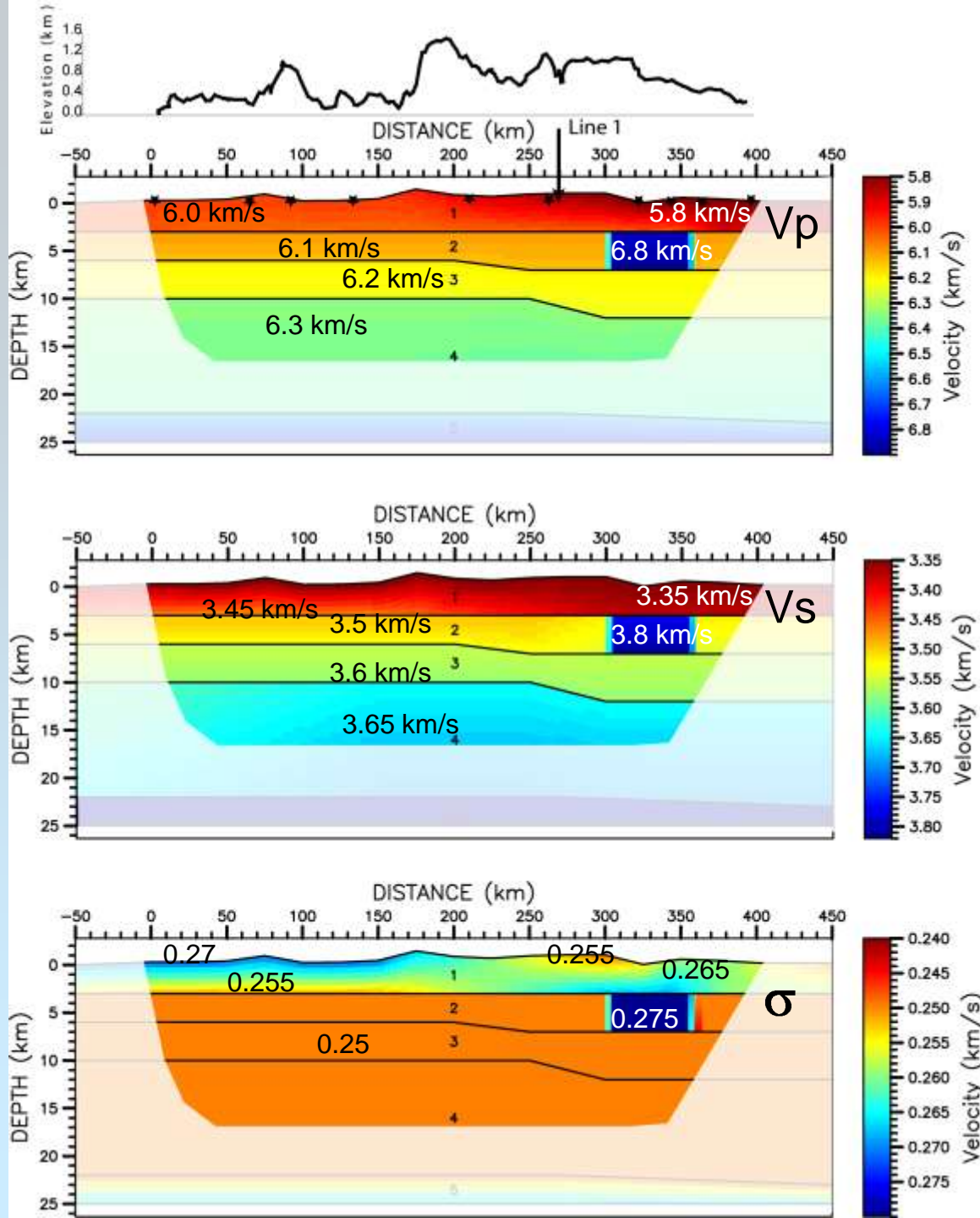
# Line 2

Profile line runs ~ north-south - along the western edge of the Faltungsgraben.

Crustal layers thicken to the north

Block of high velocity rock ~ 3 km depth beneath the upper Allochthon.

Refracted arrivals down to 15 km depth indicate the crust below the caledonides is more or less uniform  $\sigma$  at a ~10 km scale.

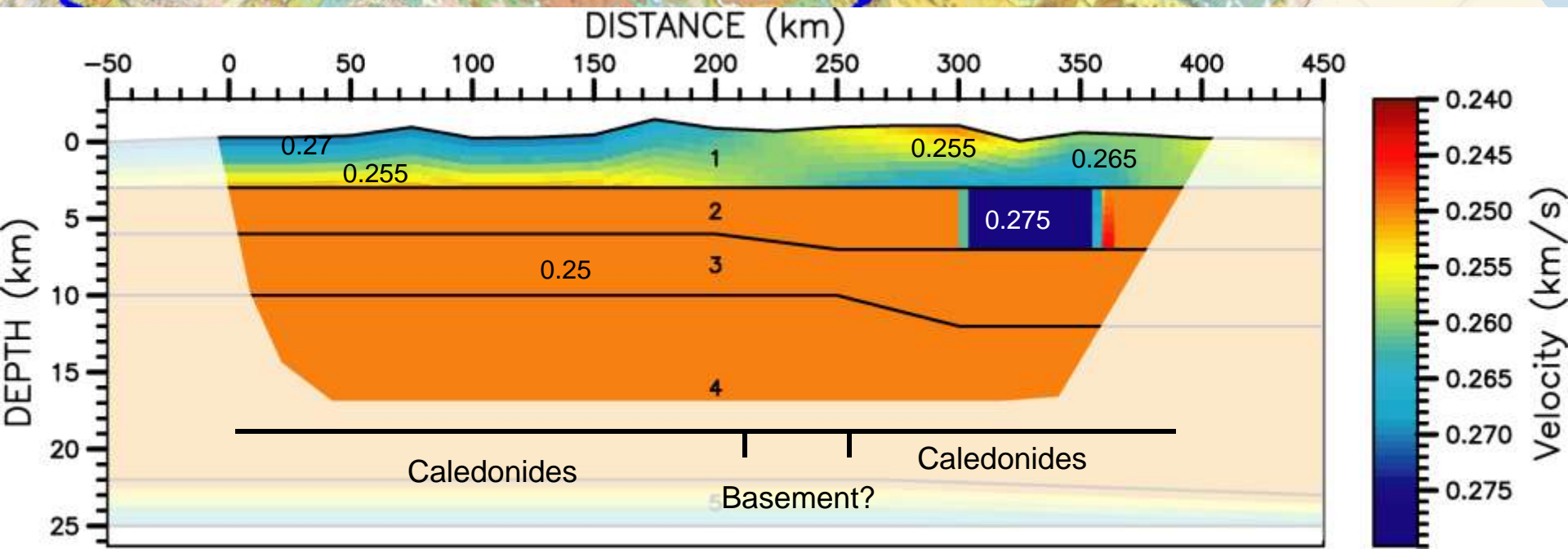
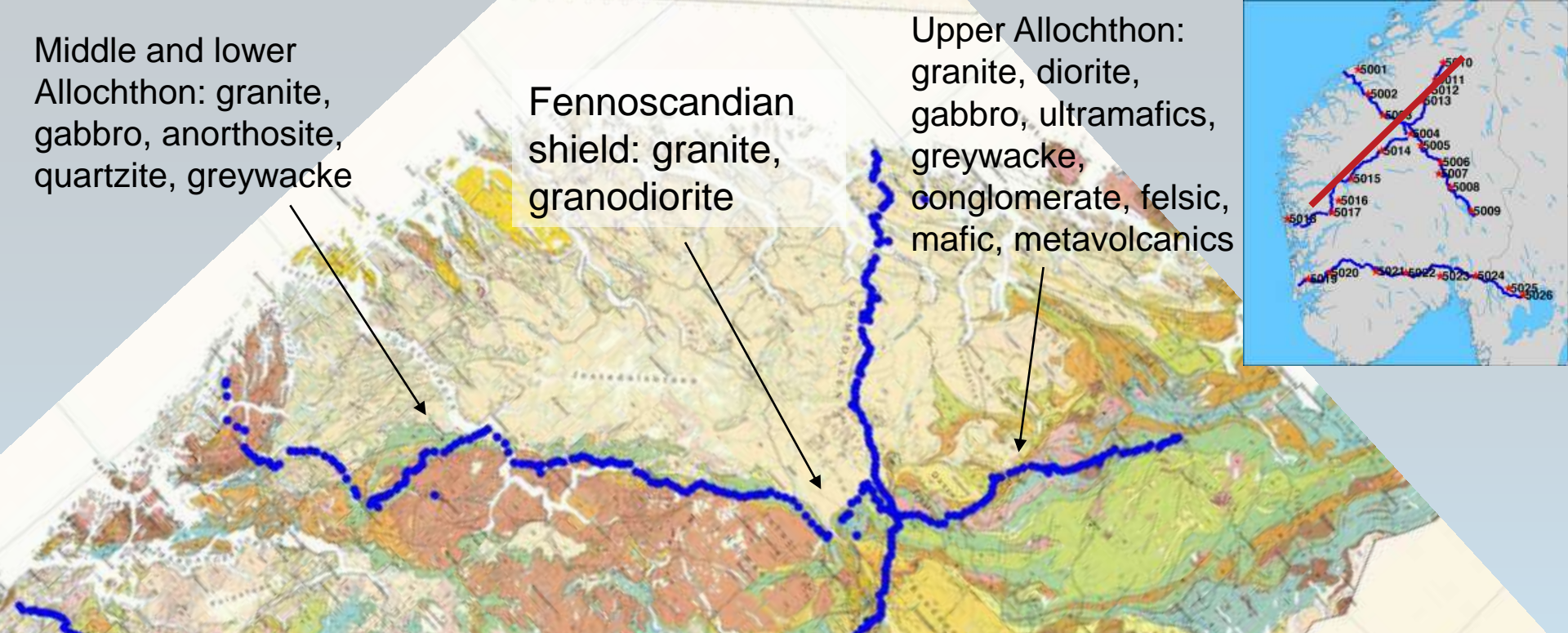




Middle and lower  
Allochthon: granite,  
gabbro, anorthosite,  
quartzite, greywacke

Fennoscandian  
shield: granite,  
granodiorite

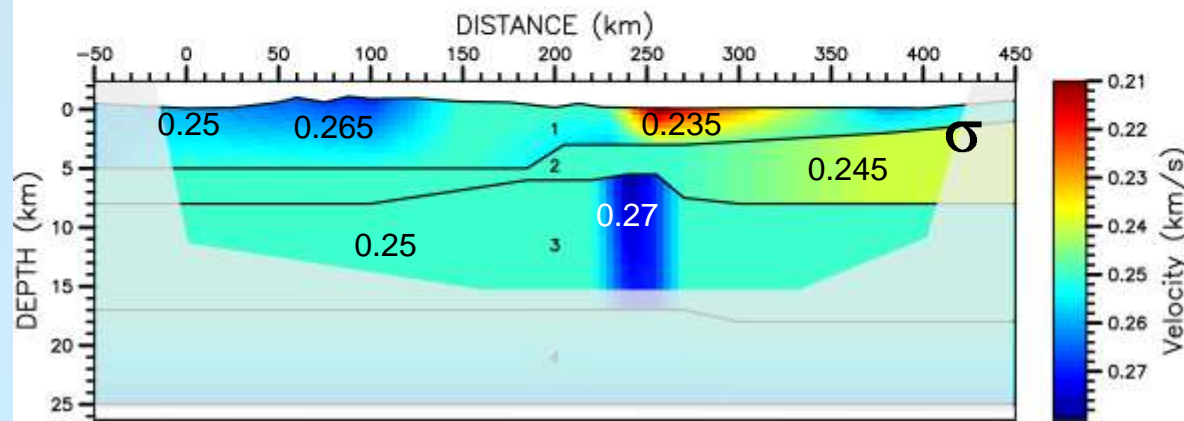
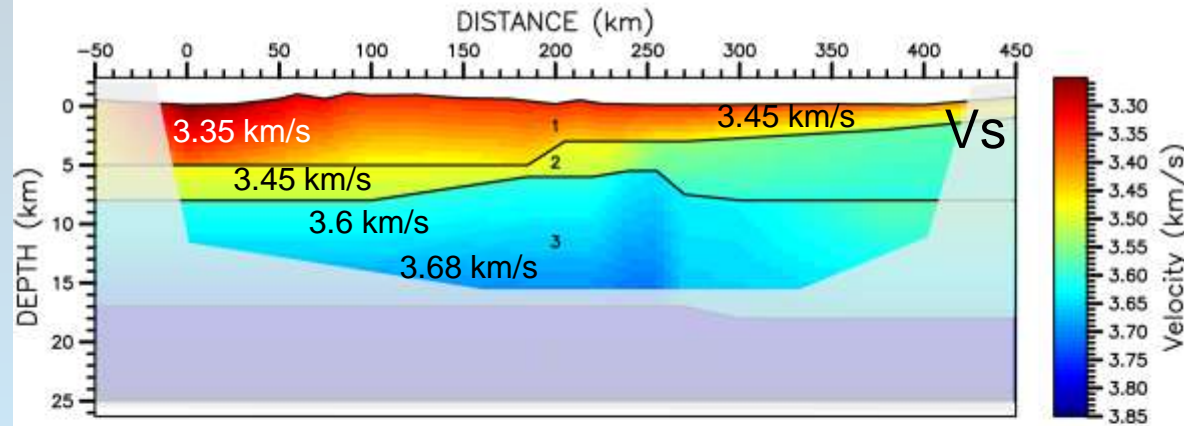
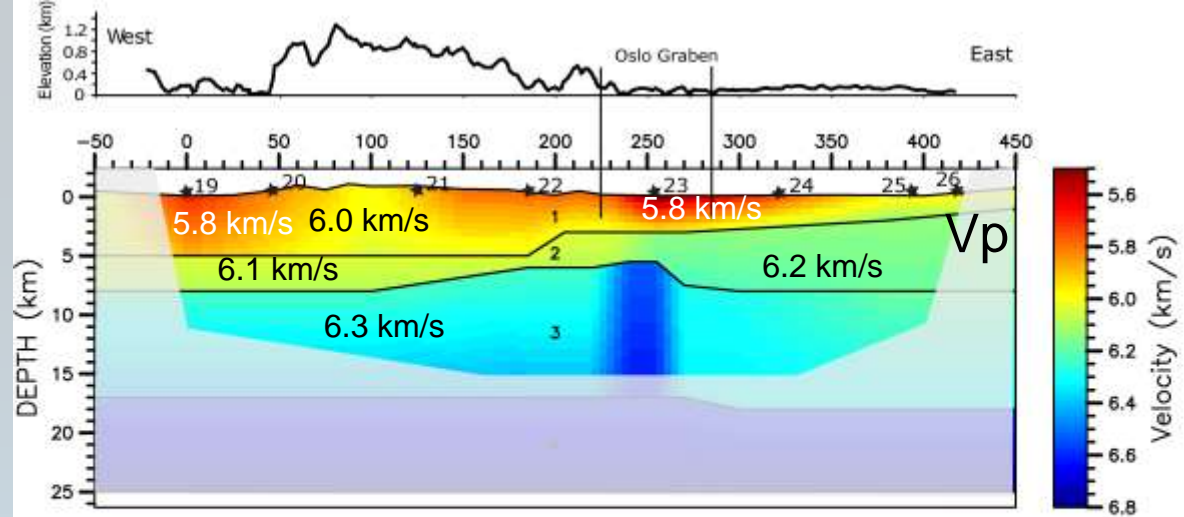
Upper Allochthon:  
granite, diorite,  
gabbro, ultramafics,  
greywacke,  
conglomerate, felsic,  
mafic, metavolcanics



Line three - across southern Norway and the Oslo Graben.

Near surface  $V_p$  is slower within the graben.

High  $V_p$  at depth.



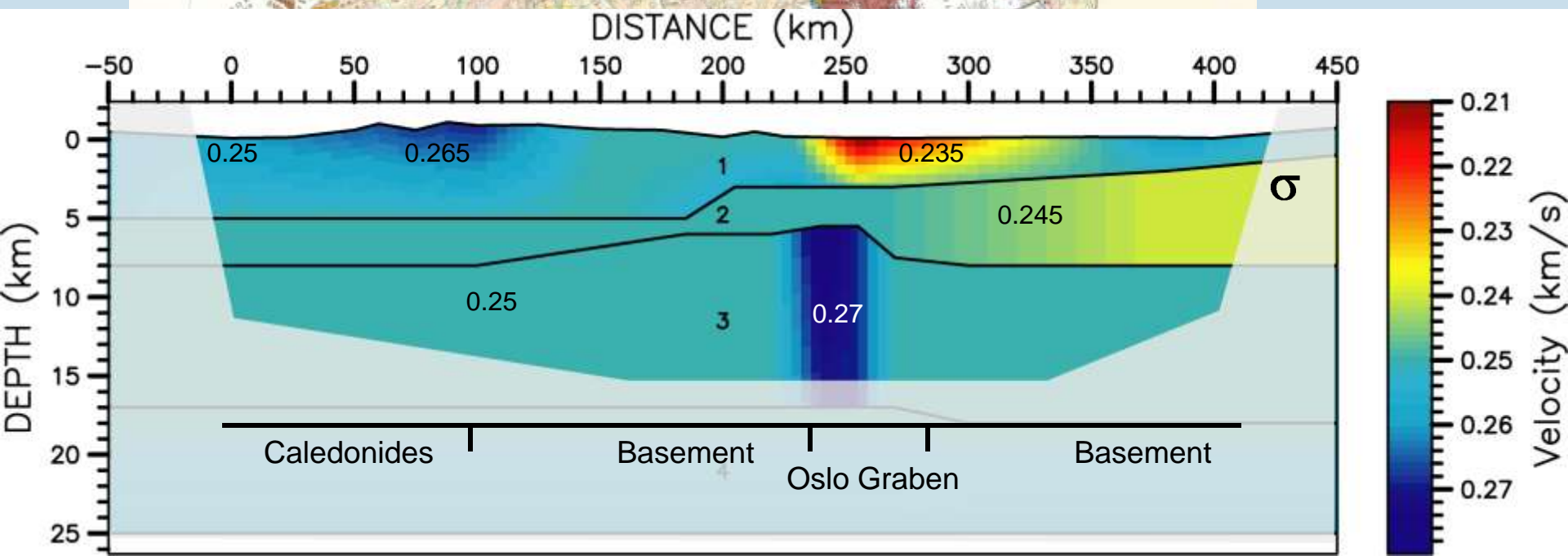


Middle and lower allochthon: granite, quartzite, greywacke, limestone

Fennoscandian shield: granite, granodiorite, mica schist

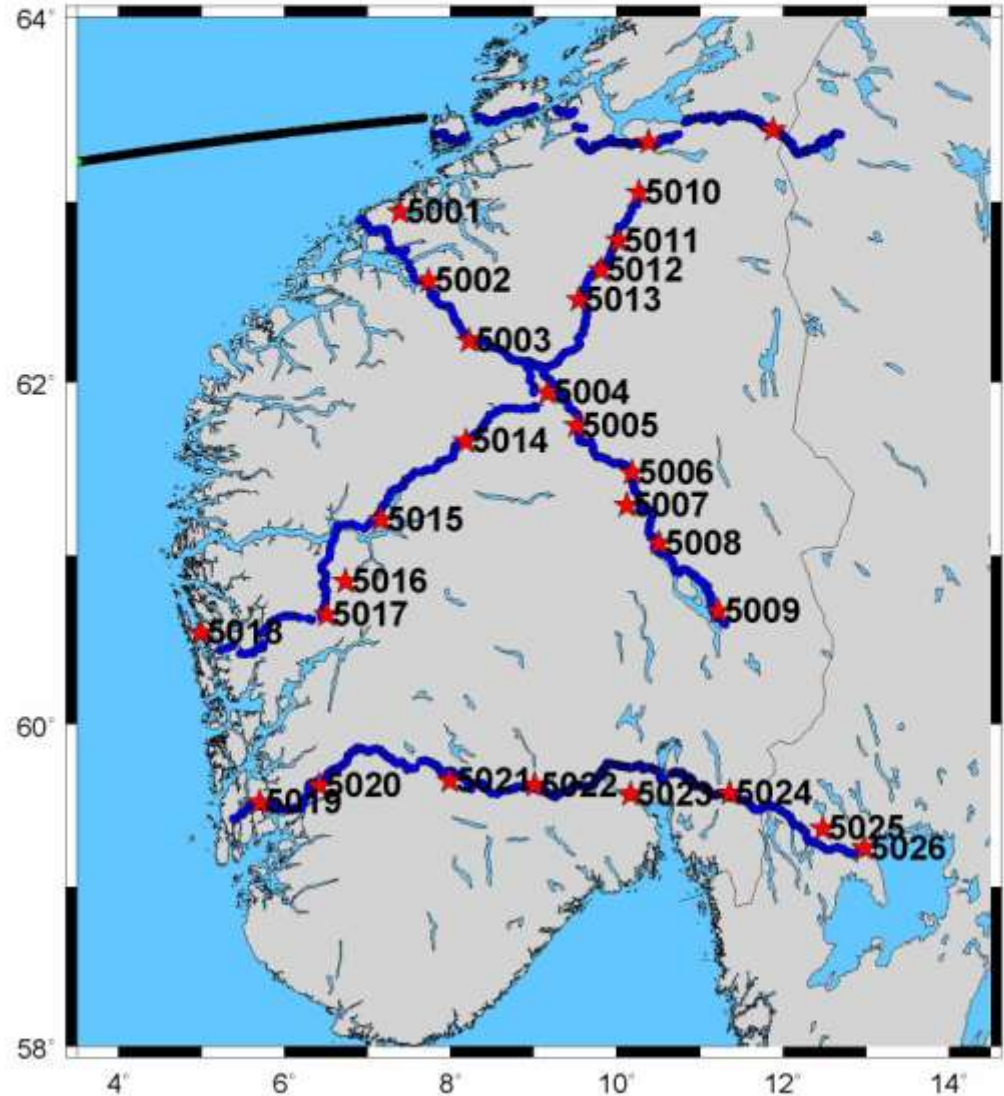
Fennoscandian shield: granite, granodiorite, metagreywacke, quartzite

Oslo Graben: granite, diorite, gabbro, rhyolite, basalt.



# Magnus-0

- Recorded in 2006. onshore/offshore refraction profile for the Trondheim area.
- 200m shot spacing onshore-offshore.
- 5 instrument clusters onshore for onshore-offshore shots.
- Onshore shots ~1km receiver spacing.





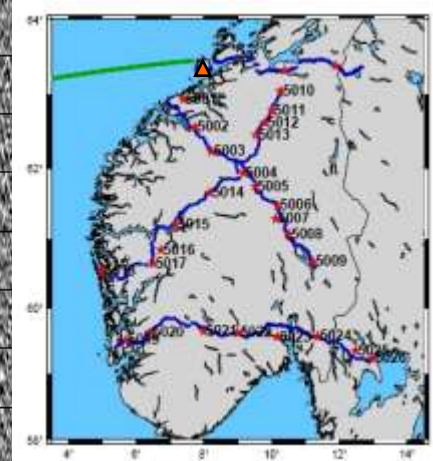
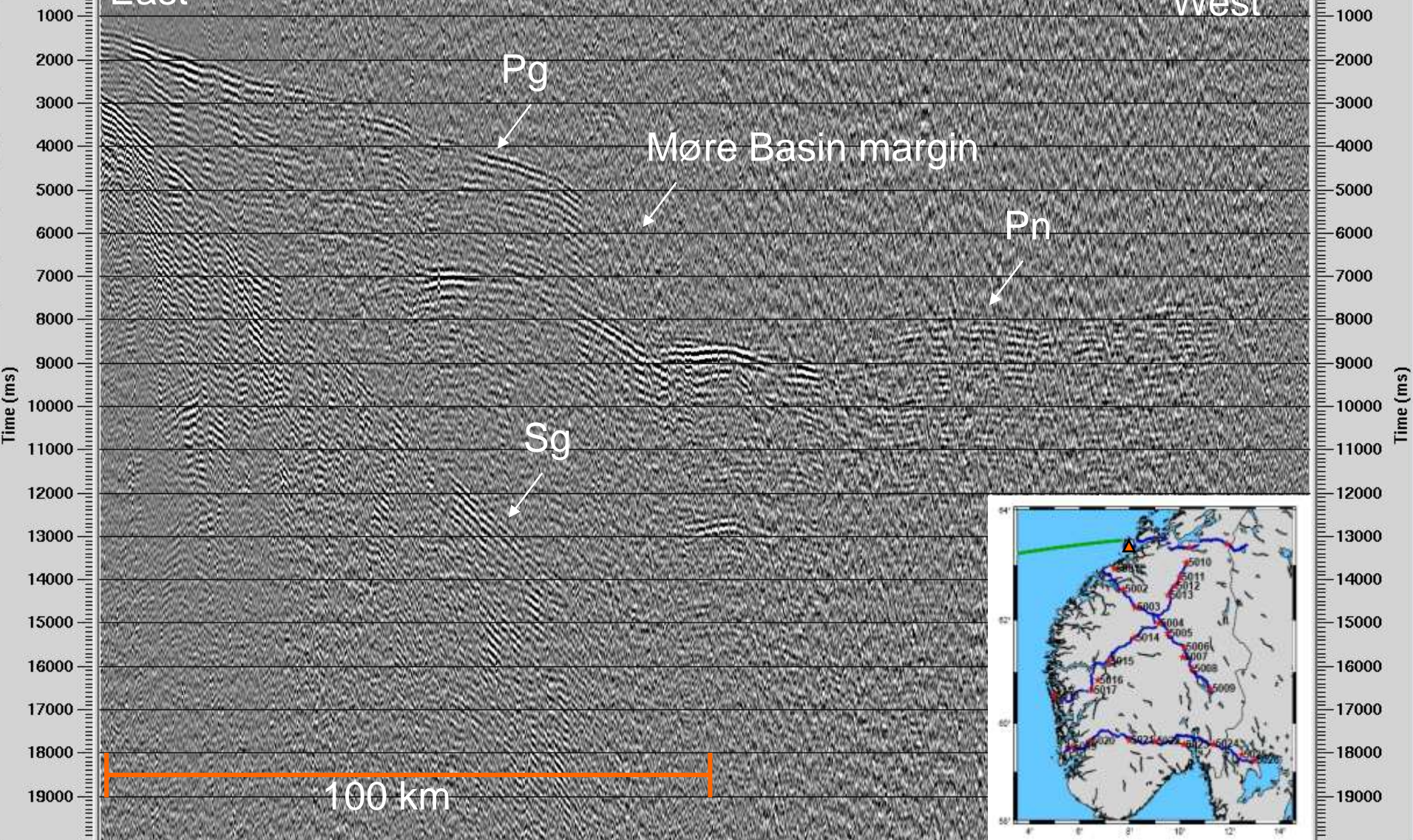
FFID

SOURCE

03 1053 1104 1155 1205 1258 1308 1360 1411 1461 1514 1564 1616 1666 1718 1769 1821 1872 1922

East

West





# Summary

- Crustal thickness beneath the southern Scandes mountains reach ~40 km.
- Crust thickens to the north
- Thins from the central mountains towards the coast and Oslo Graben, and to the south
- Strong S-wave energy is generated by shots in the caledonides, and Poisson's ratios highlight structural/compositional changes within the caldedonides.